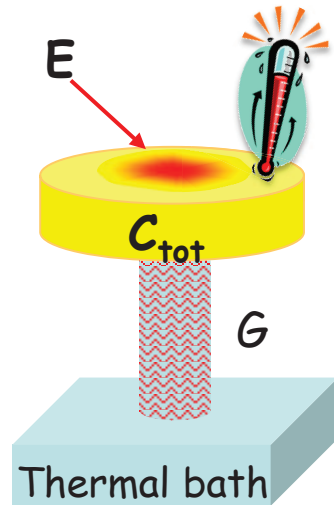


Magnetic Micro-Calorimeters

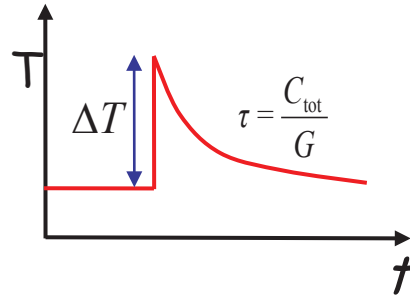
Loredana Gastaldo, S.Allgeier, A.Ferring, L.Gamer, J.Geist, S.Hähnle, C.Hassel, S. Hendricks, D.Hengstler, M.Keller, M.Krantz, W. Köntges, G. Möhl, J. Poller, C.Schötz, D.Schulz, M.Wegner, S.Kempf, A. Fleischmann and C.Enss

Heidelberg University

Low temperature micro-calorimeters



$$\Delta T \cong \frac{E}{C_{\text{tot}}}$$

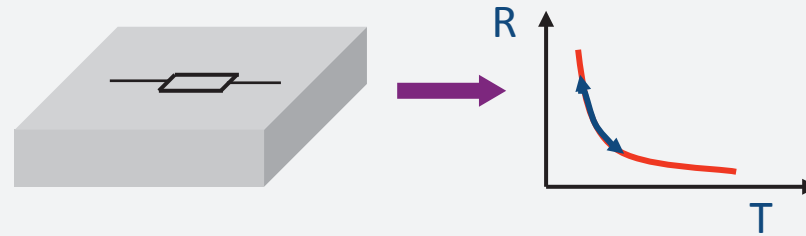


$$\left. \begin{array}{l} E = 10 \text{ keV} \\ C_{\text{tot}} = 1 \text{ pJ/K} \end{array} \right\} \rightarrow \sim 1 \text{ mK}$$

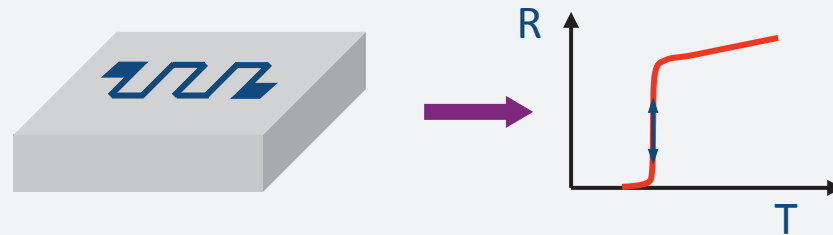
- Very small volume
- Working temperature below 100 mK
small specific heat
small thermal noise
- Very sensitive temperature sensor

Temperature sensors

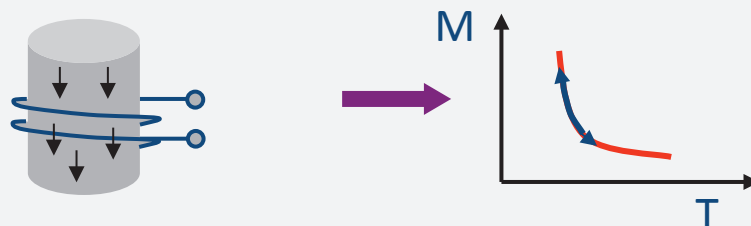
Resistance of highly doped semiconductors



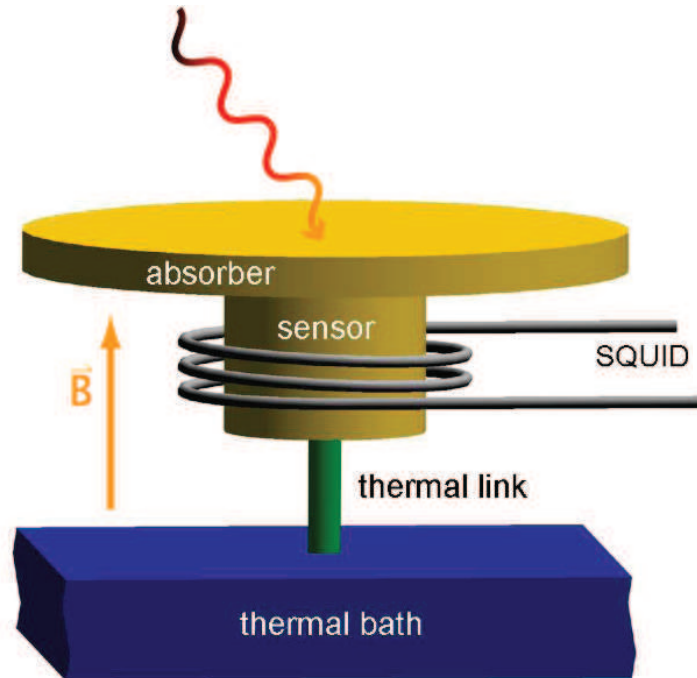
Resistance at superconducting transition, TES



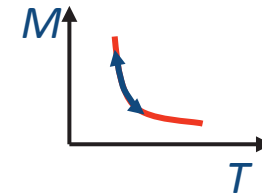
Magnetization of paramagnetic material



Metallic Magnetic Calorimeters - MMC



Paramagnetic sensor: **Au:Er_{500ppm}**



Signal size:

$$\delta M = \frac{\partial M}{\partial T} \delta T = \frac{\partial M}{\partial T} \frac{E_\gamma}{C_{\text{tot}}}$$

main differences to calorimeters with resistive thermometers

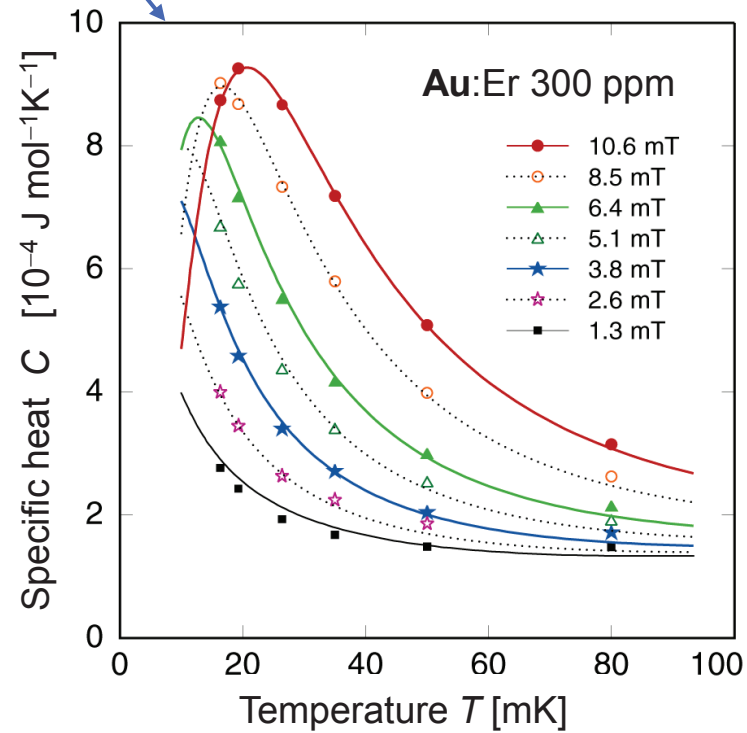
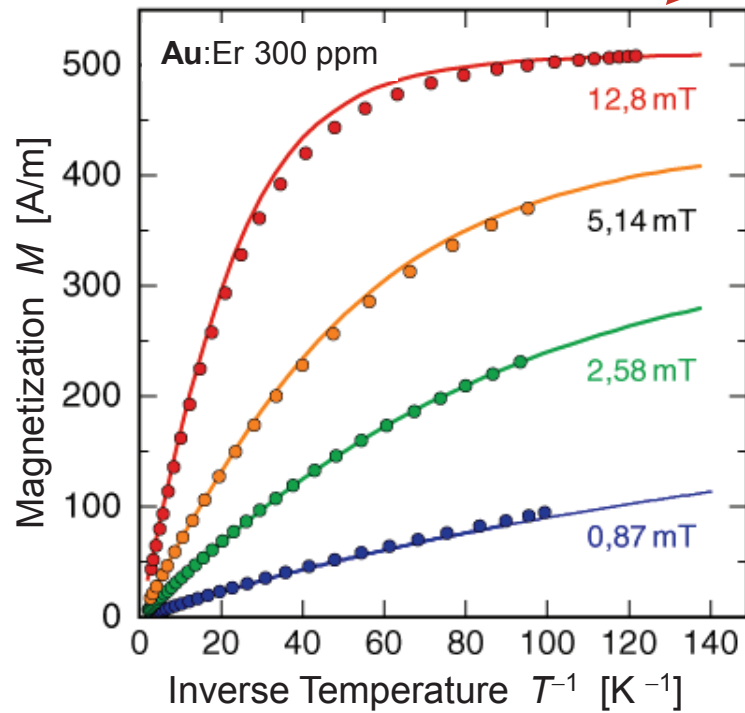
no dissipation in the sensor

no galvanic contact to the sensor

MMC: signal size

Numerical calculations based on mean field approximation are used to describe thermodynamical properties of interacting spins (RKKY)

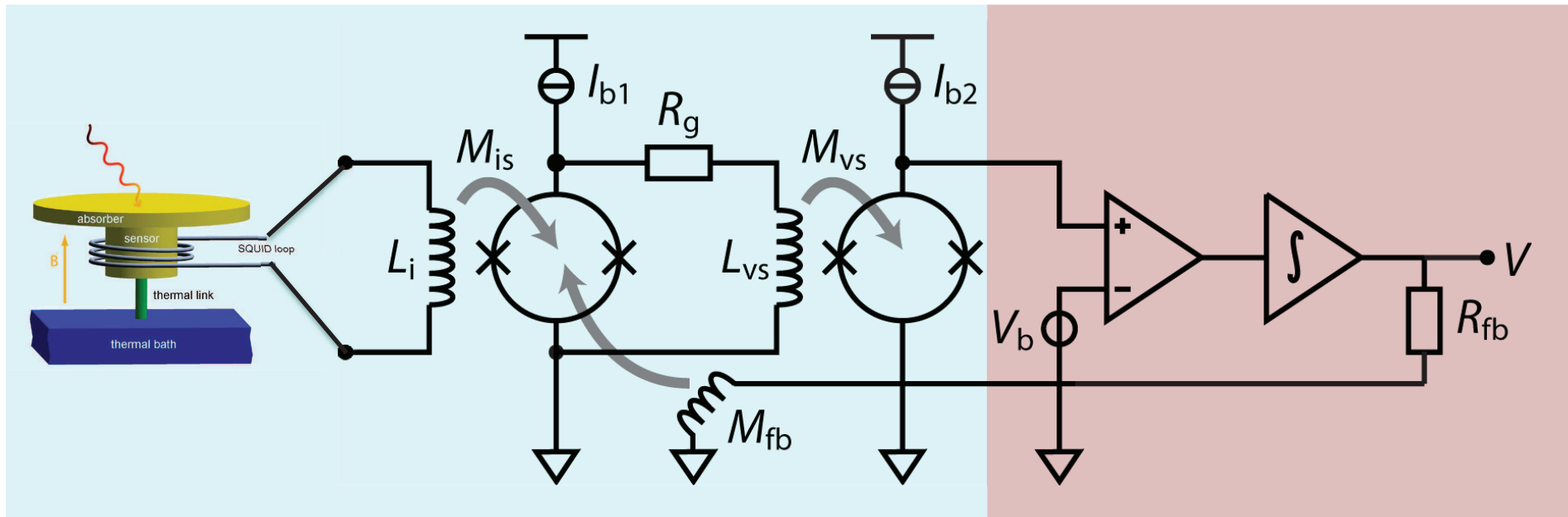
$$\delta\Phi_S \propto \frac{\partial M}{\partial T} \frac{1}{C_{\text{tot}}} \delta E$$



MMCs: Readout

$T \sim 30 \text{ mK}$

$T \sim 300 \text{ mK}$

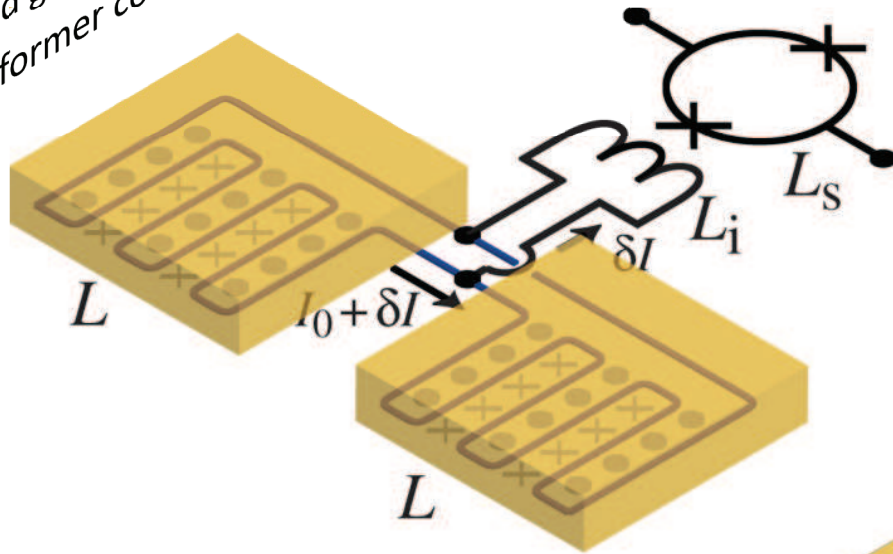


Two-stage SQUID setup with flux locked loop allows for:

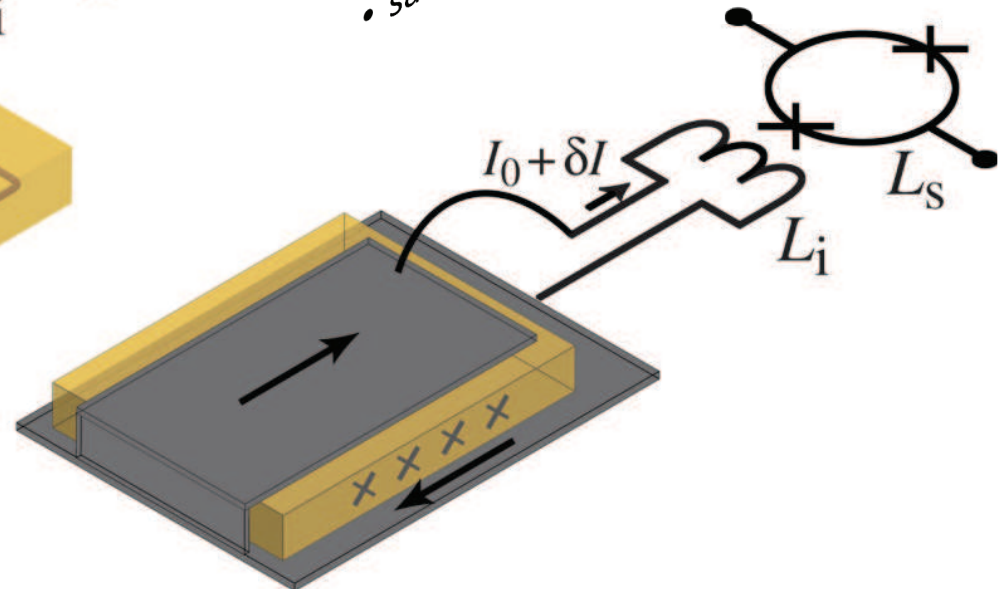
- low noise
- large bandwidth / slewrate
- small power dissipation on detector SQUID chip (voltage bias)

MMCs: Geometries

- Well established:*
- planar T -sensor
 - superconducting meander shaped pickup loop
 - B-field generated by persistent current
 - transformer coupled to SQUID



- New sandwich geometry:*
- best magn. flux coupling,
 - planar sensor
 - sandwiched between stripline

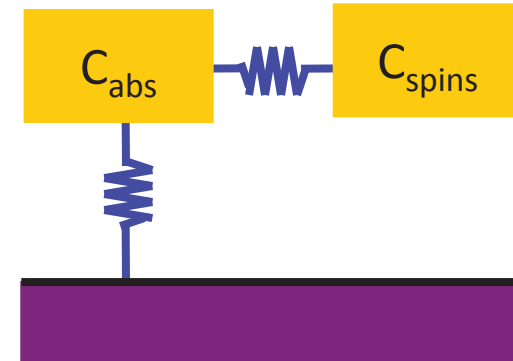


Energy resolution

- fluctuations of energy between sub-systems

$$\Delta E_{\text{FWHM}} \simeq 2,36 \sqrt{4k_B C_{\text{Abs}} T^2} \sqrt{2} \left(\frac{\tau_0}{\tau_1} \right)^{1/4}$$

(optimum for $C_{\text{abs}} = C_{\text{spins}}$)



- flux noise of SQUID-magnetometer

$$S_{\Phi} = 2 \varepsilon L, \quad \text{required:} \quad \varepsilon < 50 \hbar \dots 300 \hbar$$

- magnetic Johnson noise

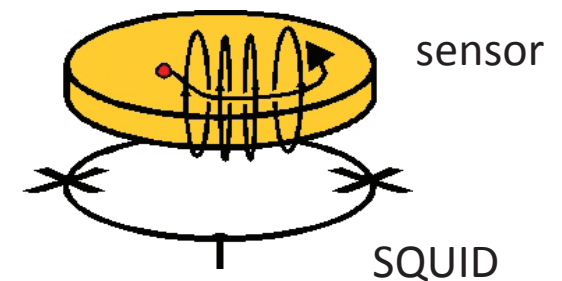
- thermal currents in the metallic components
- marginal in all present detectors

- excess noise

$$S_{\Phi} \sim N_{\text{Er}}$$

$$S_{\Phi} \sim 1/f, \quad S_m|_{1\text{Hz}} \approx 0.023 \mu_{\text{Er}}^2/\text{Hz}$$

temperature independent (20mK – 4K)



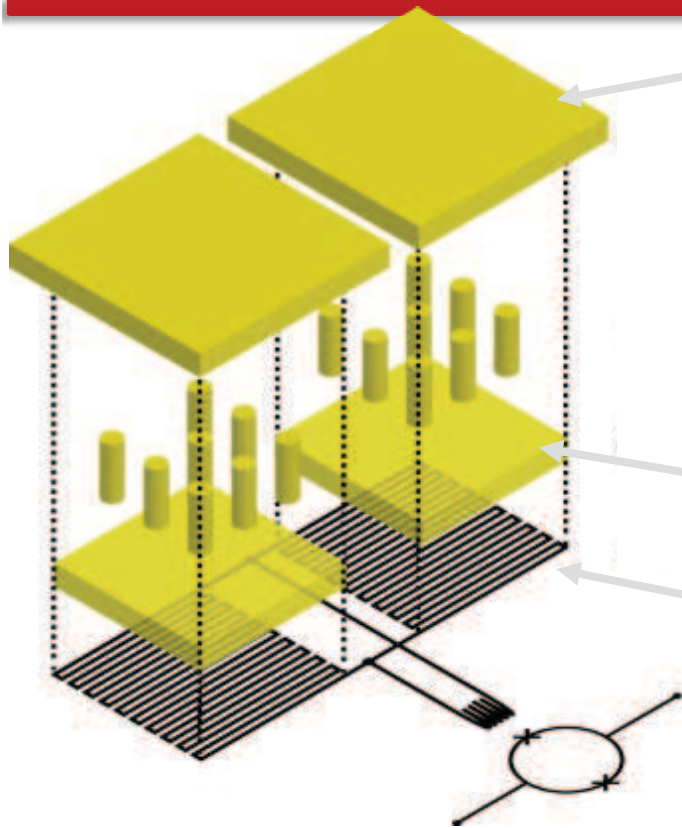
MMCs: Microfabrication

Cleanroom facilities at KIP



- Flexibility in design and fabrication
- Reliable processes for thin films
- Presently 10 different designs (6-16 layers) processed in parallel

maXs20: 1d-array for soft x-rays



- **1×8 x-ray absorbers**

- 250 μm ×250 μm gold, 5 μm thick
- 98% Qu.-Eff. @ 6 keV
- electroplated into photoresist mold (RRR>15)
- mech/therm contact to sensor by stems to prevent loss of initially hot phonons

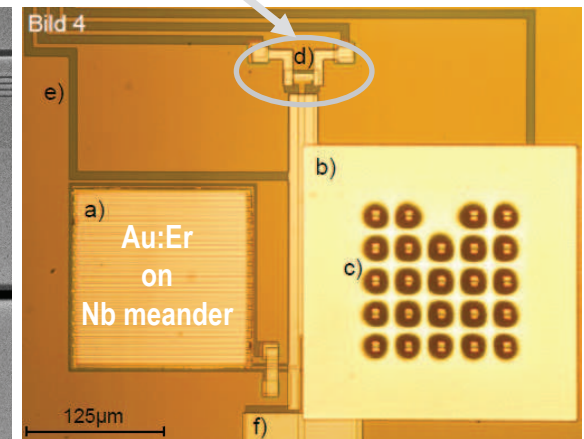
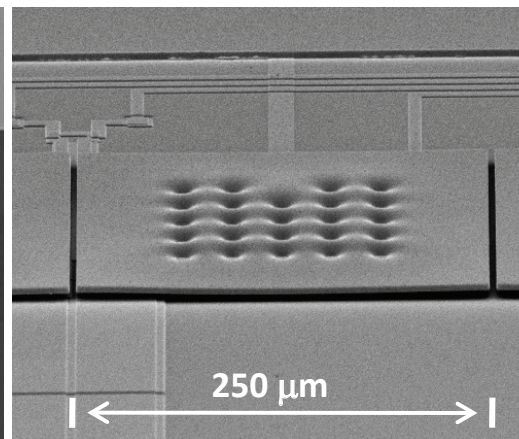
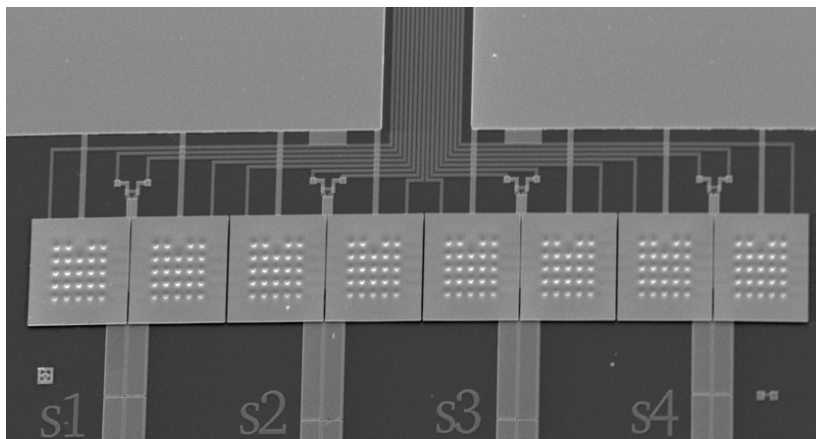
- **Au:¹⁶⁶Er_{300ppm} temperature sensors**

- co-sputtered from pure Au and high conc. AuEr target

- **Meander shaped pickup coils**

- 2.5 μm wide Nb lines
- $I_c \approx 100\text{mA}$

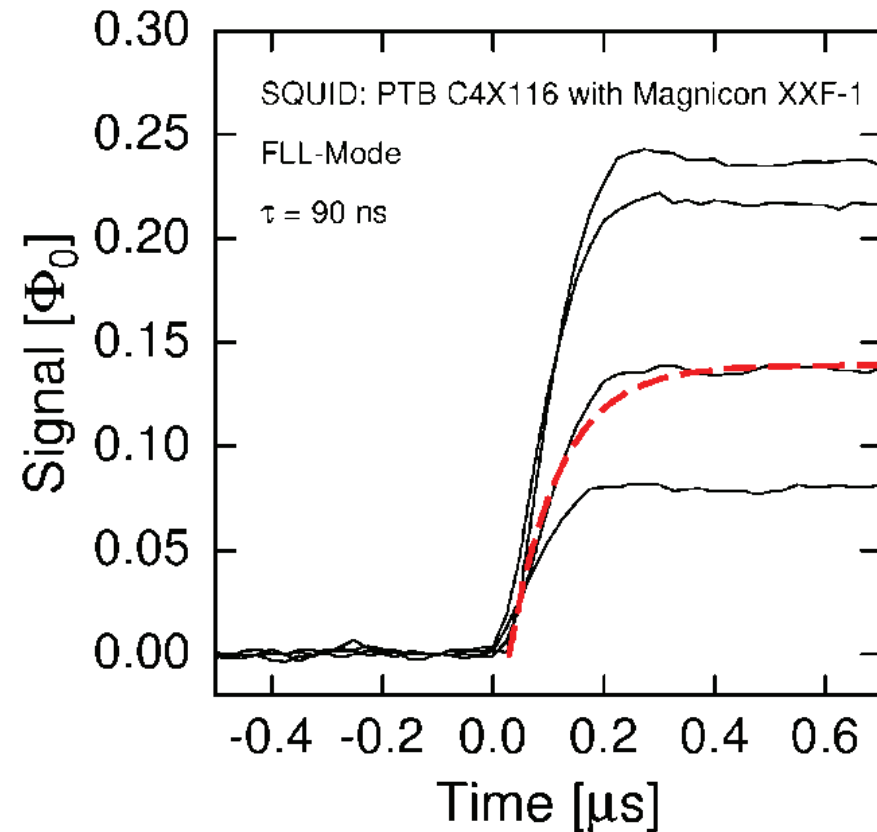
- **On-chip persistent current switch (AuPd)**



maXs20: 1d-array for soft x-rays

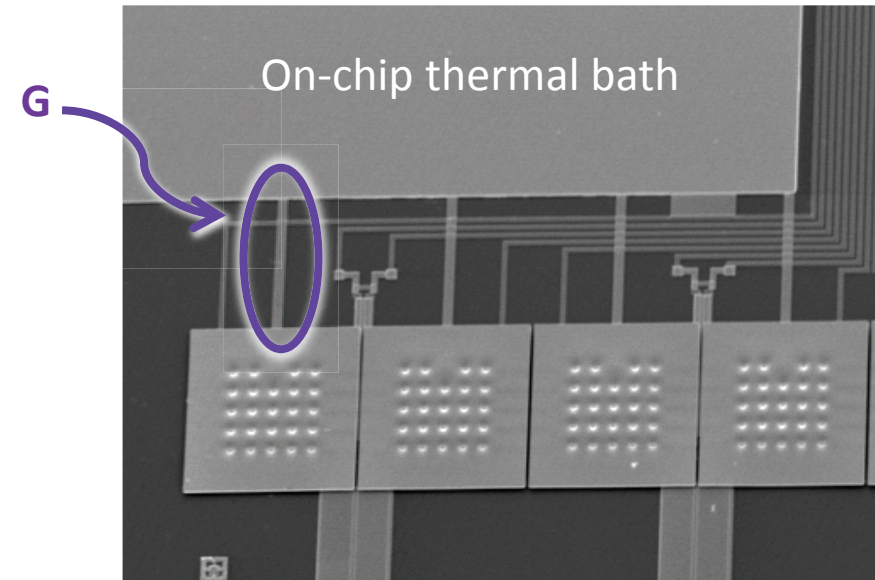
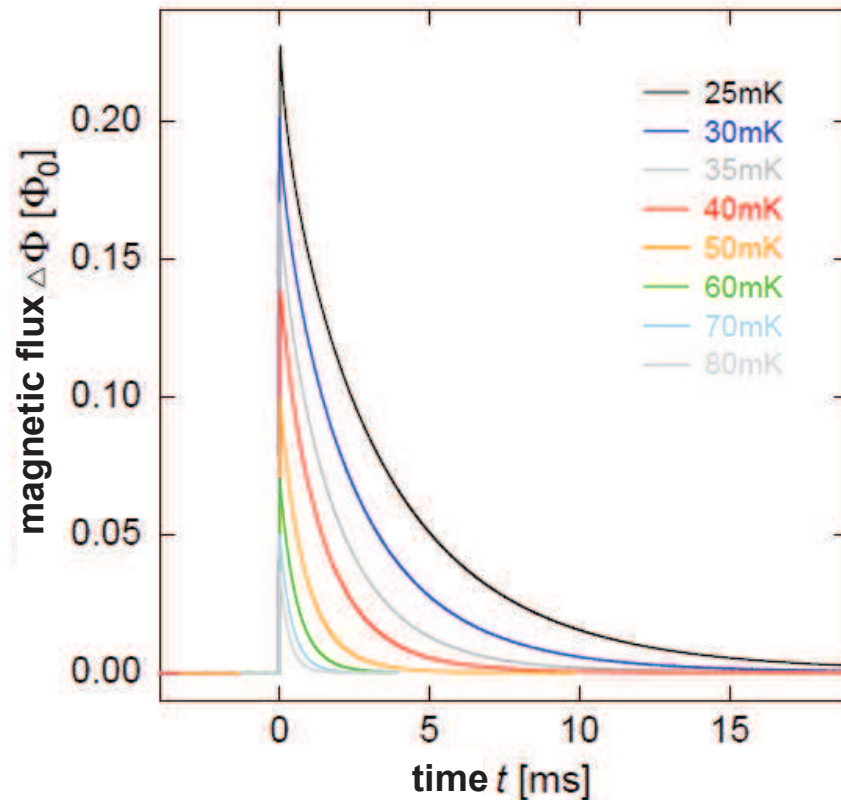
- **rise time: 90 ns @ 30 mK,**
as expected for the **spin-electron-relaxation**
from Korringa-constant of Er in Au

**Fastest rise-time among
 μ -cal for x-ray detection**



maXs20: 1d-array for soft x-rays

- **decay time** here: **3 ms @ 30 mK**
- nearly single exponential decay

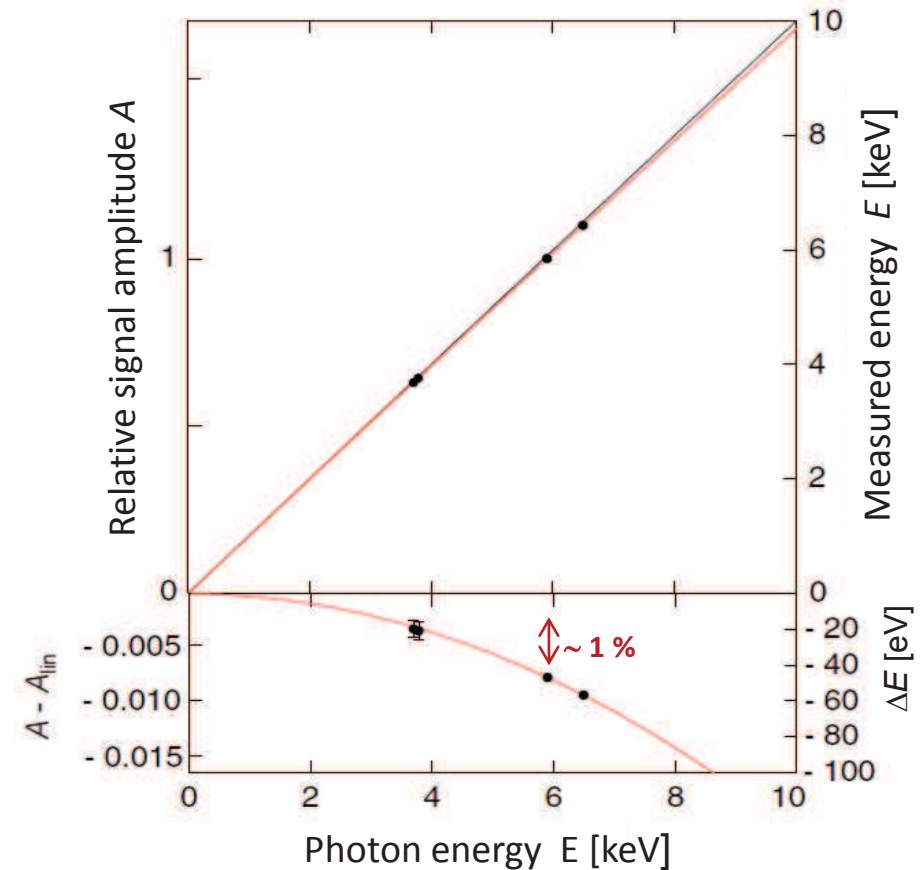


adjusted by sputtered
thermal link (Au)

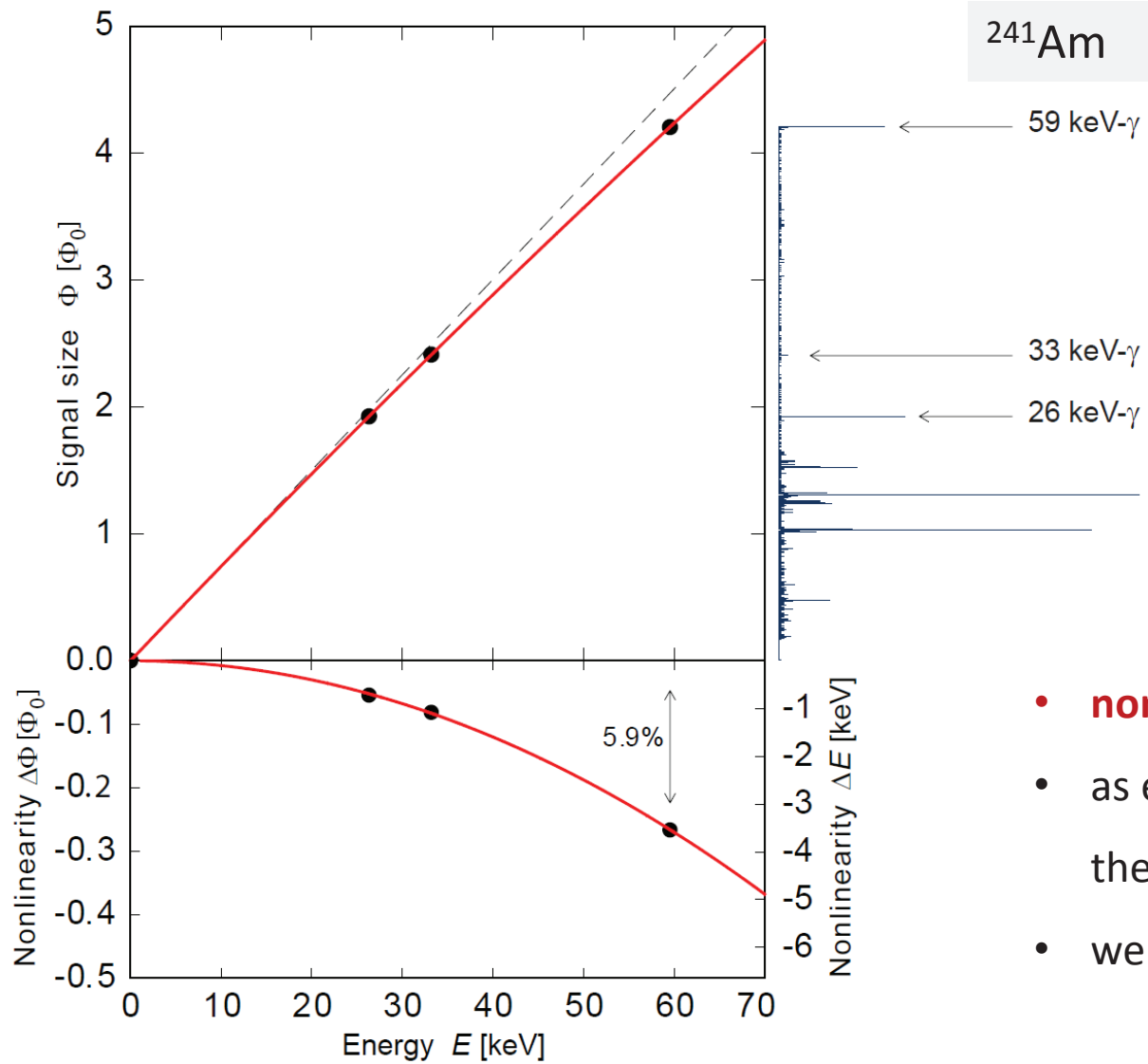
maXs20: 1d-array for soft x-rays

- **non-linearity: 1% at 6 keV**
- as expected from thermodynamical properties
- well described by quadratic term

The energy scale is defined with high precision



maXs20: 1d-array for soft x-rays

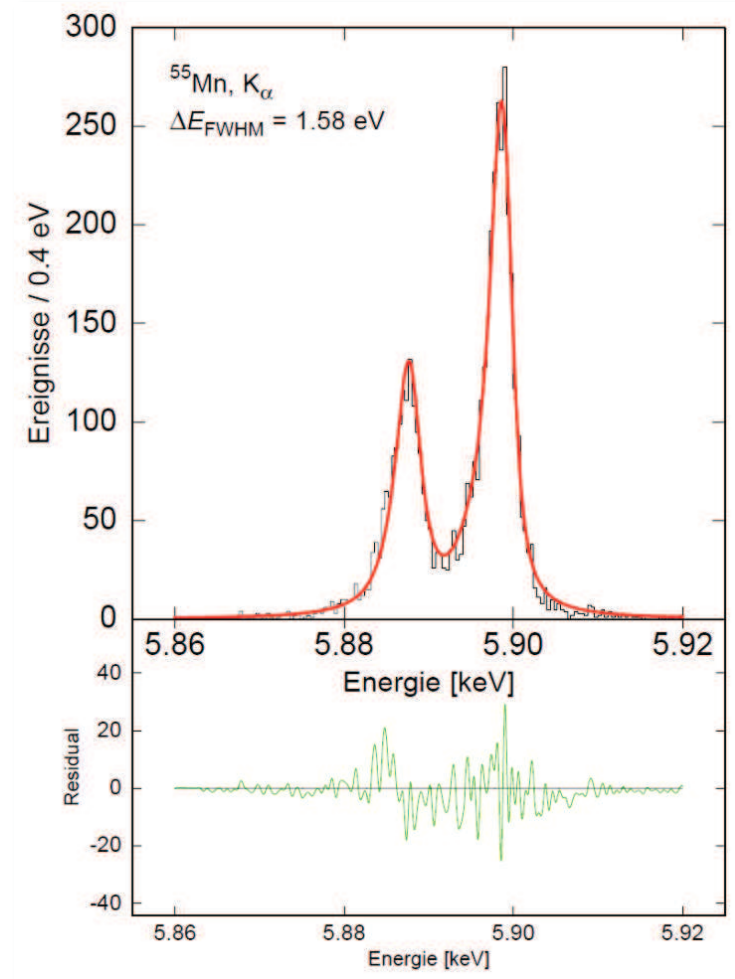


- **non-linearity: 6% at 60 keV**
- as expected from thermodynamical properties
- well described by quadratic term

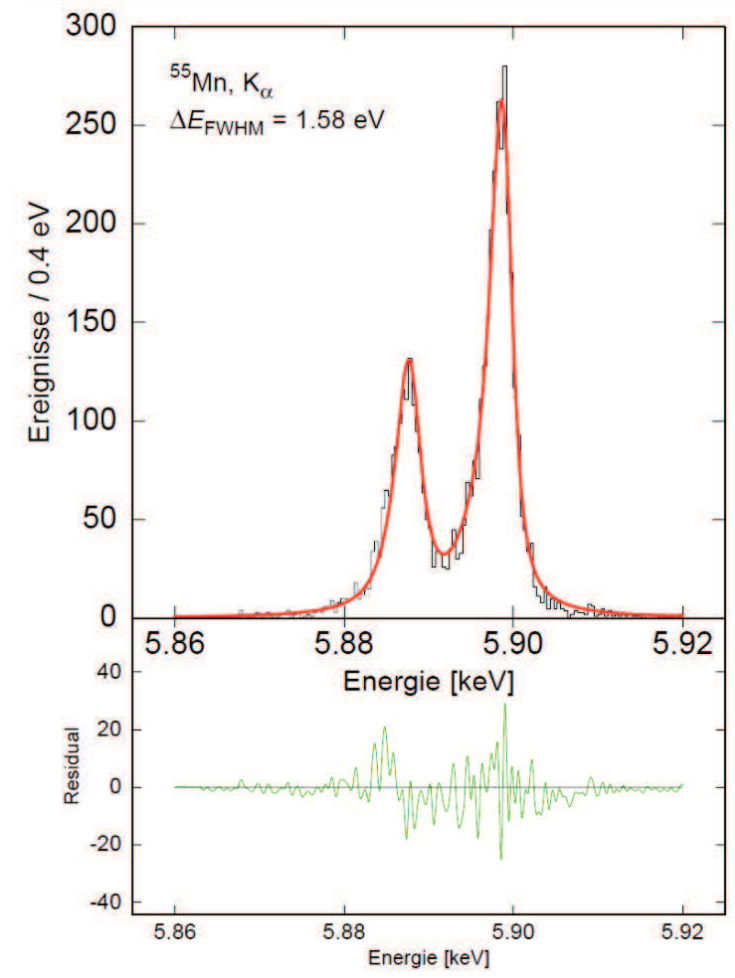
maXs20: 1d-array for soft x-rays ($T=20$ mK)

- Very good energy resolution

$$\Delta E_{\text{FWHM}} = 1.6 \text{ eV @ } 6 \text{ keV}$$



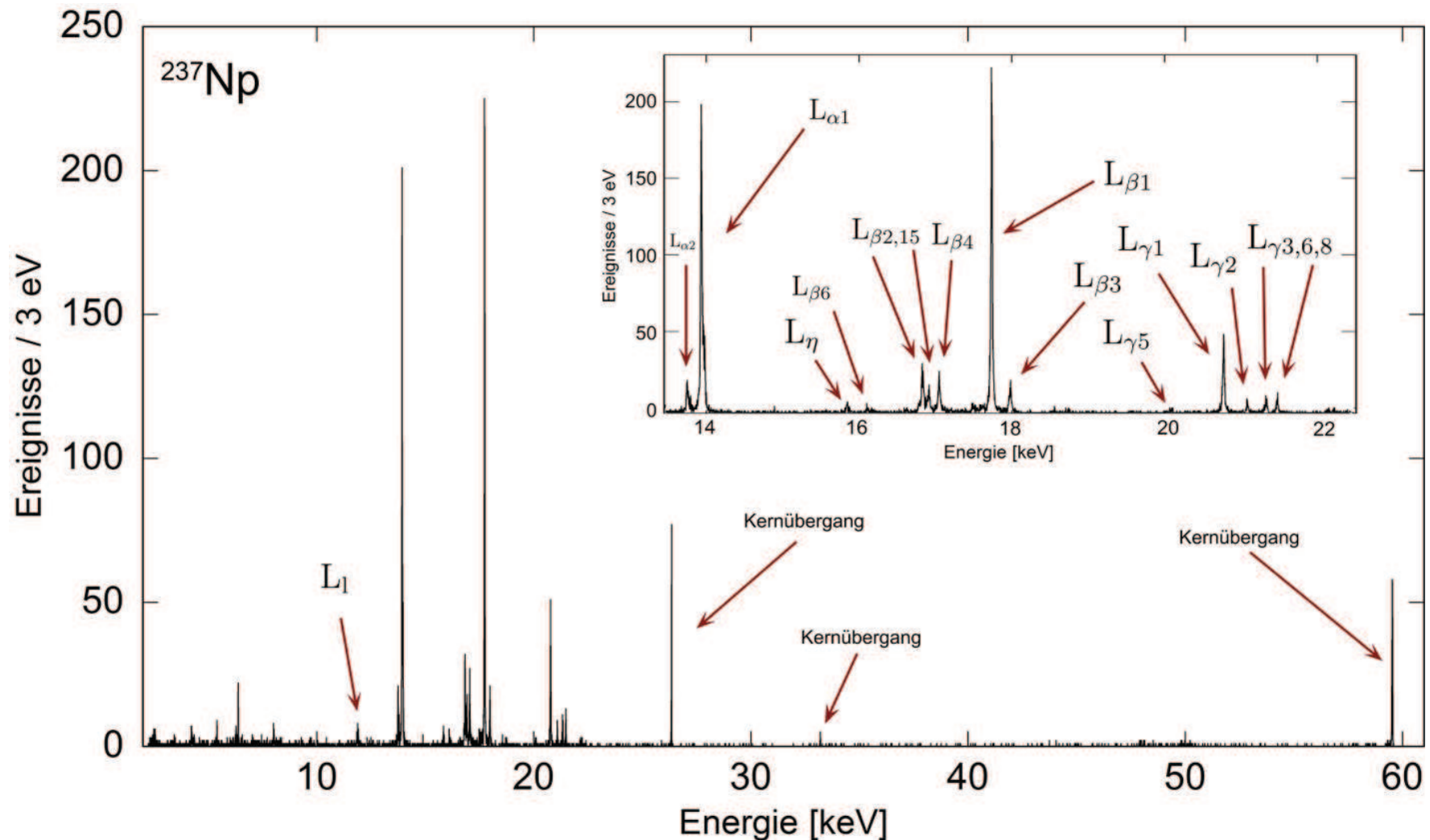
maXs20: 1d-array for soft x-rays ($T=20$ mK)



World record together with TES-sensors of NASA-GSFC!

maXs20: 1d-array for soft x-rays ($T=20$ mK)

- Large dynamic range



maXs200: 1d-array for hard x-rays

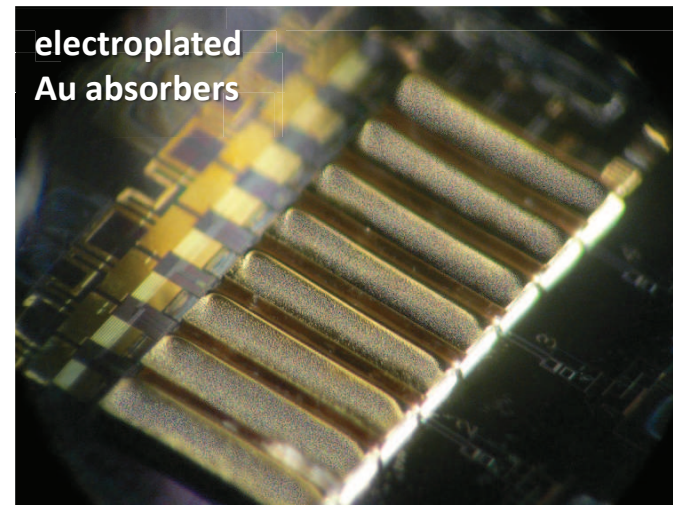
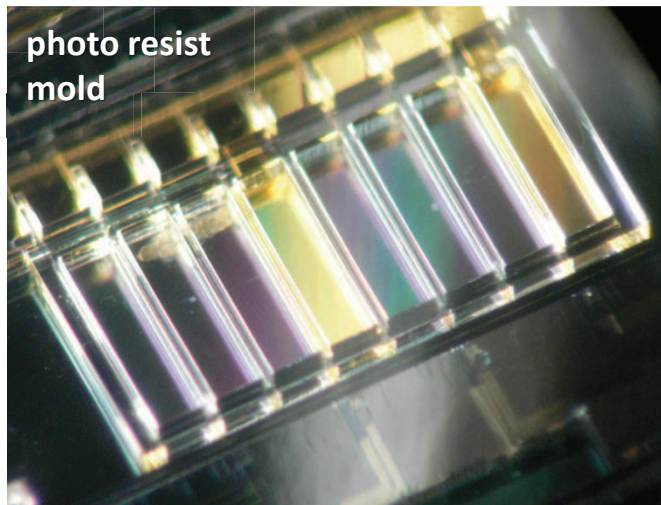
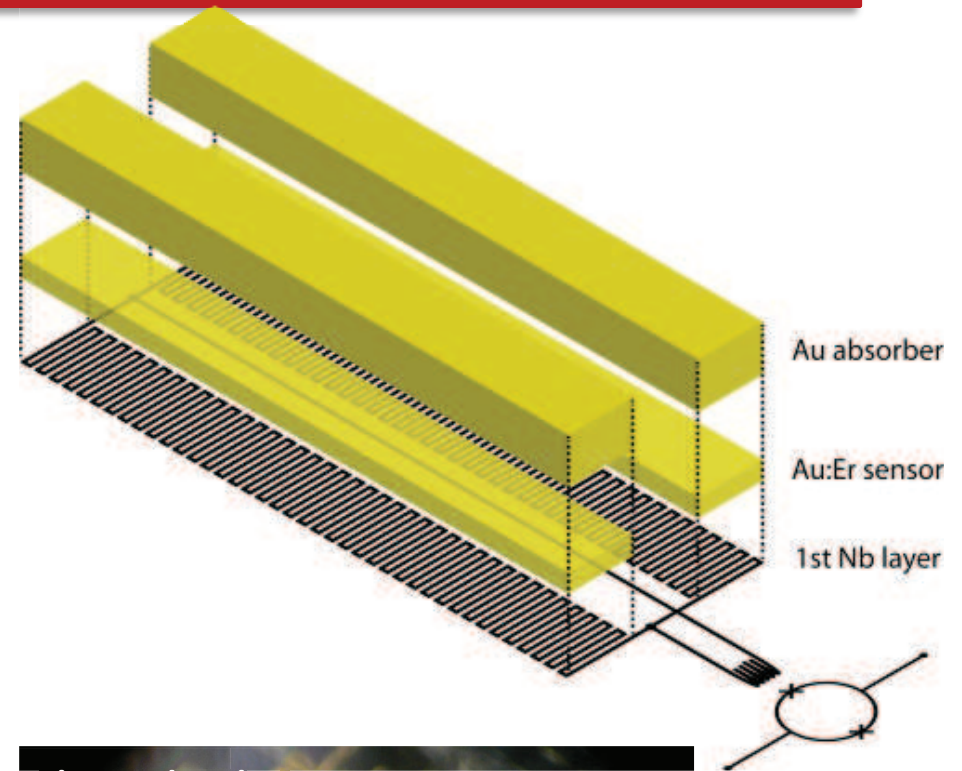
1x8 x-ray absorbers

2 mm x 0.5 mm

200 μm thick electroplated Au

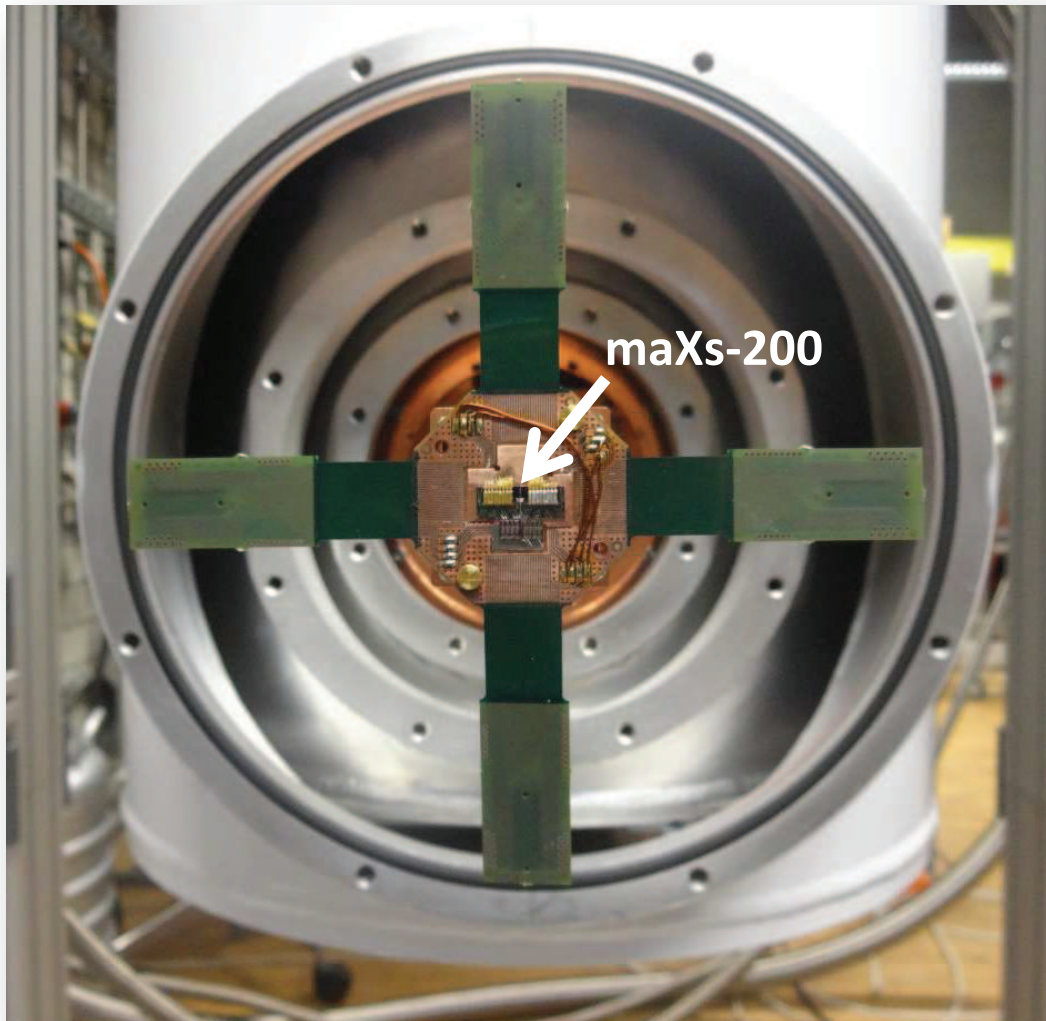
80% QE at 100 keV

$\Delta E_{\text{FWHM}} < 50 \text{ eV}$

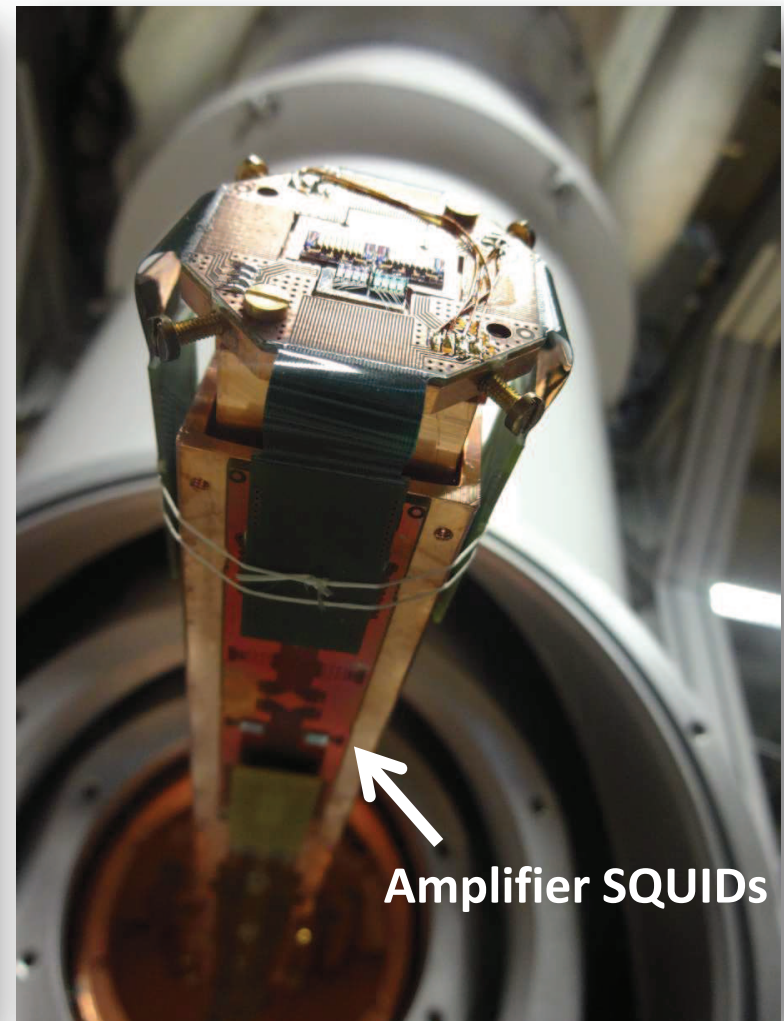


maXs200 on cold finger of a dry cryostat

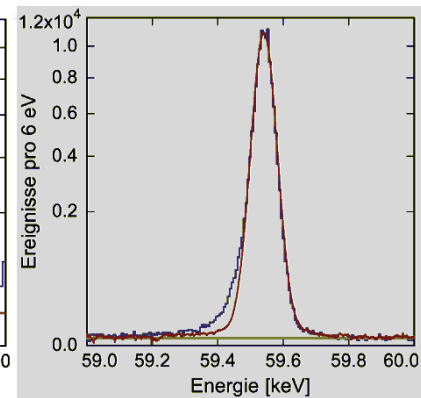
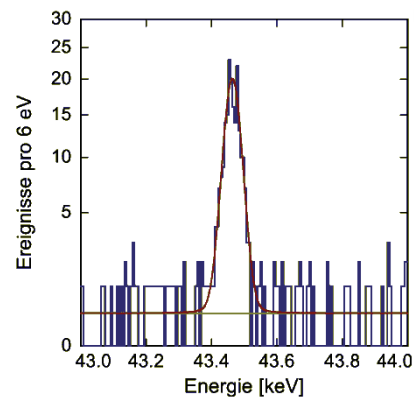
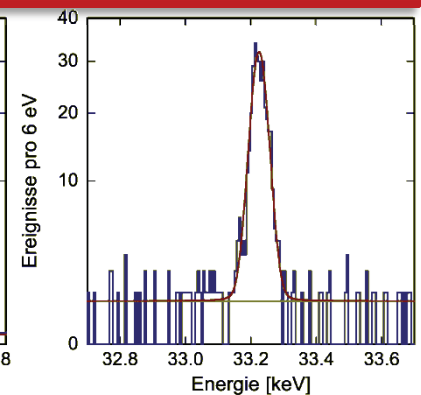
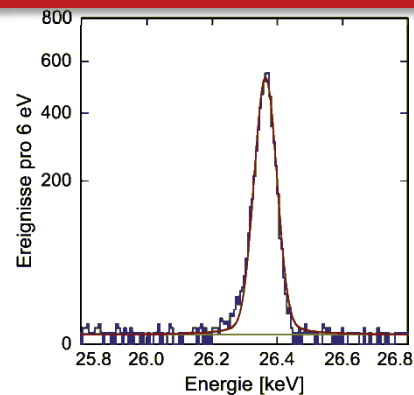
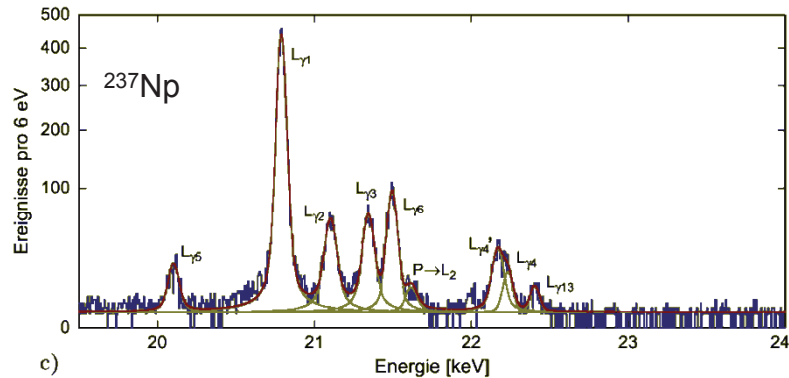
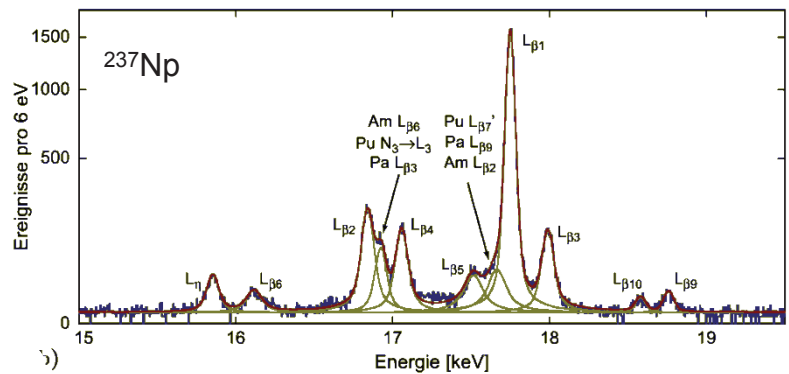
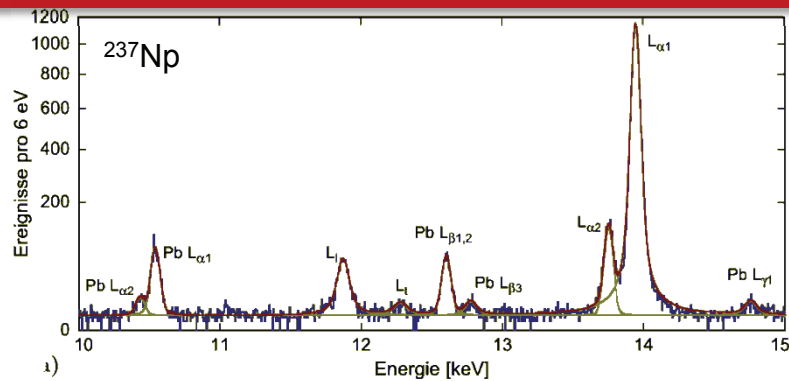
16 pixels **maXs-200** on detector platform



side arm without radiation shields



maXs200 – ²⁴¹Am calibration source



$$\Delta E_{FWHM} = 60 \text{ eV @ } 60 \text{ keV}$$

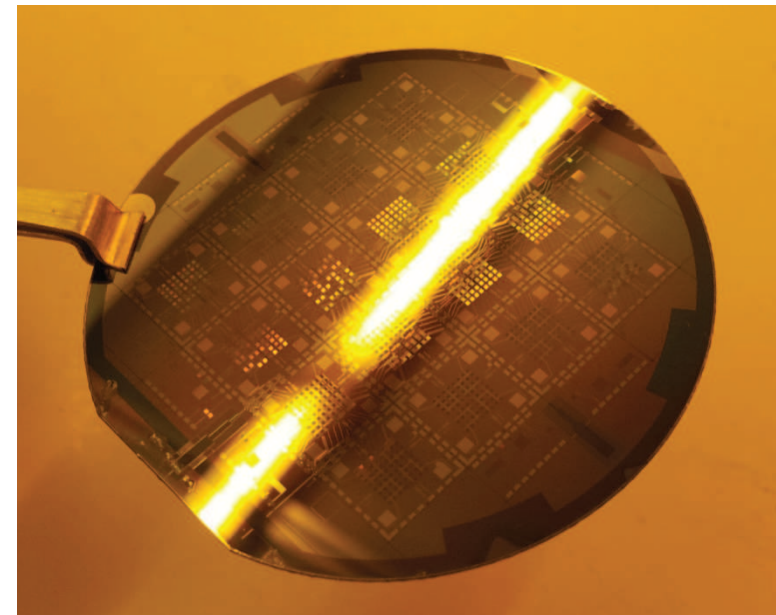
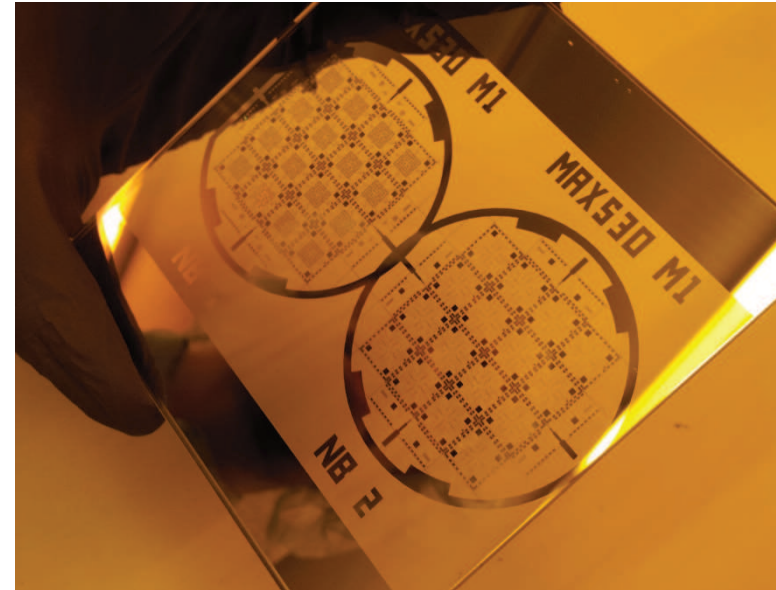
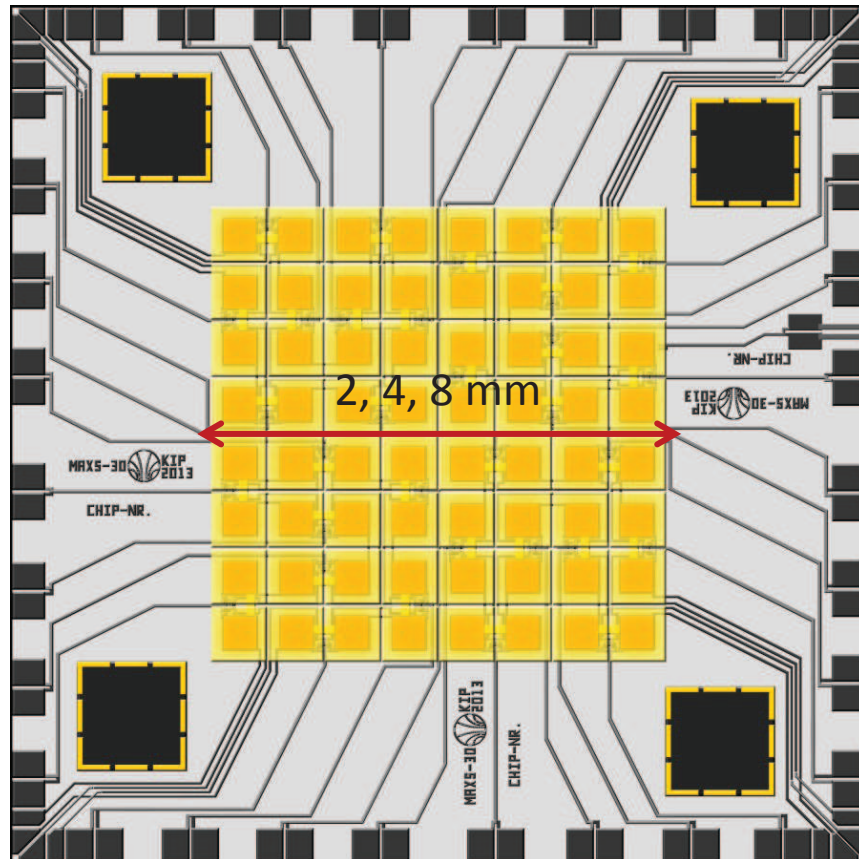
Slight degradation towards higher energies

due to position dependence,

Meanwhile fixed by stems between absorber and sensor

$$\Delta E_{FWHM} = 40 \text{ eV @ } 0\text{-}20 \text{ keV}$$

Large Arrays – parallel read-out

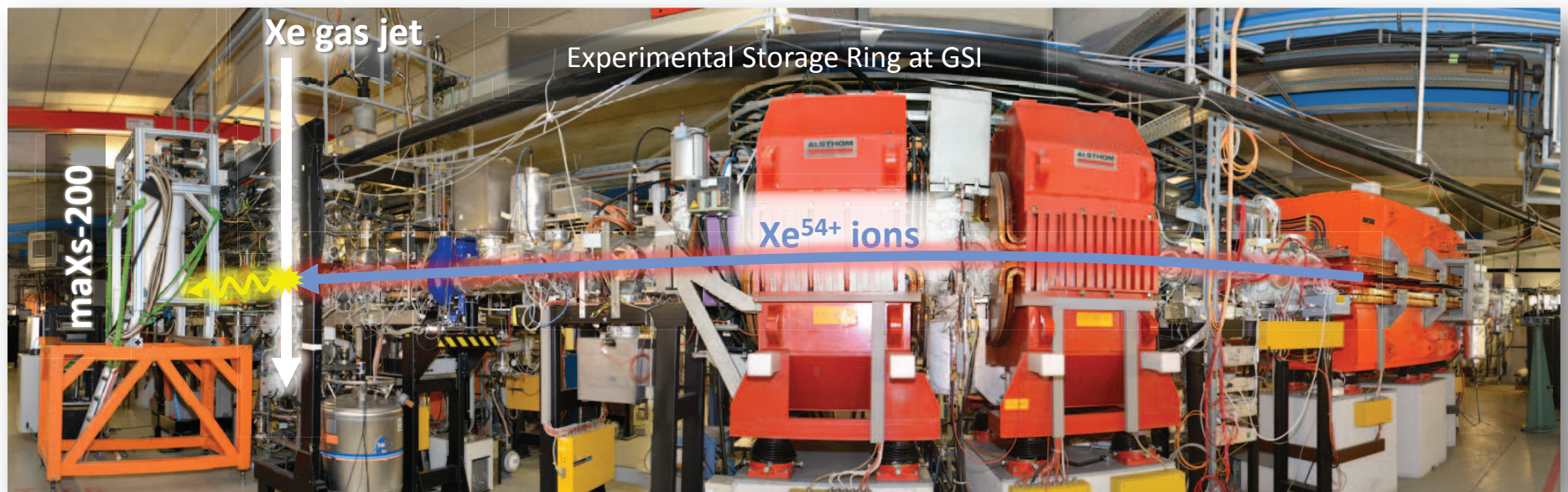
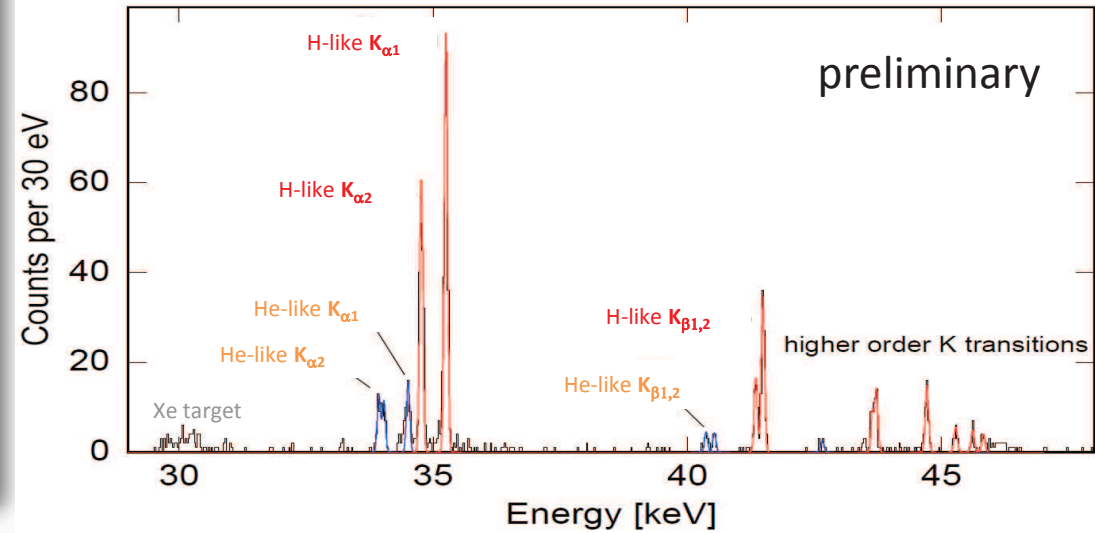


maXs-20/30/200:

- 8x8 pixels for photons up to 20/30/200 keV
- with $\Delta E_{\text{FWHM}} = 2/5/30$ eV
- 32 two-stage dc-SQUIDS

MMCs on the road to GSI- Darmstadt

Heidelberg University, Friedrich-Schiller University Jena
and Helmholtz-Institute Jena for SPARC



MMCs meet massive particles from CSR at MPIK

Creation of lattice defects might spoil energy resolution drastically !

→ **first linewidth tests with ions** onto maXs-200 detectors at MPI-K

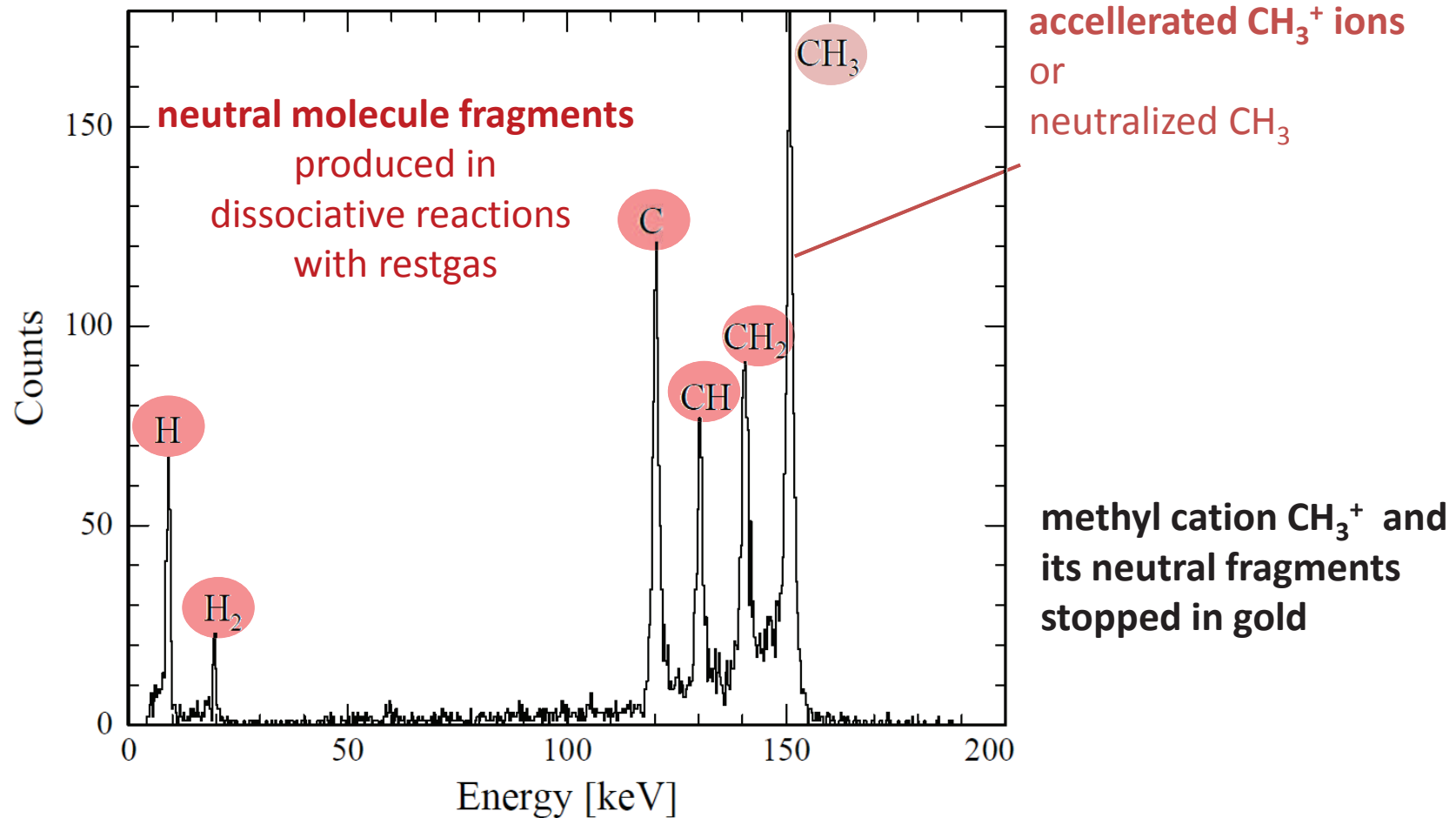


maXs-200
@ 20 mK

- atomic ions H^+ , C^+ , O^+ , Ar^+
- molecular ions N_2^+ , Ar_2^+ , CH_3^+ , Acetone $^+$

... accelerated by 0...150 kV

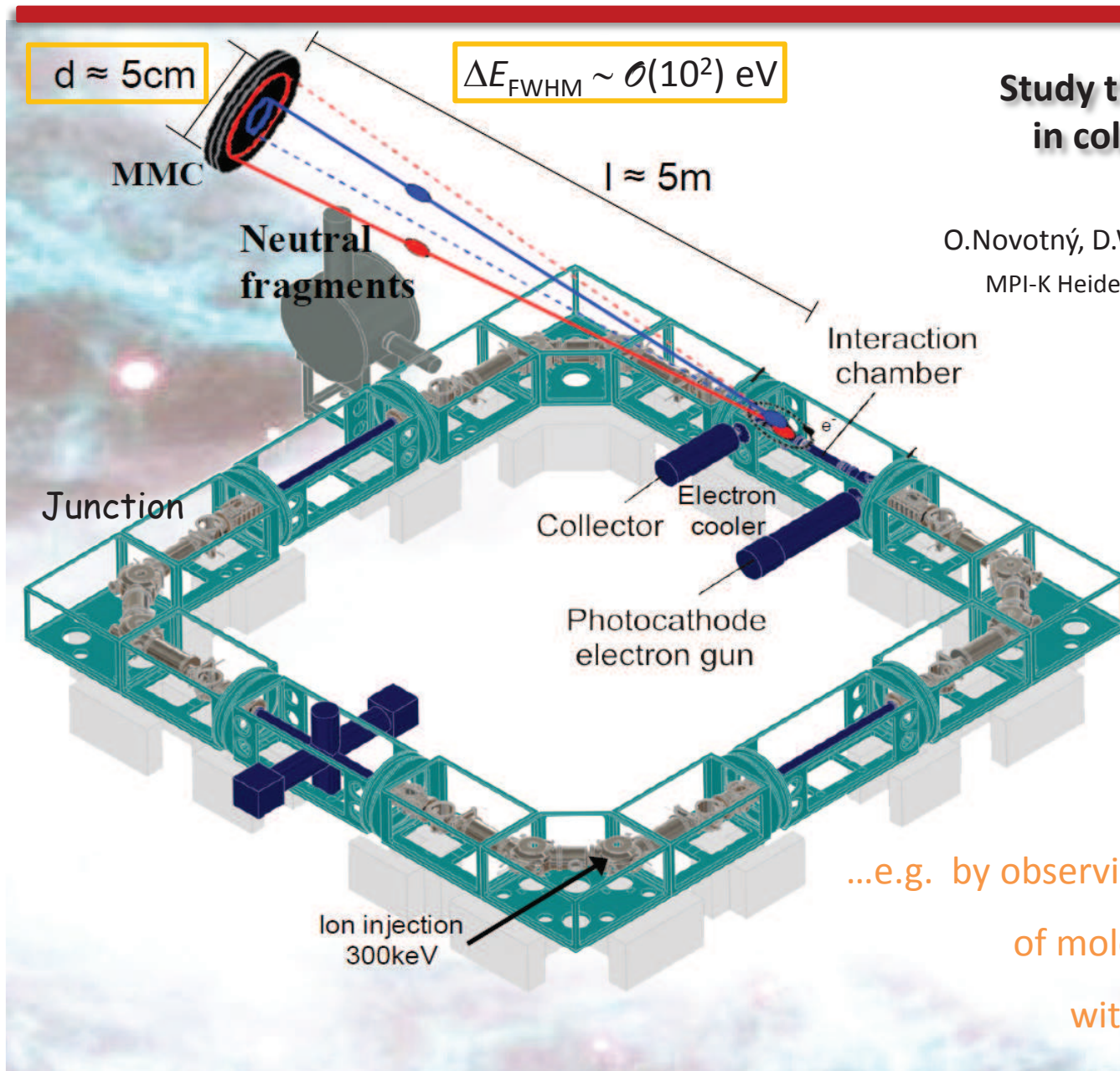
MMCs meet massive particles from CSR at MPIK



- All fragments clearly identified by mass (= kinetic energy)
- $\Delta E_{\text{FWHM}} = 750 \text{ eV @ } 150 \text{ keV}$ allows full identification for molecule masses up to 200

Sometimes nature is more friendly than the simulation ☺

Cold chemistry @ CSR MPI-K



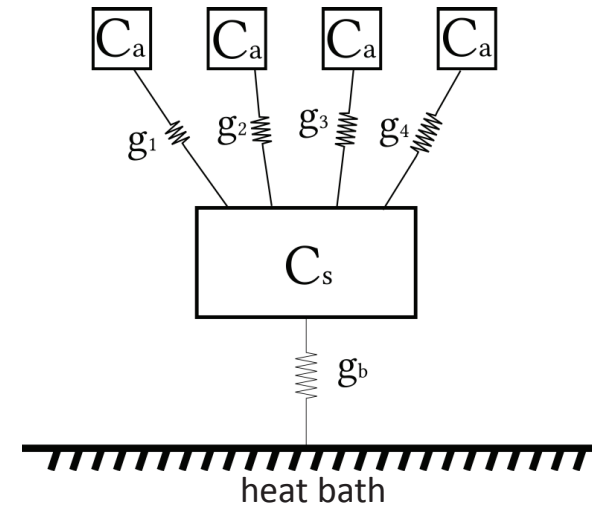
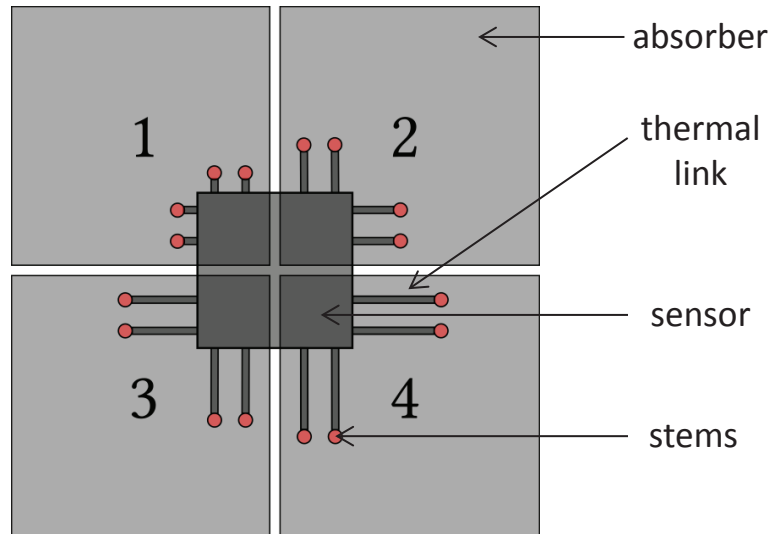
**Study the complex chemistry
in cold interstellar clouds**

with

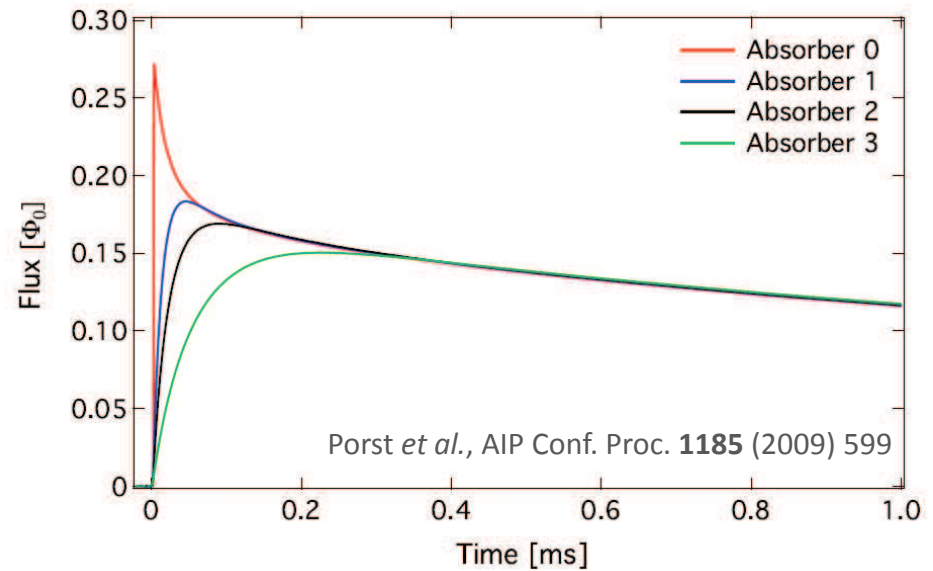
O. Novotný, D.W. Savin, C. Krantz, A. Wolf, K. Blaum,
MPI-K Heidelberg, Columbia Astrophysics Lab NY

...e.g. by observing neutral reaction products
of molecular ions after neutralizing
with an electron in the e-cooler

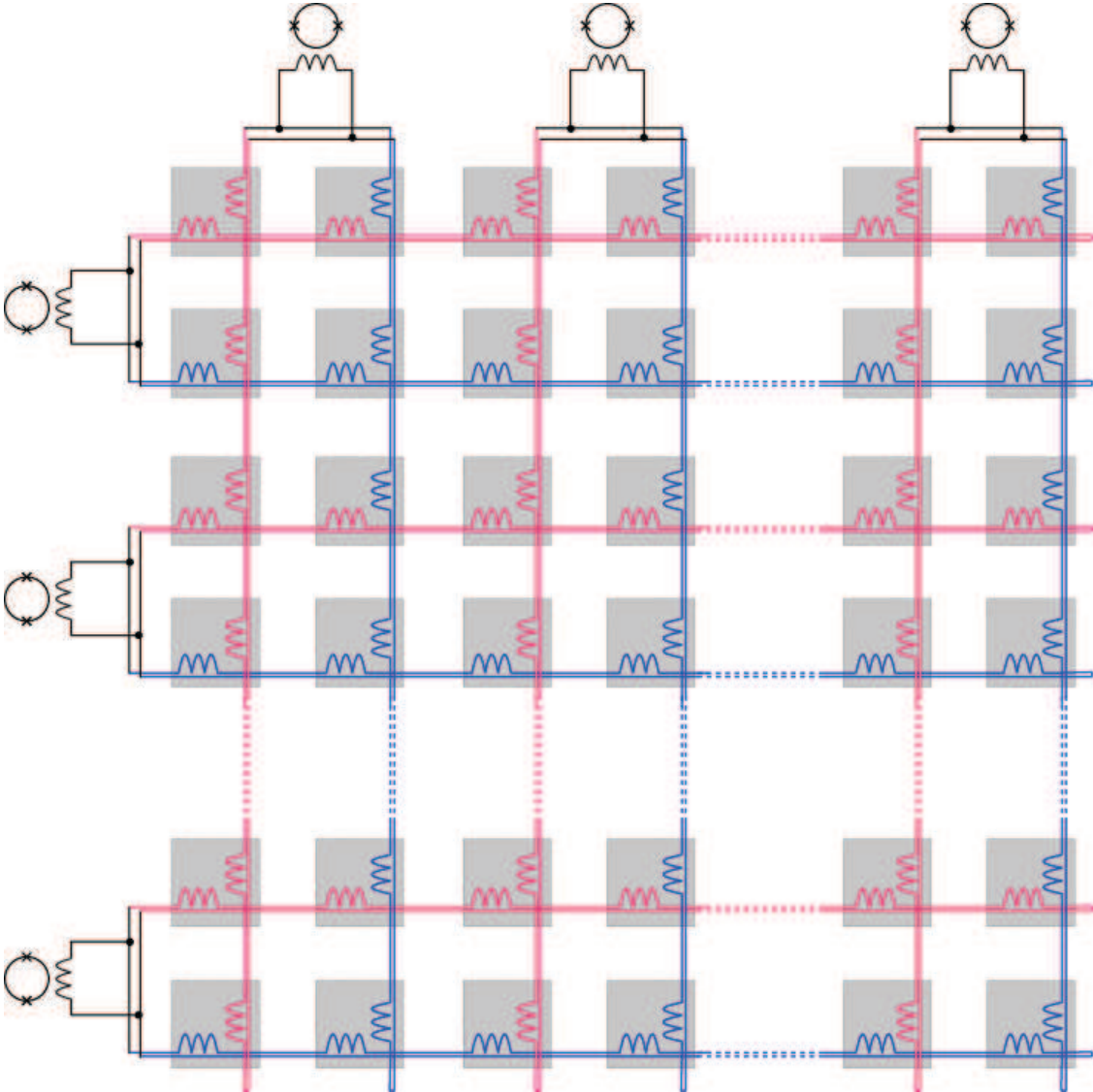
The Hydra principle



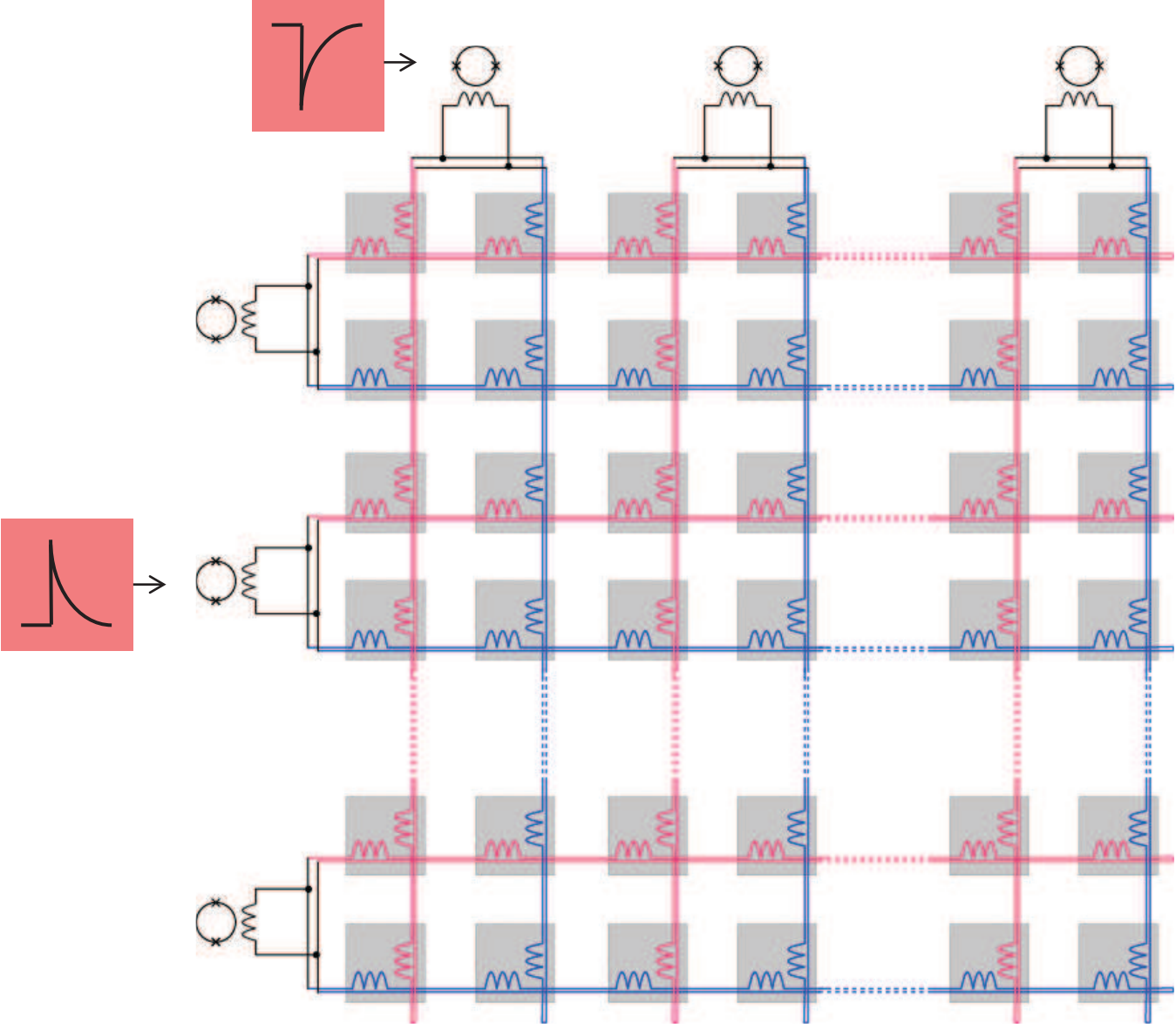
Pixel identification
via rise-time
of the detector signal



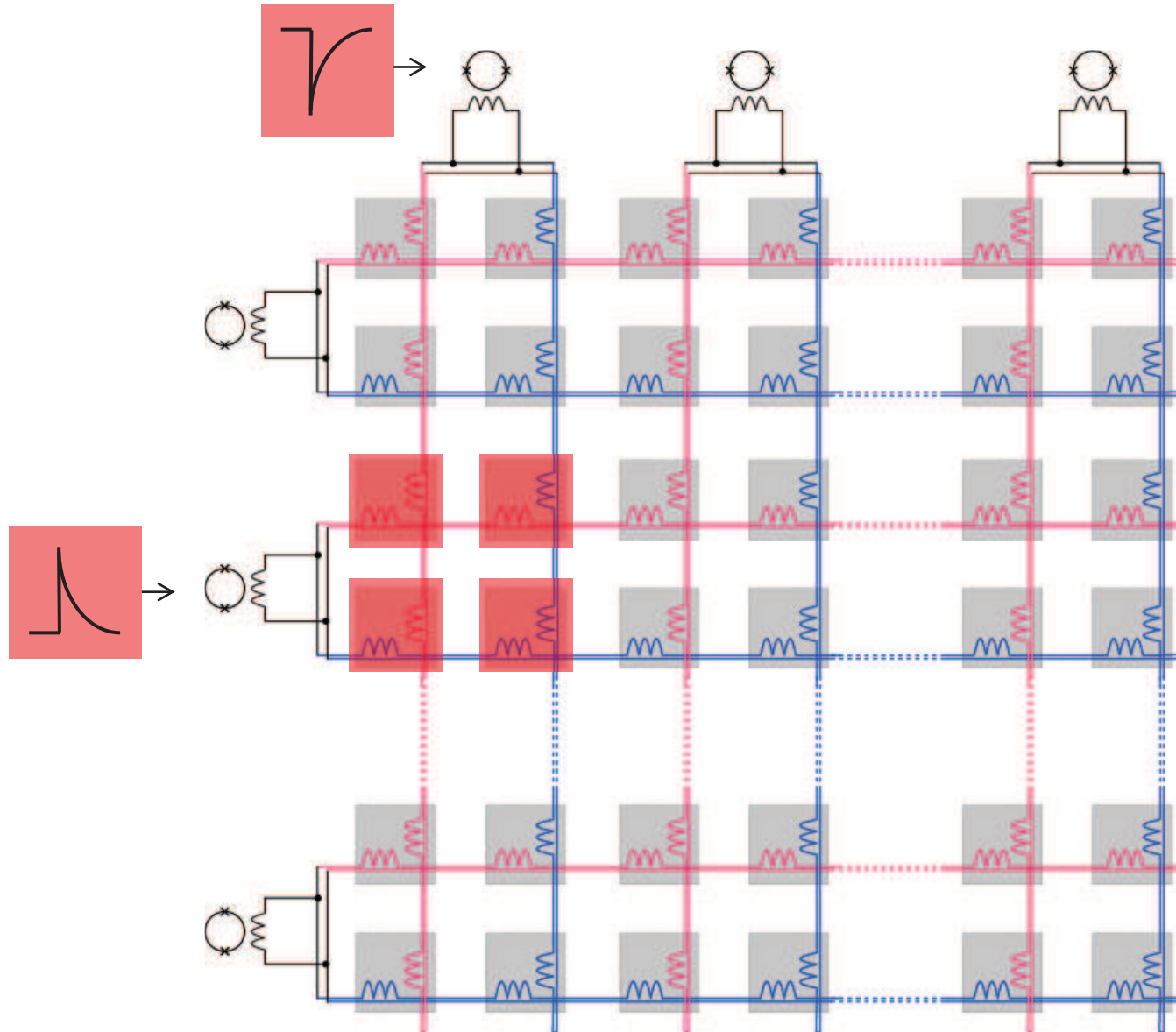
Read out scheme



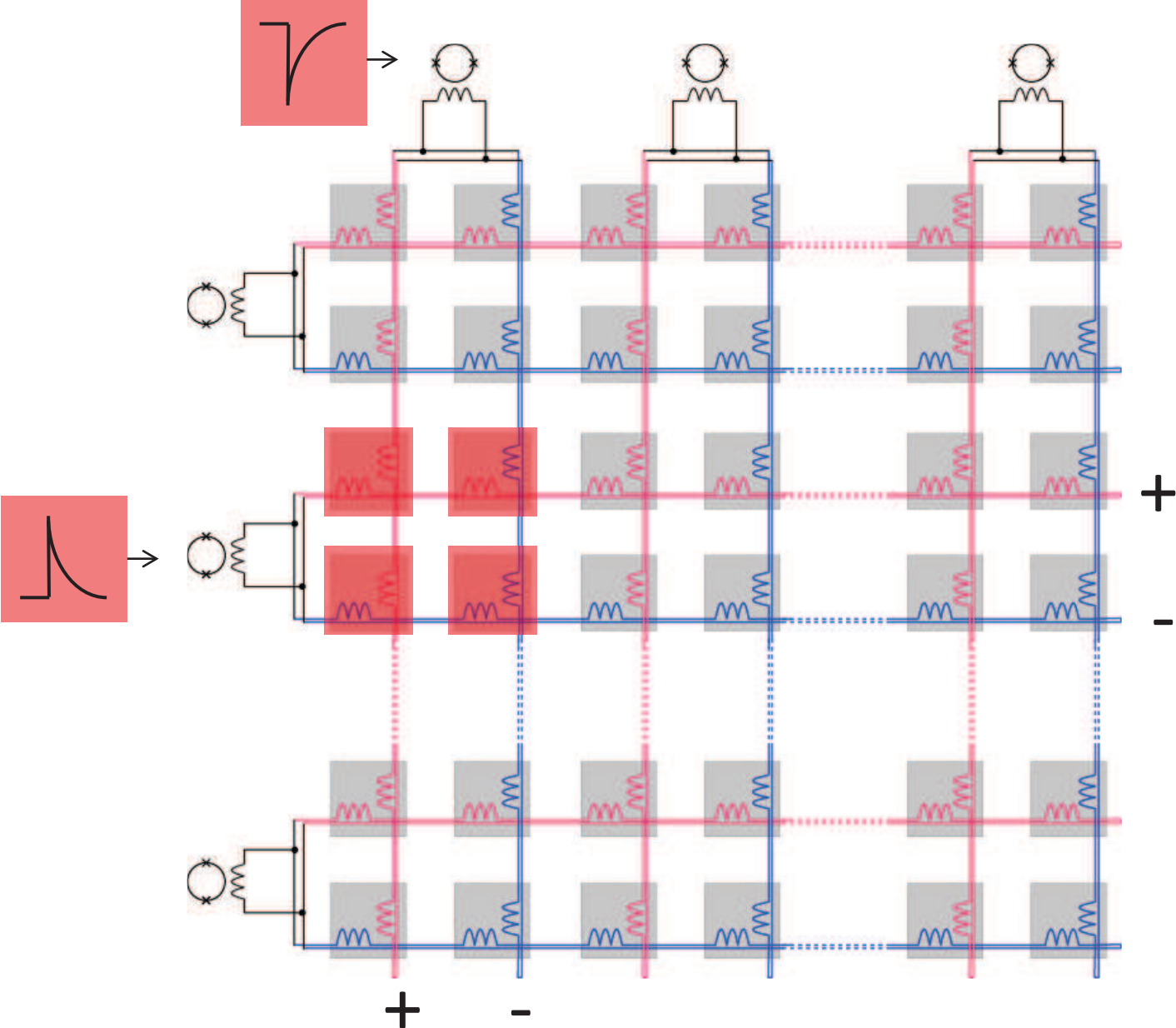
Read out scheme



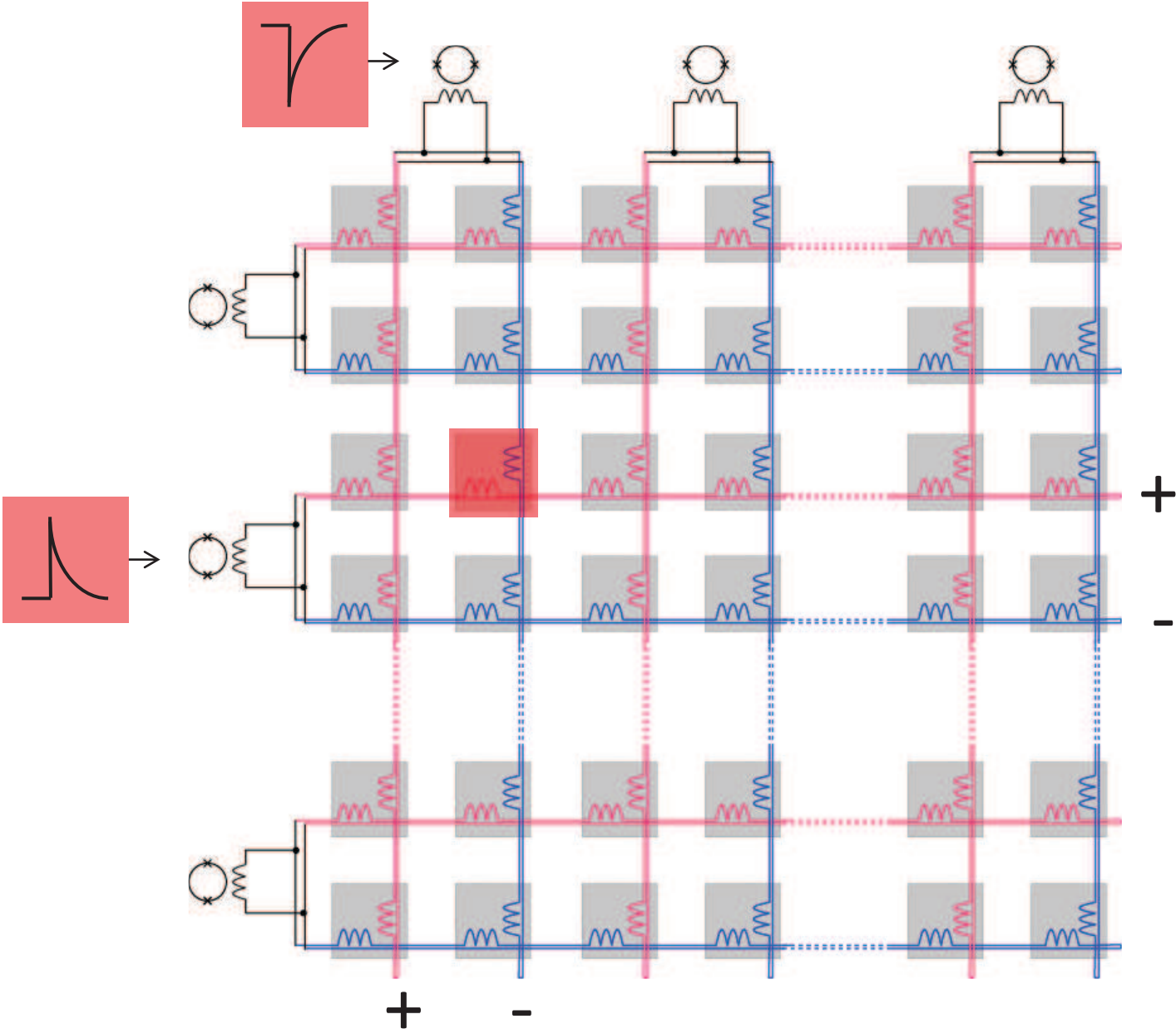
Read out scheme



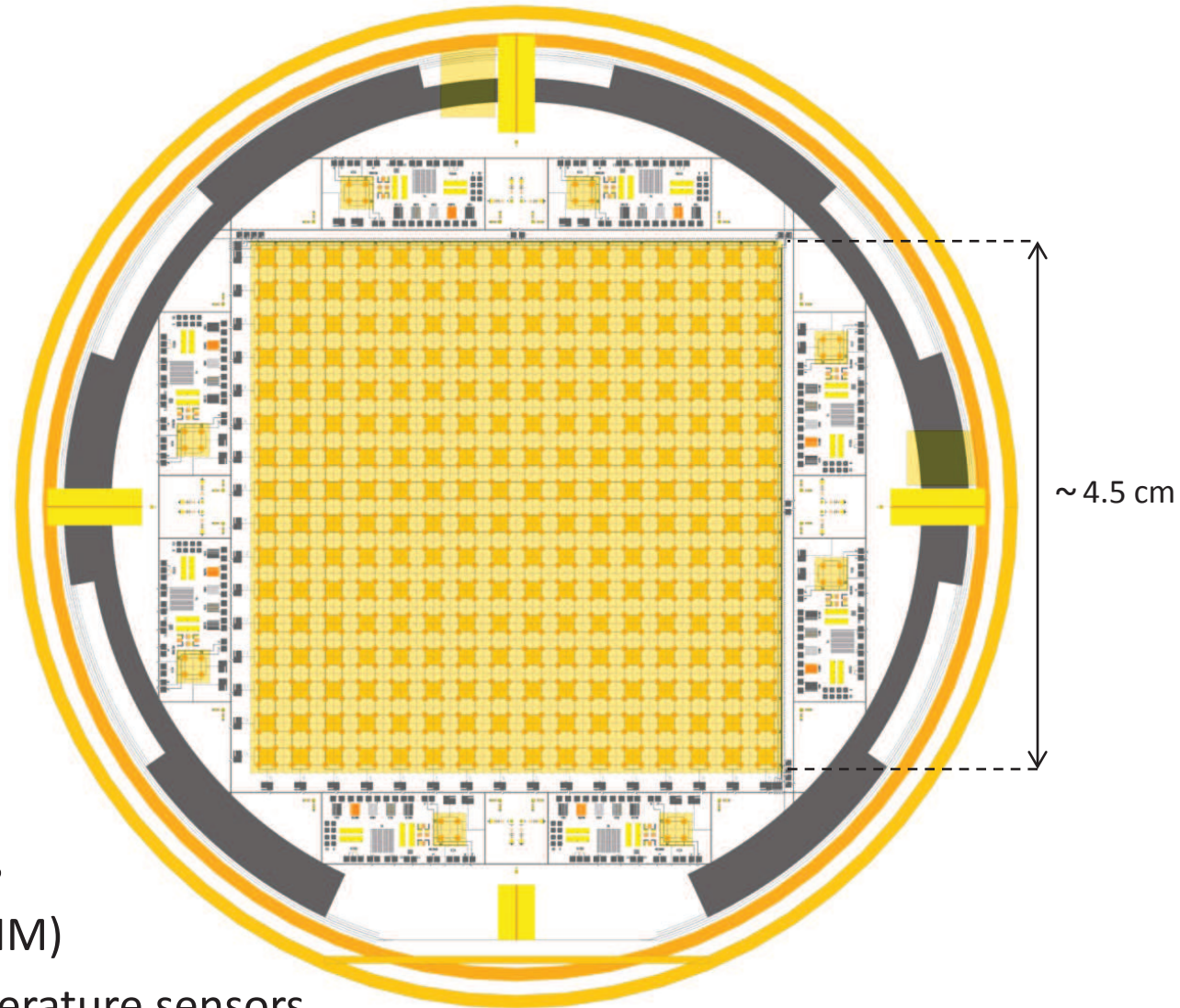
Read out scheme



Read out scheme



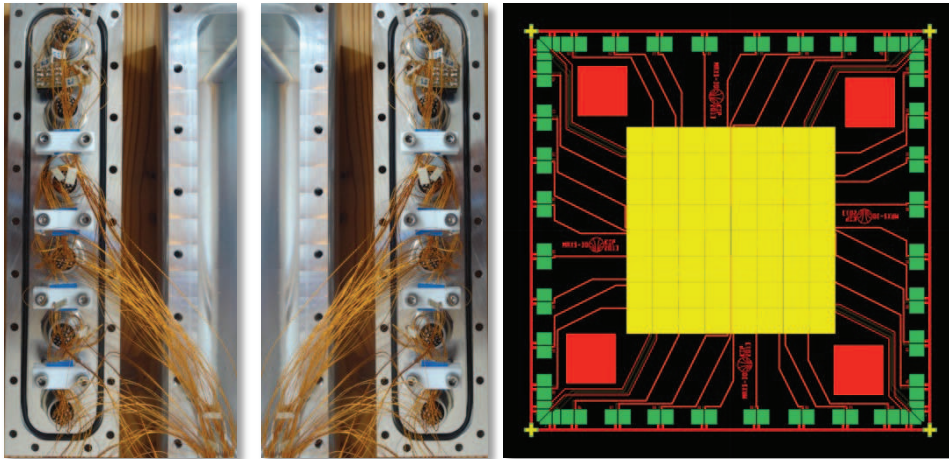
MOCCA: a 4k-pixel MMC camera



- 64×64 pixels
- 200 eV (FWHM)
- 32×32 temperature sensors
- Read out by 16+16 SQUIDs

How to read-out 1000s of detectors?

Simplest idea: duplicate single channel readout

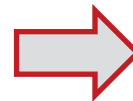


Brute force read-out of 8x8 x-ray detectors

But:

- number of wires
- parasitic heat load
- costs
- complexity

$\sim N$

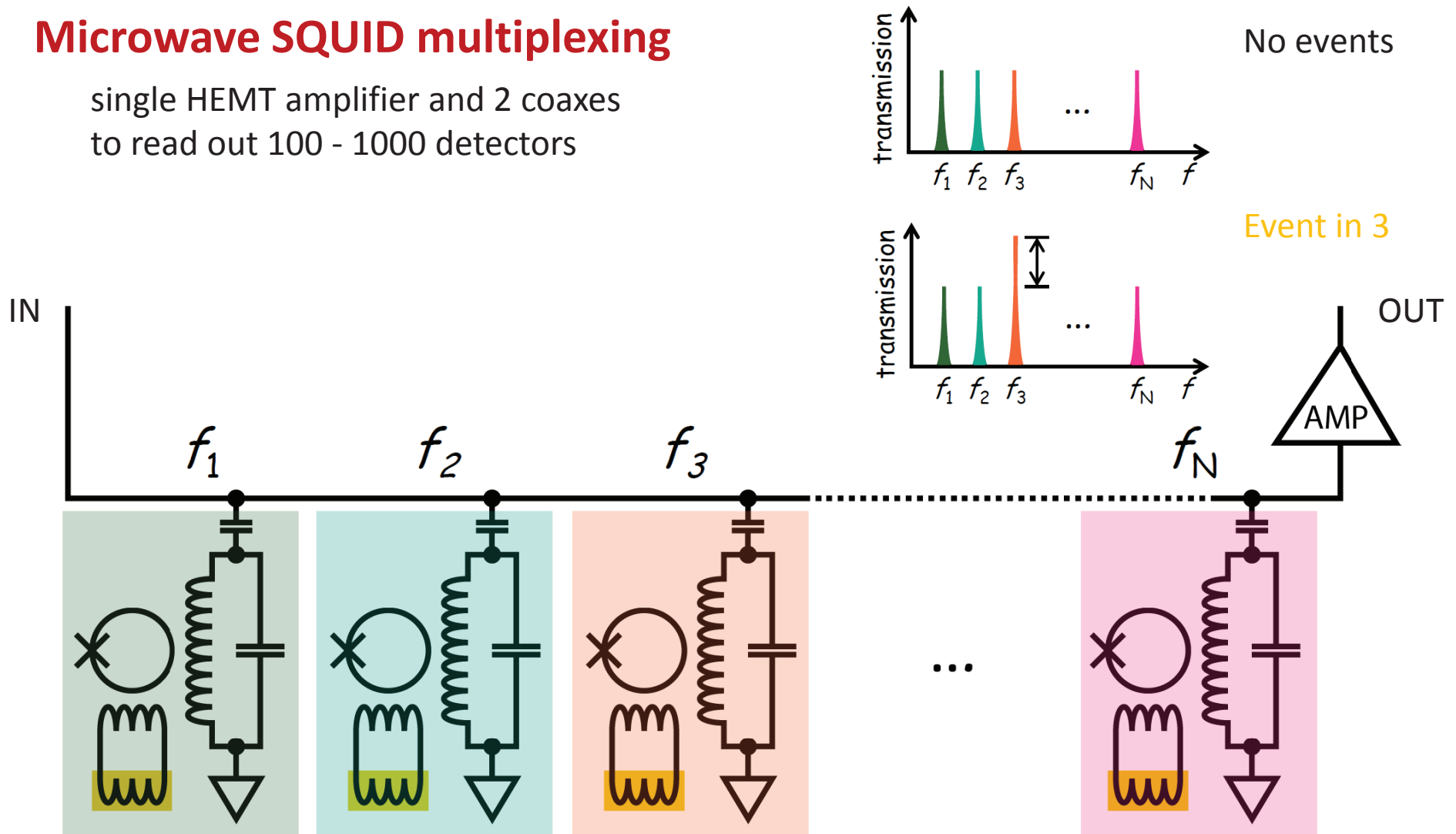


multiplexing scheme required
(TDM, FDM, CDM, ...)

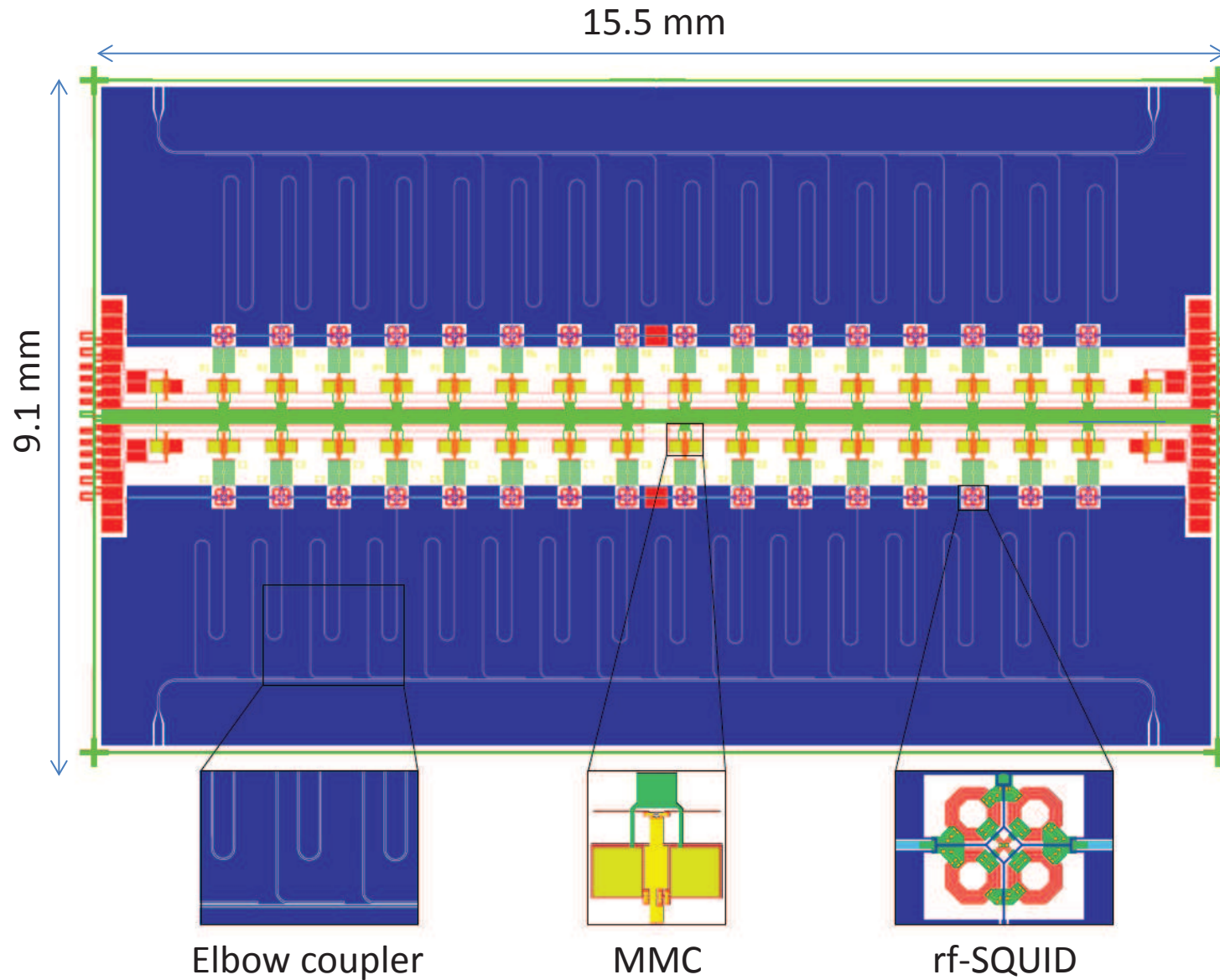
MMCs: Microwave SQUID multiplexing

Microwave SQUID multiplexing

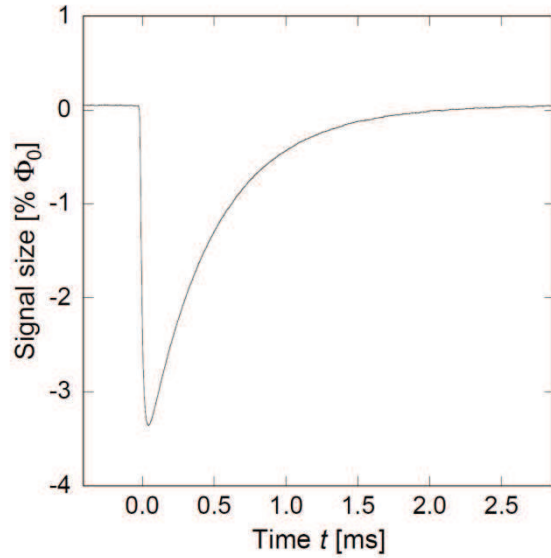
single HEMT amplifier and 2 coaxes
to read out 100 - 1000 detectors



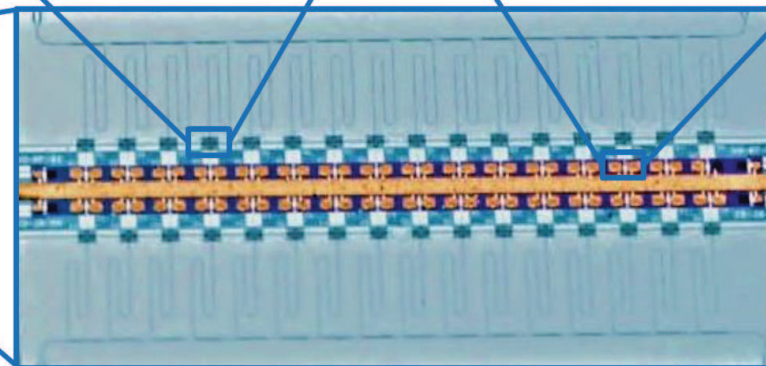
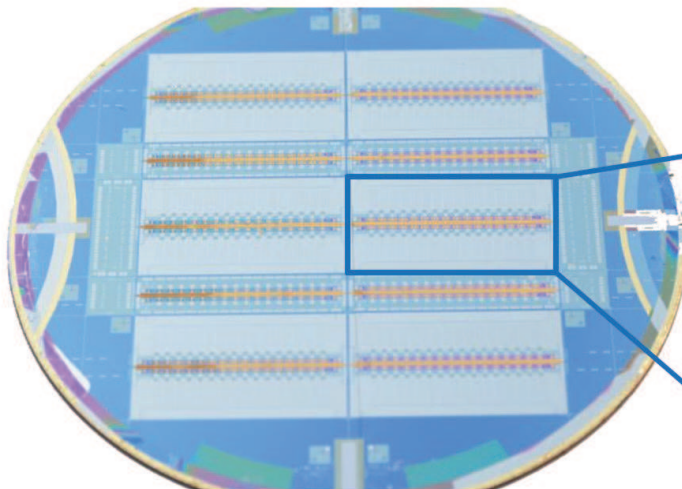
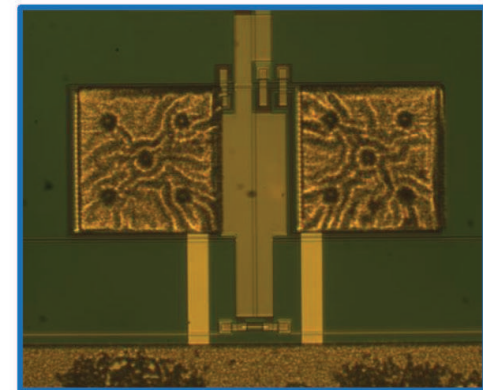
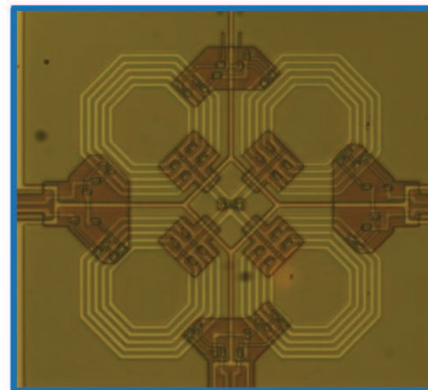
MMCs: Microwave SQUID multiplexing



MMCs: Microwave SQUID multiplexing



Successful production and test of the first prototype



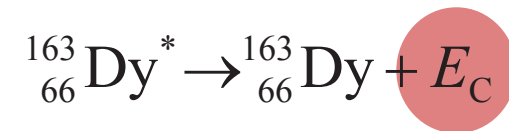
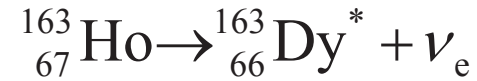
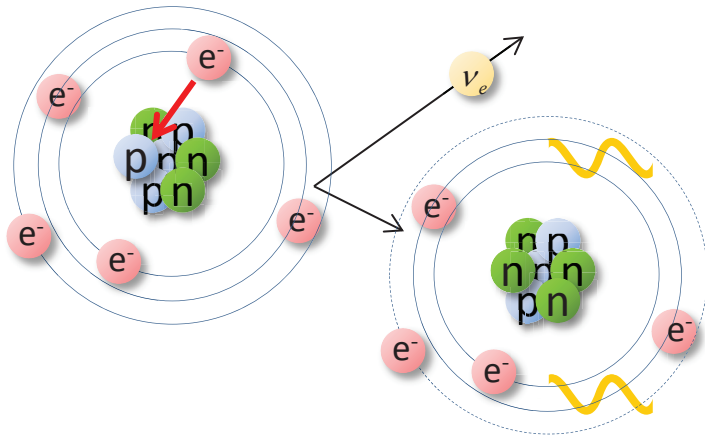
Neutrino mass determination



The Electron Capture in Ho-163 experiment

- Department of Nuclear Physics, Comenius University, Bratislava, Slovakia
- Department of Physics, Indian Institute of Technology Roorkee, India
- Institute for Nuclear Chemistry, Johannes Gutenberg University Mainz
- Institute of Nuclear Research of the Hungarian Academy of Sciences
- Institute of Nuclear and Particle Physics, TU Dresden, Germany
- Institute for Physics, Johannes Gutenberg-Universität
- Institute for Theoretical and Experimental Physics Moscow, Russia
- Institute for Theoretical Physics, University of Tübingen, Germany
- Kirchhoff-Institute for Physics, Heidelberg University, Germany
- Max-Planck Institute for Nuclear Physics Heidelberg, Germany
- Petersburg Nuclear Physics Institute, Russia
- Physics Institute, University of Tübingen, Germany
- Saha Institute of Nuclear Physics, Kolkata, India

^{163}Ho and neutrino mass



- $\tau_{1/2} \cong 4570$ years ($2 \cdot 10^{11}$ atoms for 1 Bq)
- $Q_{EC} = (2.555 \pm 0.016)$ keV *

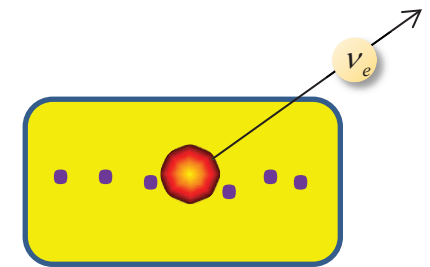
*M. Wang, G. Audi et al., *Chinese Phys. C* **36**, 1603, (2012)

A non-zero neutrino mass affects the **de-excitation energy spectrum**

Atomic de-excitation:

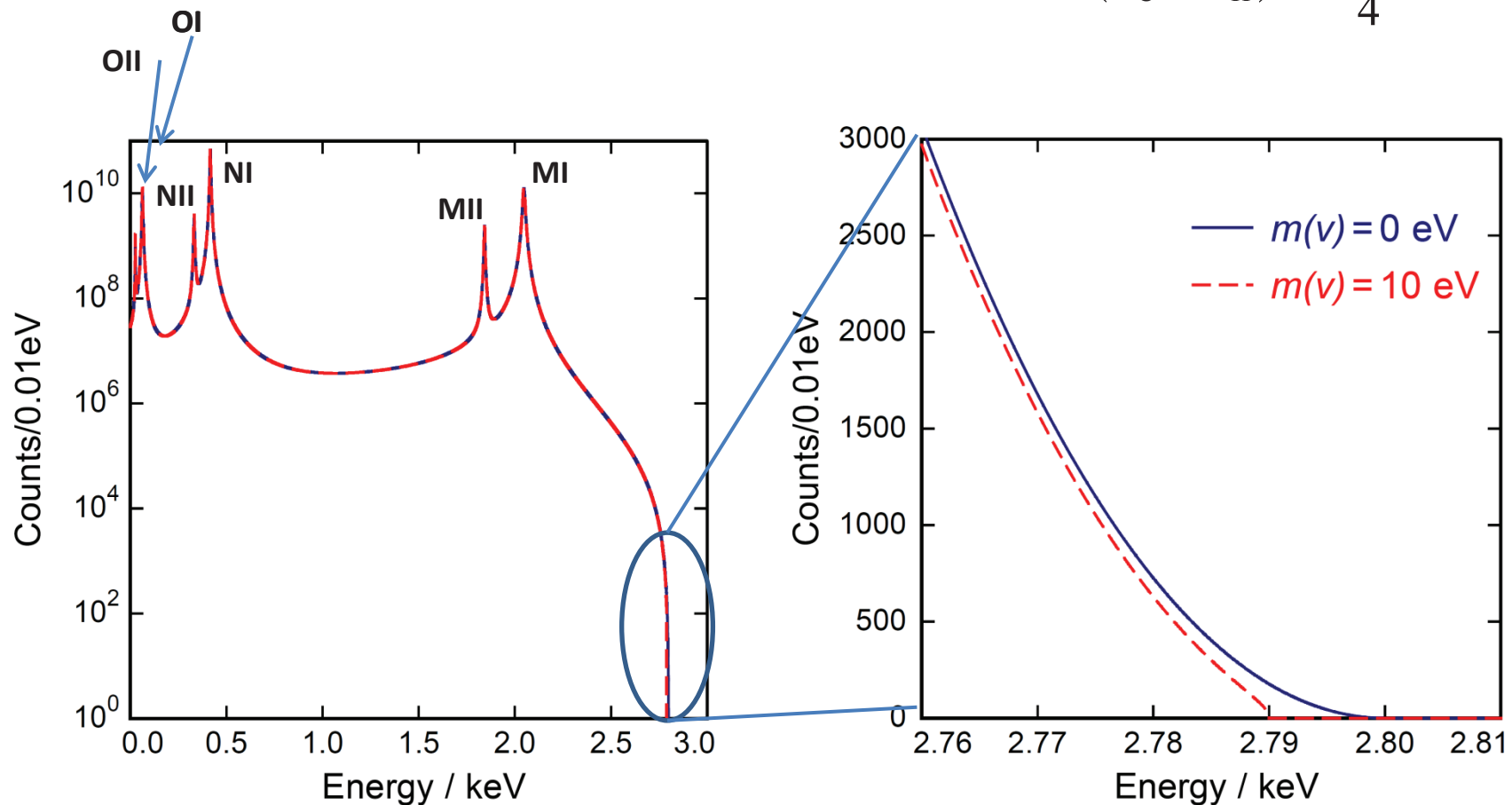
- X-ray emission
- Auger electrons
- Coster-Kronig transitions

Calorimetric measurement



^{163}Ho and neutrino mass

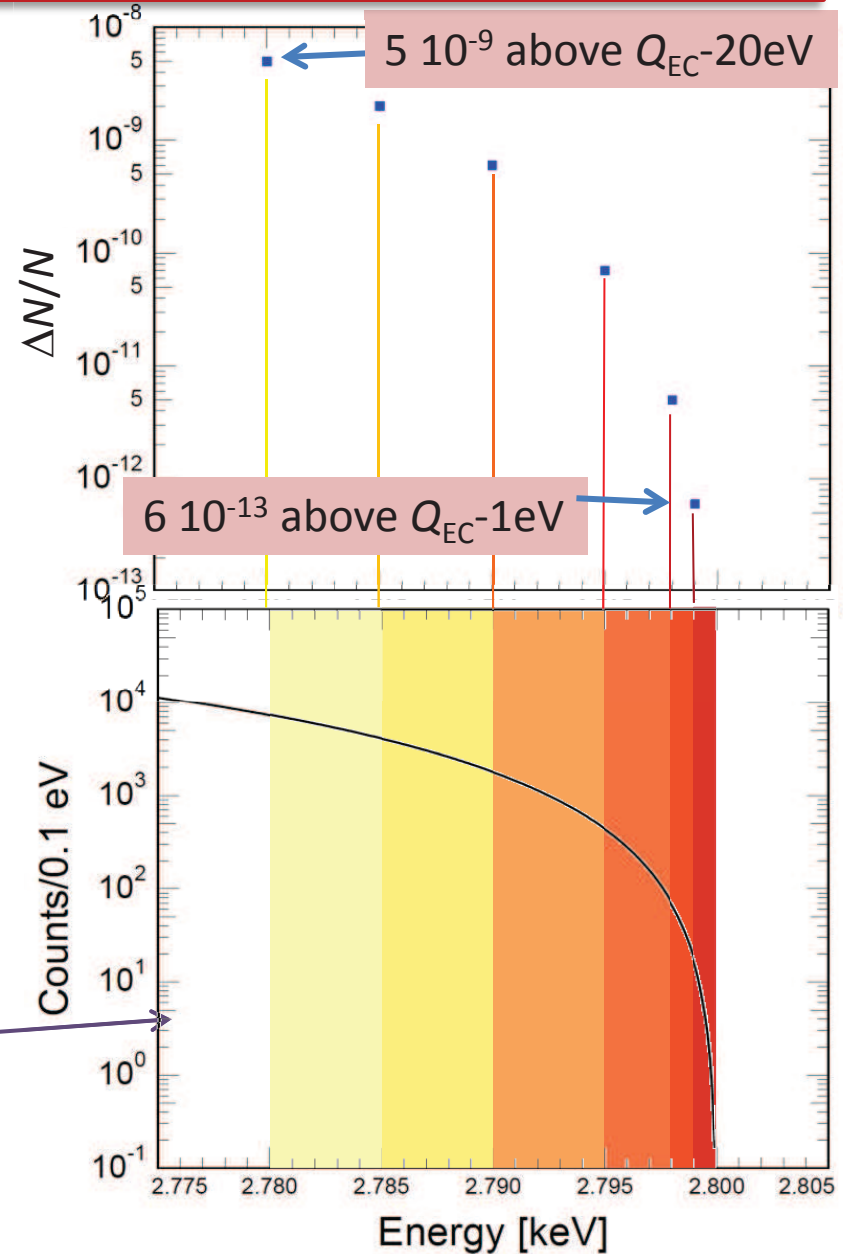
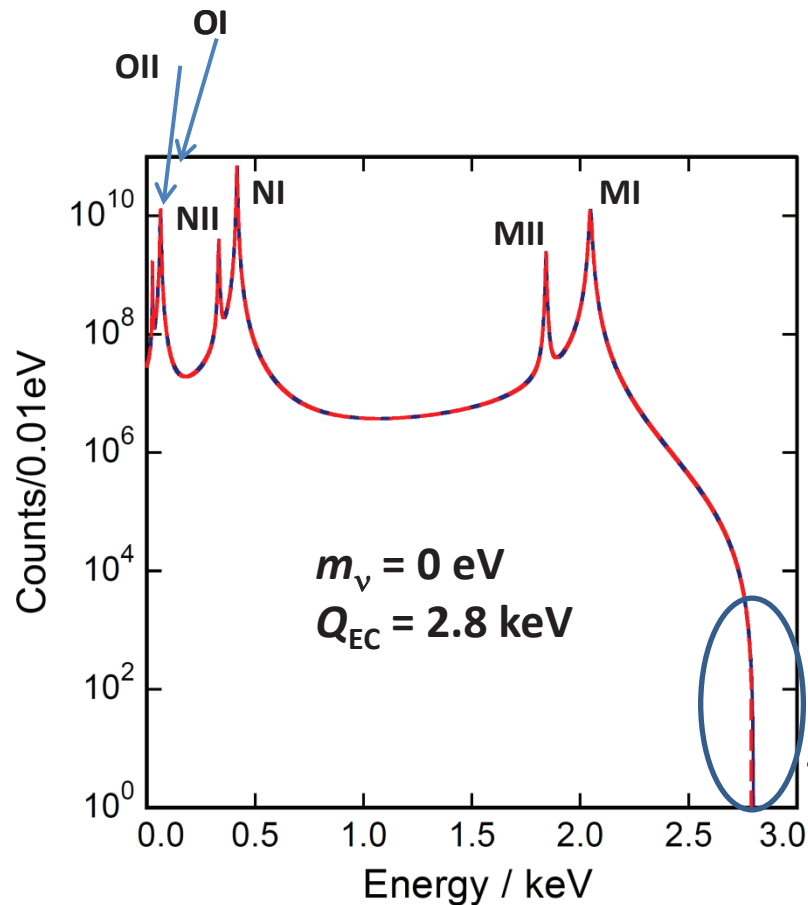
$$\frac{dW}{dE_C} = A(Q_{\text{EC}} - E_C)^2 \sqrt{1 - \frac{m_\nu^2}{(Q_{\text{EC}} - E_C)^2}} \sum_{\text{H}} B_{\text{H}} \phi_{\text{H}}^2(0) \frac{\frac{\Gamma_{\text{H}}}{2\pi}}{(E_C - E_{\text{H}})^2 + \frac{\Gamma_{\text{H}}^2}{4}}$$



^{163}Ho and neutrino mass: sub-eV sensitivity

Statistics in the end point region

- $N_{\text{ev}} > 10^{14} \rightarrow A \approx 1 \text{ MBq}$



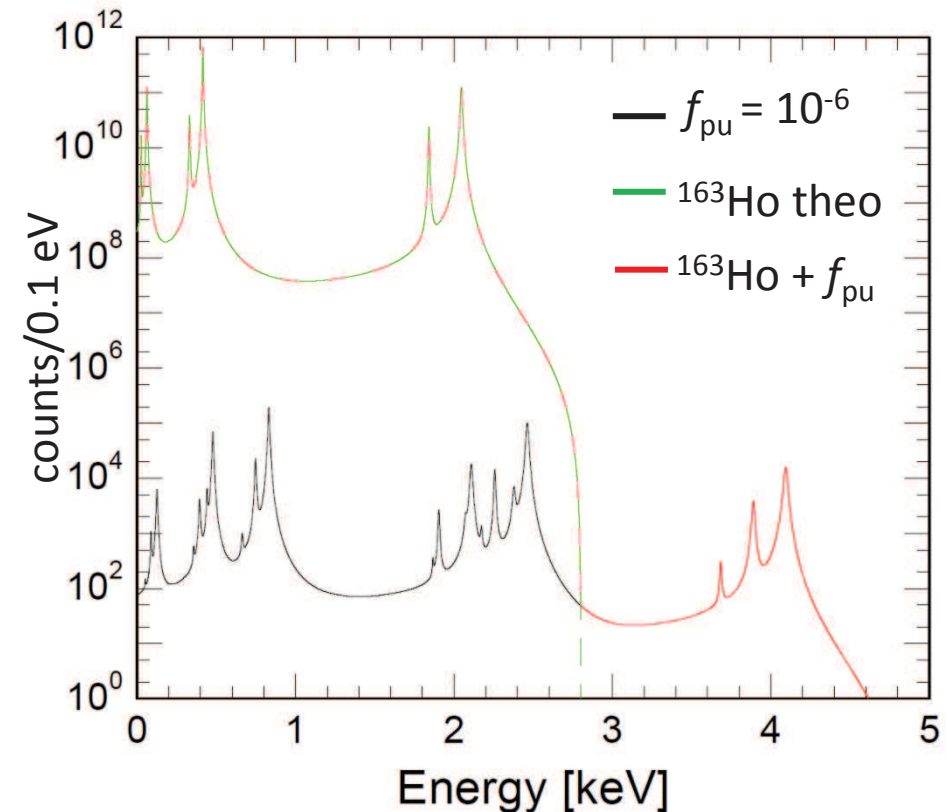
^{163}Ho and neutrino mass: sub-eV sensitivity

Statistics in the end point region

- $N_{\text{ev}} > 10^{14} \rightarrow A \approx 1 \text{ MBq}$

Unresolved pile-up ($f_{\text{pu}} \sim a \cdot t$)

- $f_{\text{pu}} < 10^{-5}$
- $\tau_r < 1 \mu\text{s} \rightarrow a < 10 \text{ Bq}$
- 10^5 pixels



^{163}Ho and neutrino mass: sub-eV sensitivity

Statistics in the end point region

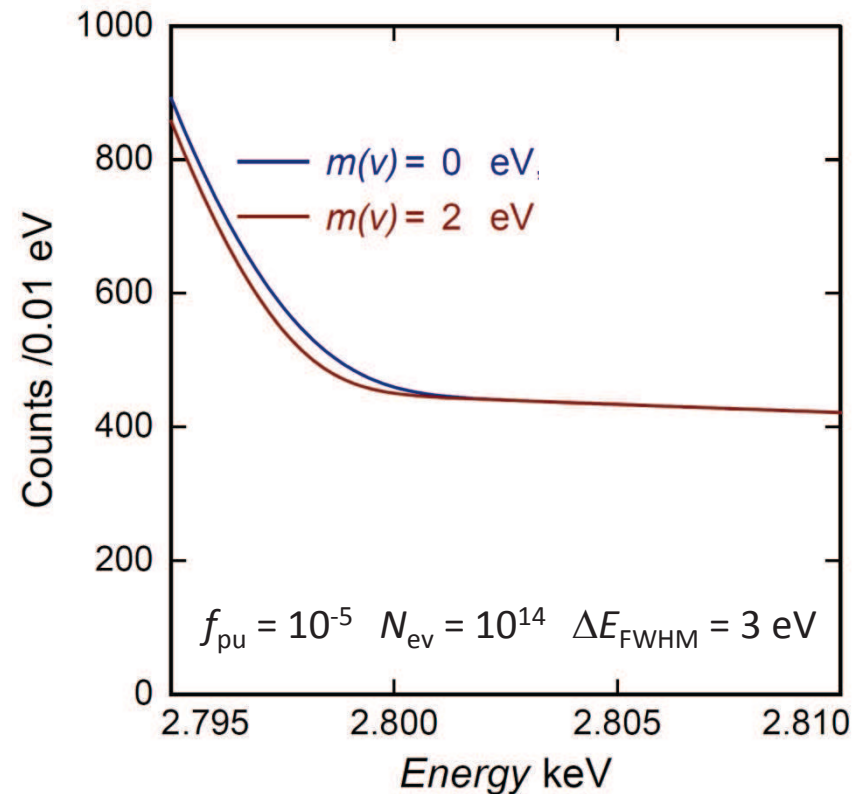
- $N_{\text{ev}} > 10^{14} \rightarrow A \approx 1 \text{ MBq}$

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Precision characterization of the endpoint region

- $\Delta E_{\text{FWHM}} < 2 \text{ eV}$



^{163}Ho and neutrino mass: sub-eV sensitivity

Statistics in the end point region

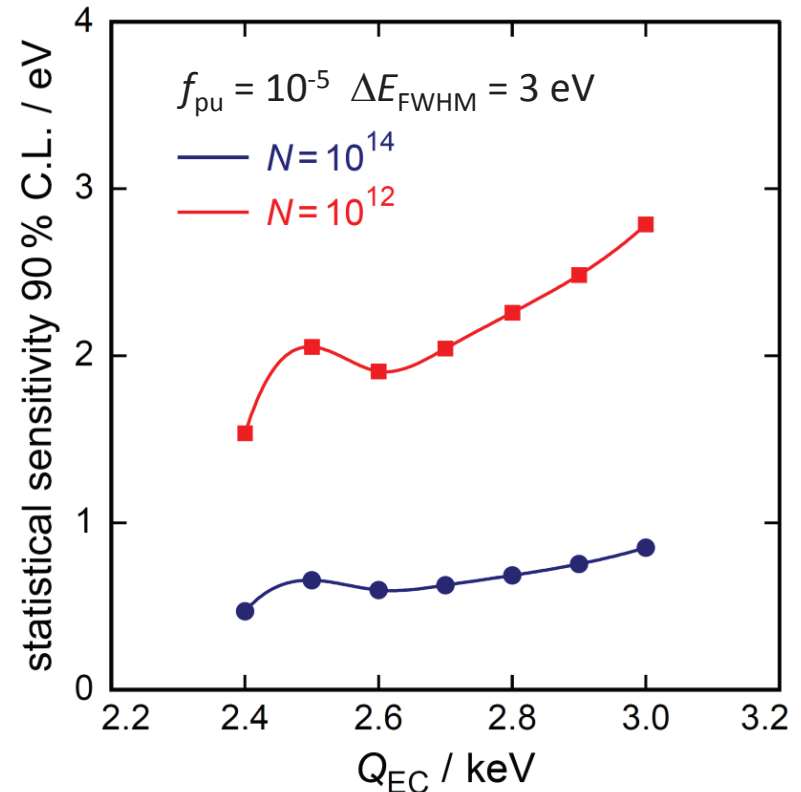
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Precision characterization of the endpoint region

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Q_{EC} determination of ^{163}Ho

Penning Trap mass spectrometry

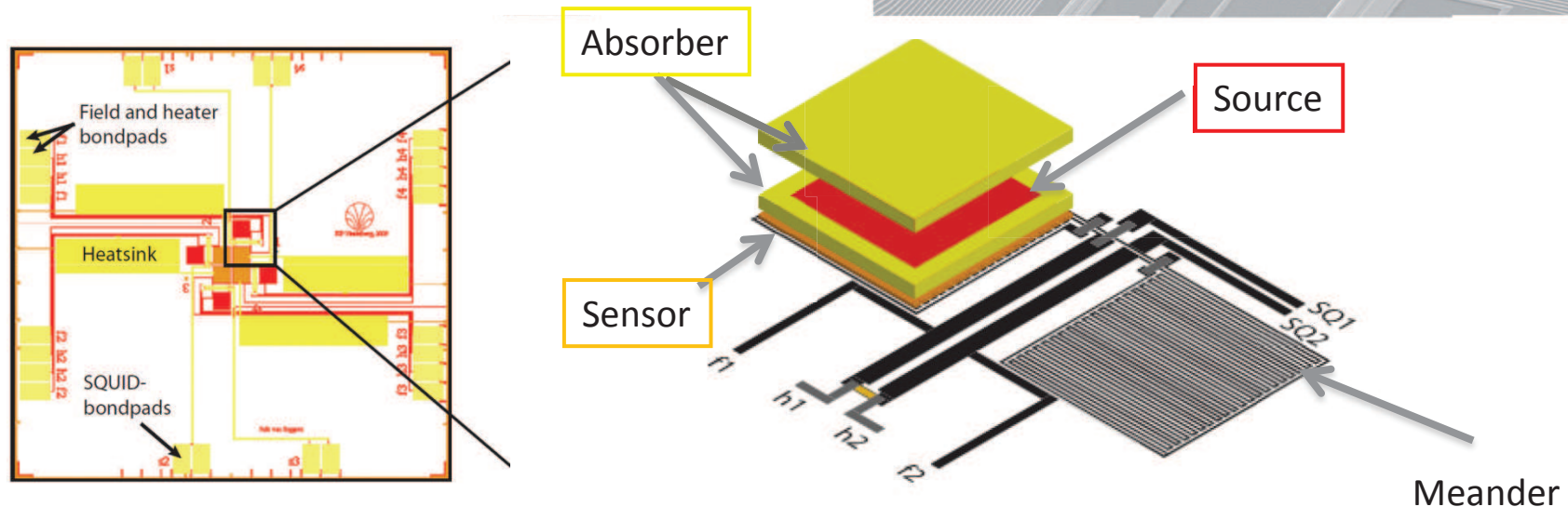
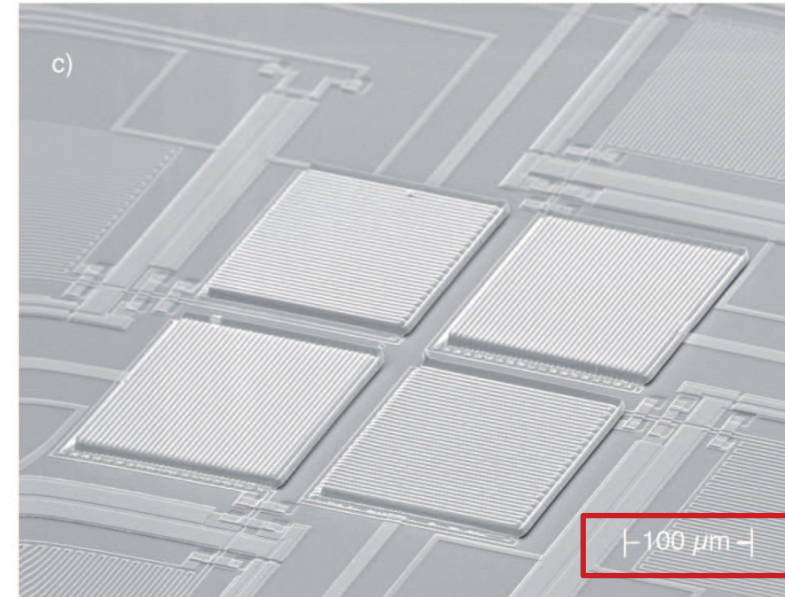
- First experiments at **TRIGATRAP (Uni-Mainz)** in 2014 *
 - Development of efficient Ho ion source using laser ablation
 - Uncertainties on ^{163}Dy and ^{163}Ho mass reduced by a factor of 2
 - Know-how to be applied in SHIPTRAP
- In a few months: **SHIPTRAP (GSI)**
 - Q_{EC} determination within **30 - 100 eV**
- In a few years: **PENTATRAP (MPI-K HD)**
 - Q_{EC} determination within **1 eV**

*Preparatory studies for a high-precision Penning trap measurement of the ^{163}Ho electron capture Q-value
F. Schneider et al., submitted to EPJ



First detector prototype

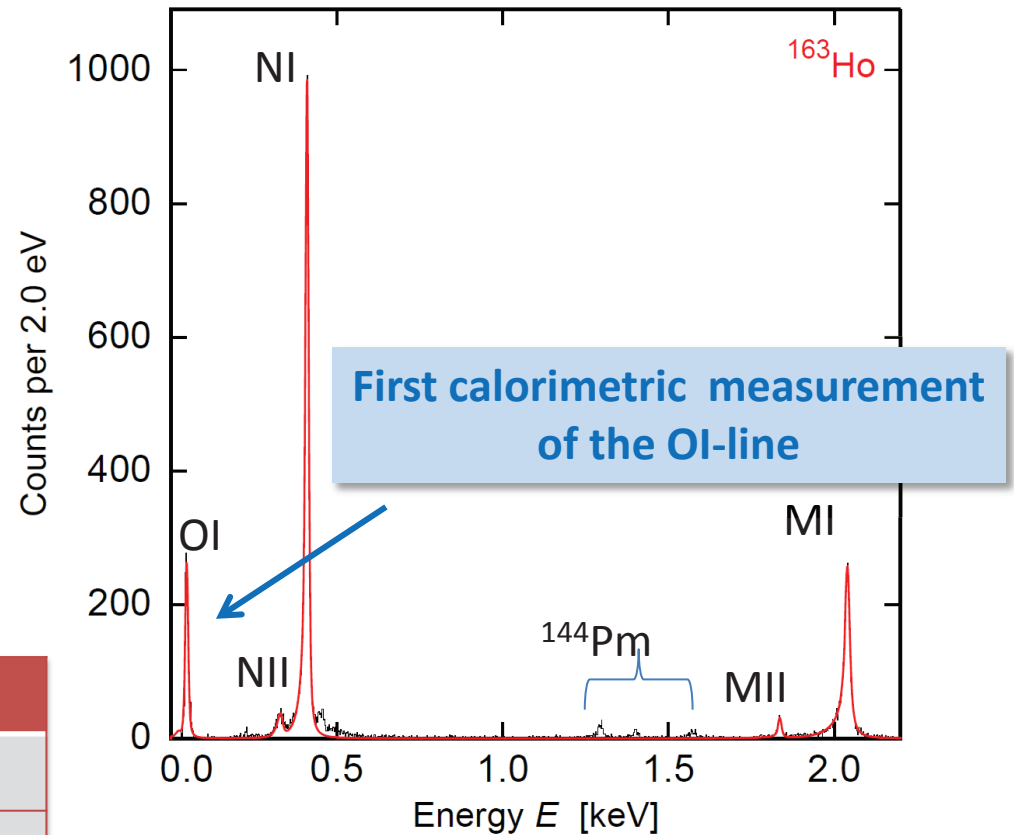
- Absorber for calorimetric measurement
→ ion implantation @ ISOLDE-CERN
- About 0.01 Bq per pixel
- Two pixels have been simultaneously measured



Calorimetric spectrum

- Rise Time ~ 130 ns
- $\Delta E_{\text{FWHM}} = 7.6$ eV @ 6 keV (2013)
 $\Delta E_{\text{FWHM}} = 2.4$ eV @ 0 keV (2014)
- Non-Linearity $< 1\%$ @ 6keV
- Synchronized measurement of 2 pixels
- Presently most precise ^{163}Ho spectrum

	E_{H} bind.	E_{H} exp.	Γ_{H} lit.	Γ_{H} exp
MI	2.047	2.040	13.2	13.7
MII	1.845	1.836	6.0	7.2
NI	0.420	0.411	5.4	5.3
NII	0.340	0.333	5.3	8.0
OI	0.050	0.048	5.0	4.3



$$Q_{\text{EC}} = (2.843 \pm 0.009^{\text{stat}} - 0.06^{\text{syst}}) \text{ keV}$$

Where to improve

Detector design and fabrication:

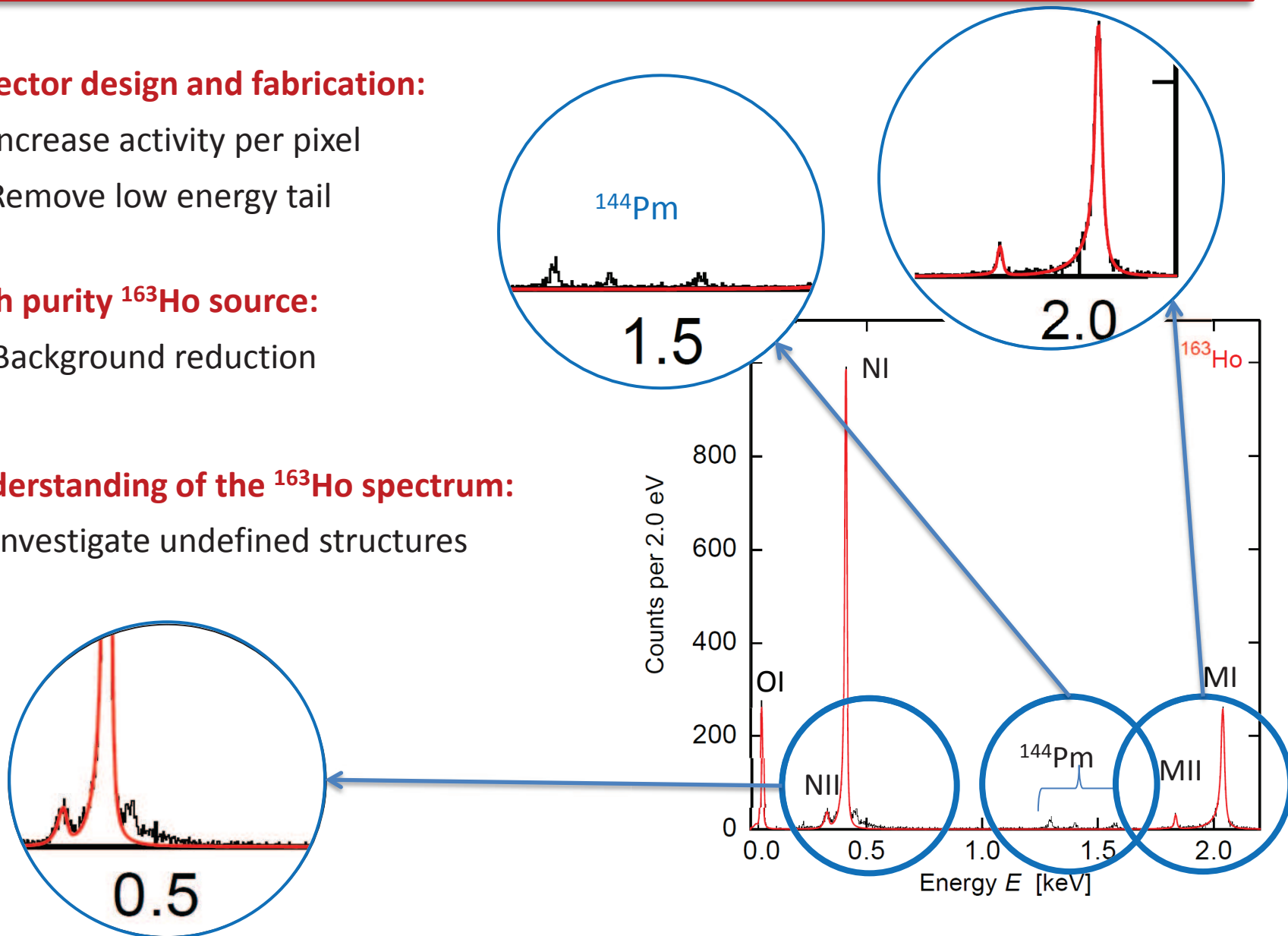
- Increase activity per pixel
- Remove low energy tail

High purity ^{163}Ho source:

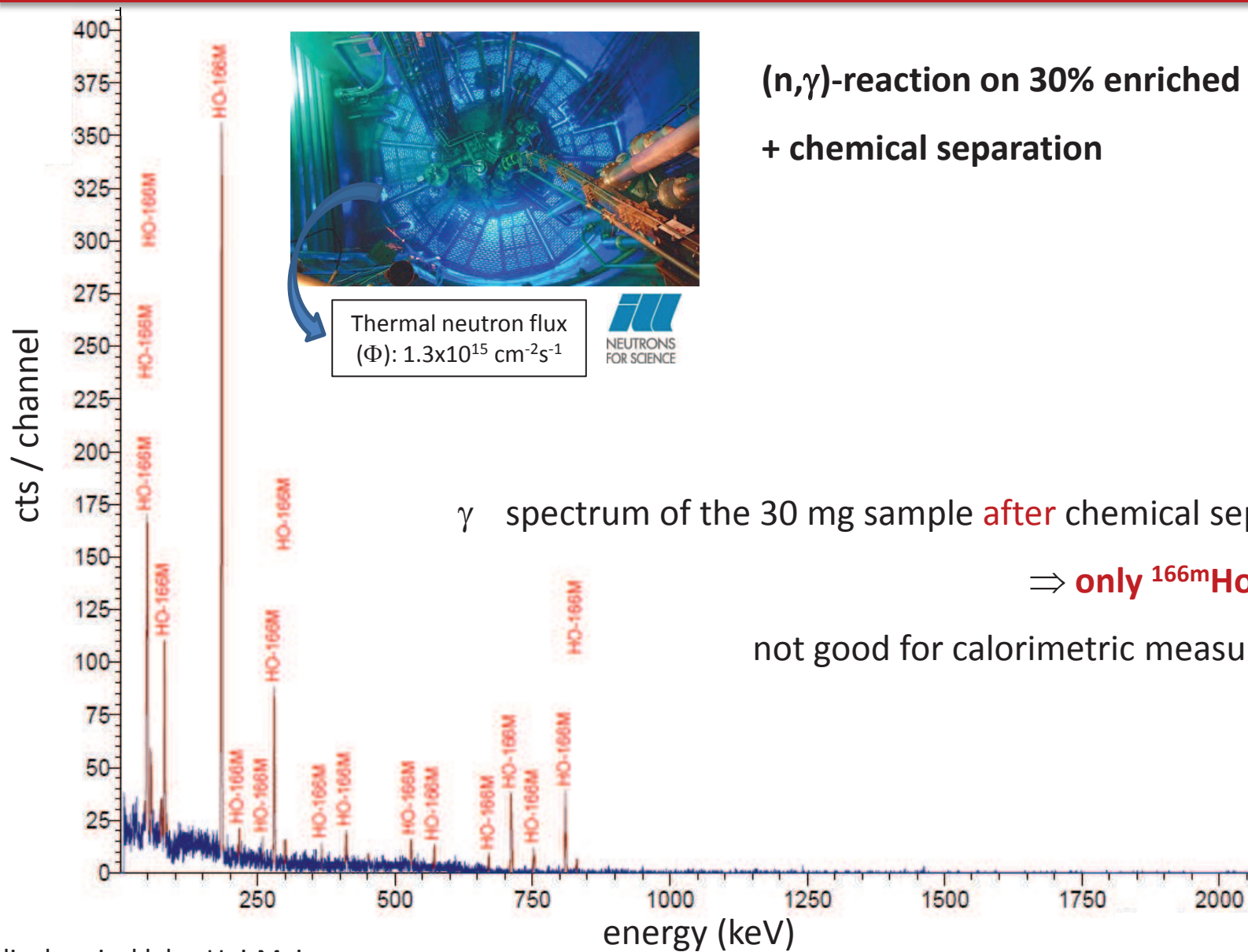
- Background reduction

Understanding of the ^{163}Ho spectrum:

- Investigate undefined structures



High purity ^{163}Ho source: Chemical separation



(n, γ)-reaction on 30% enriched ^{162}Er
+ chemical separation

γ spectrum of the 30 mg sample **after** chemical separation:

\Rightarrow **only $^{166\text{m}}\text{Ho}$ visible**

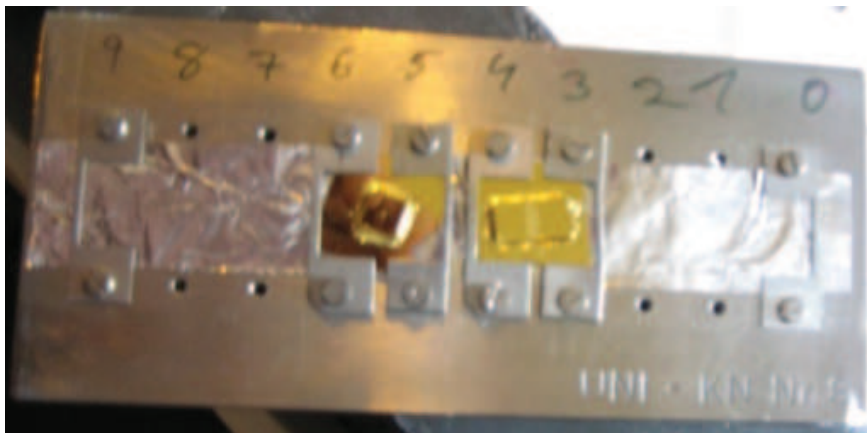
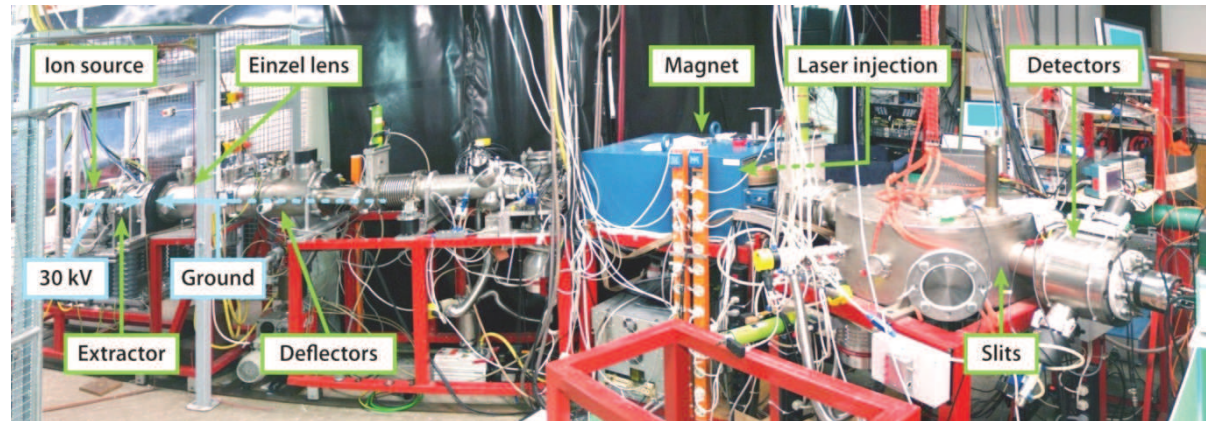
not good for calorimetric measurements!

High purity ^{163}Ho source: Mass separation

Goal: $^{166\text{m}}\text{Ho}/^{163}\text{Ho} \leq 10^{-9}$

- **RISIKO** mass-separator
@ Uni-Mainz

✓ First successful test
with natural Ho



- Separation at **CERN/ISOLDE** December 2014
 - ✓ 2 new chips to test!
... each with 16 pixel detector arrays
for soft x-rays
 - higher activity per pixel ≈ 1 Bq
 - no radioactive contaminants

Background

Background sources:

- Environmental radioactivity
- Cosmic rays
- Induced secondary radiation by cosmic rays
- Radioactivity in the detector

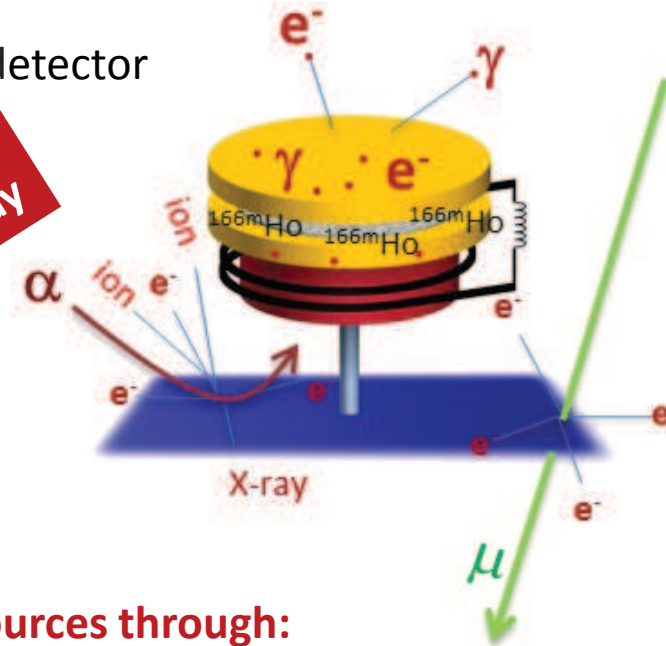
→ Material screening

→ [Underground labs
μ-Veto]



Underground measurements in [Modane](#)

Background level
 5×10^{-5} counts/eV/det/day



Study of background sources through:

- Monte Carlo simulations
- Dedicated experiments

Screening facilities

- Uni-Tübingen
- Felsenkeller



ECHo overview

- Prove **scalability** with medium large experiment **ECHo-1K**
 - $A \sim 1000$ Bq High purity ^{163}Ho source (produced at reactor)
 - $\Delta E_{\text{FWHM}} < 5$ eV
 - $\tau_r < 1$ μs
 - multiplexed arrays \rightarrow microwave SQUID multiplexing
- 1 year measuring time $\rightarrow 10^{10}$ counts = Neutrino mass sensitivity $m_\nu < 10$ eV

Just approved

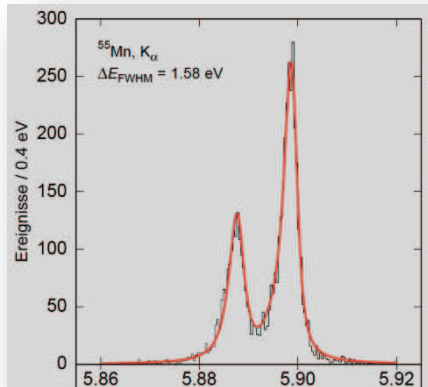
Research Unit FOR 2202/1

„Neutrino Mass Determination by Electron Capture in Holmium-163 – ECHo“

DFG Deutsche
Forschungsgemeinschaft

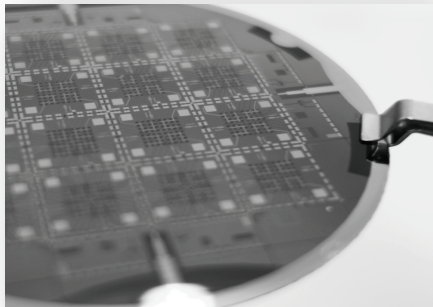
- **ECHo-1M** towards sub-eV sensitivity

Summary and Outlook



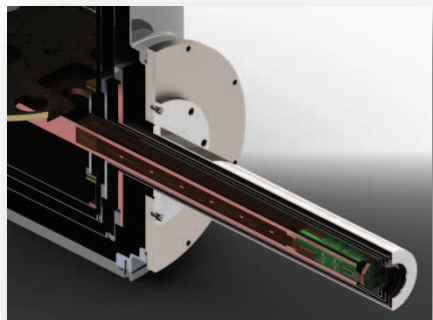
metallic magnetic calorimeters

- are versatile low temperature detectors
- high resolution for all kinds of particles
- wide range of energies
- impressive resolving powers in reach



micro-fabrication works

- first detector arrays fabricated
- designed performance is reached
- 2d arrays are on the way



multiplexing

- demonstrated principles

Low temperature physics @KIP



Thank you!