

The Scintillating Fibre Tracker for the LHCb Upgrade

DESY Joint Instrumentation Seminar

Presented by

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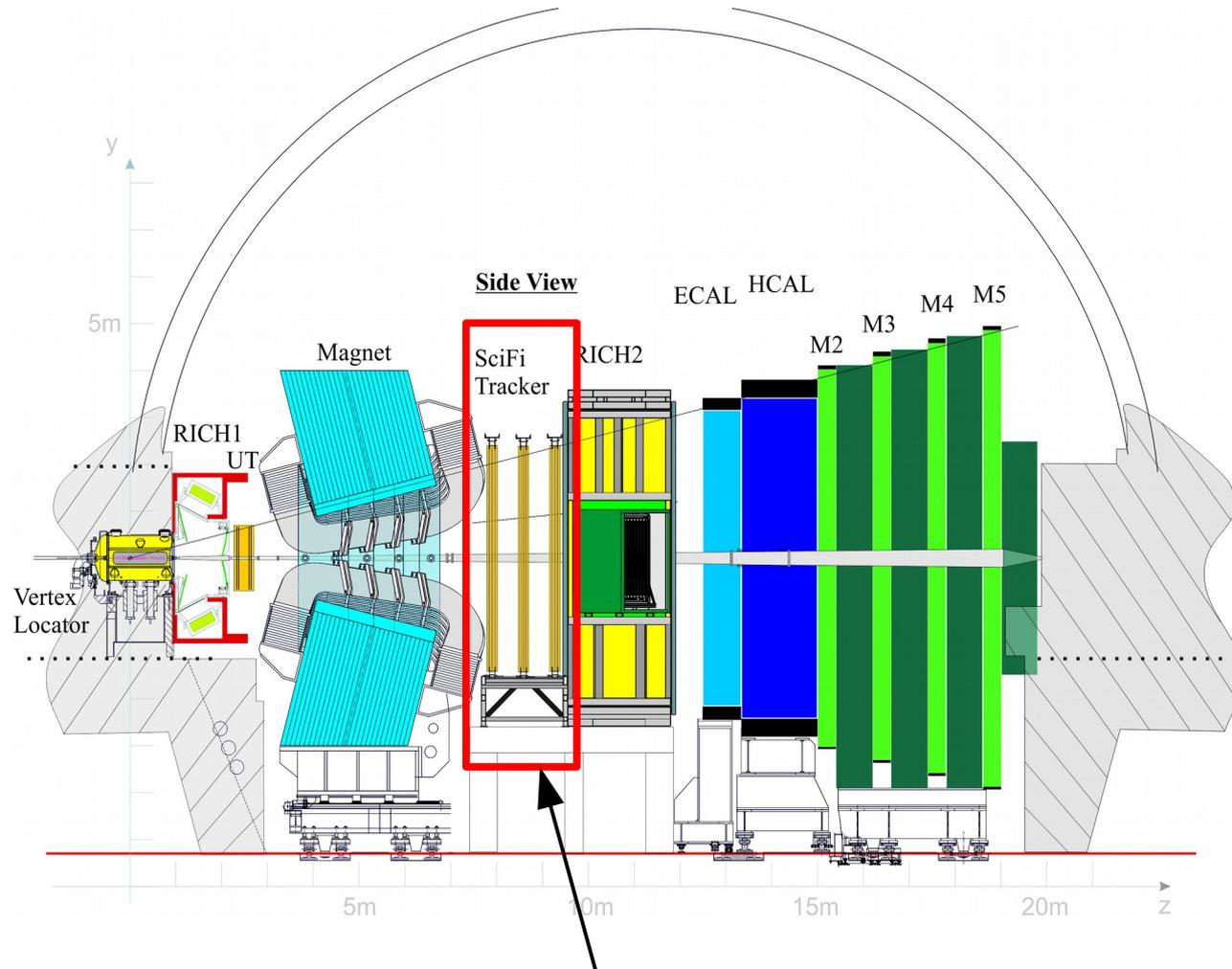
UNIVERSITÄT
HEIDELBERG
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SEIT 1386

on behalf of the LHCb SciFi Tracker group



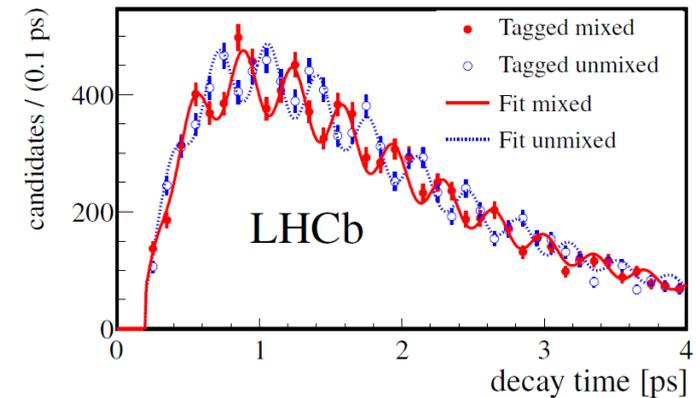
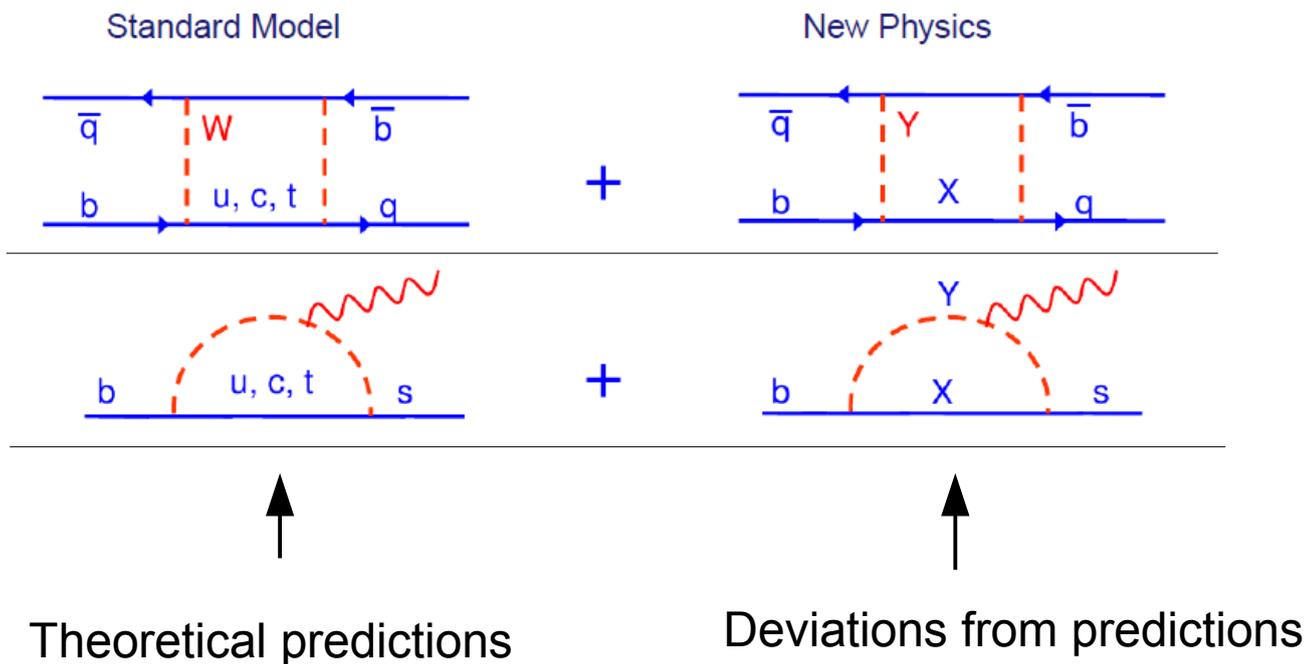
Outline

- LHCb and the Upgrade overview
- The SciFi Tracker
 - Detector basics
 - Challenges

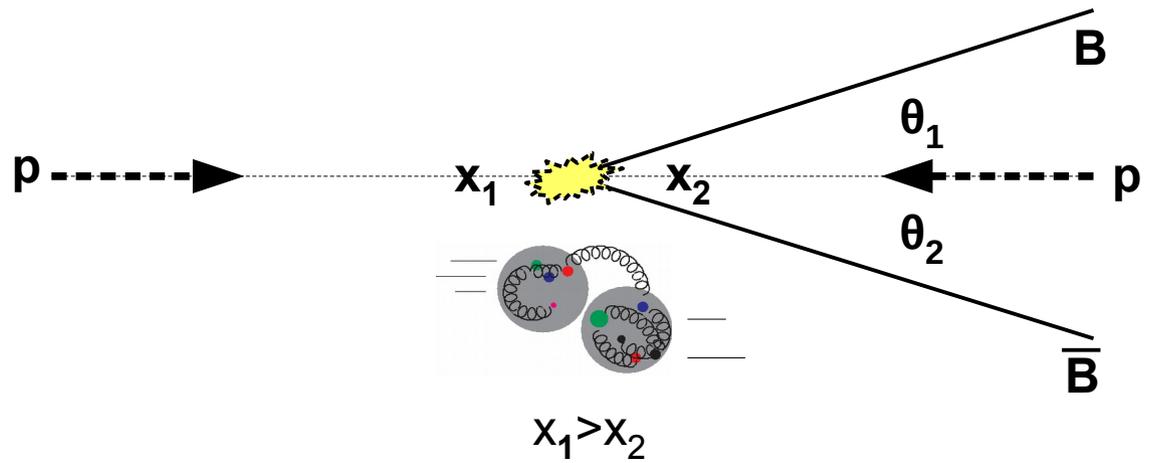
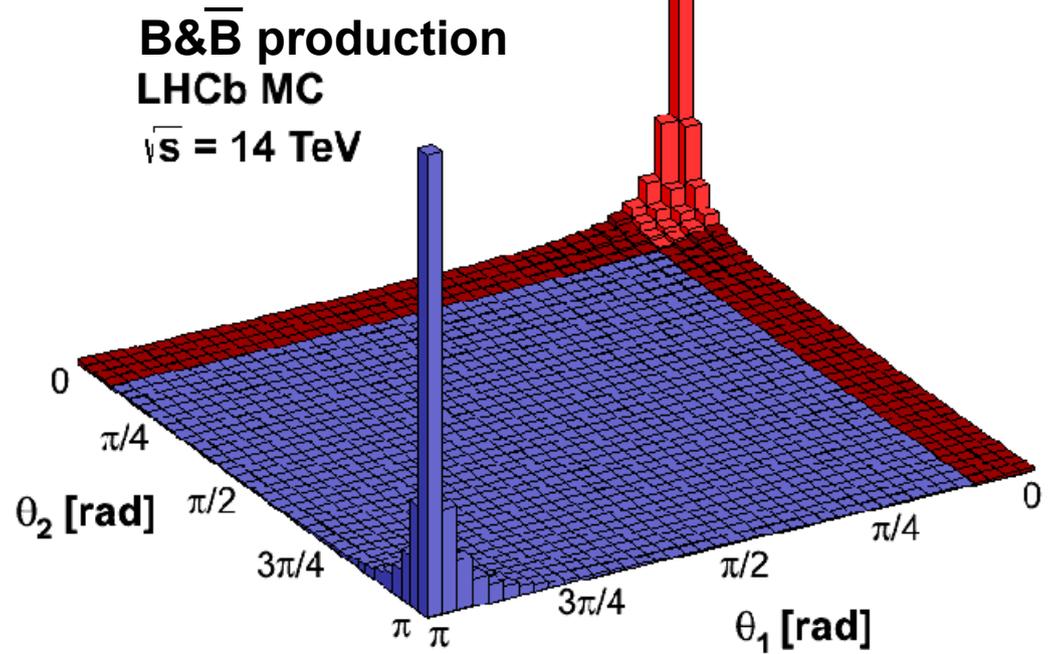
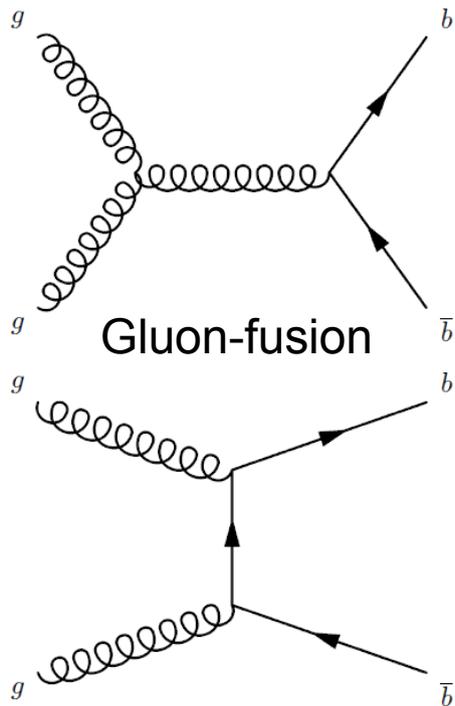


The LHCb detector

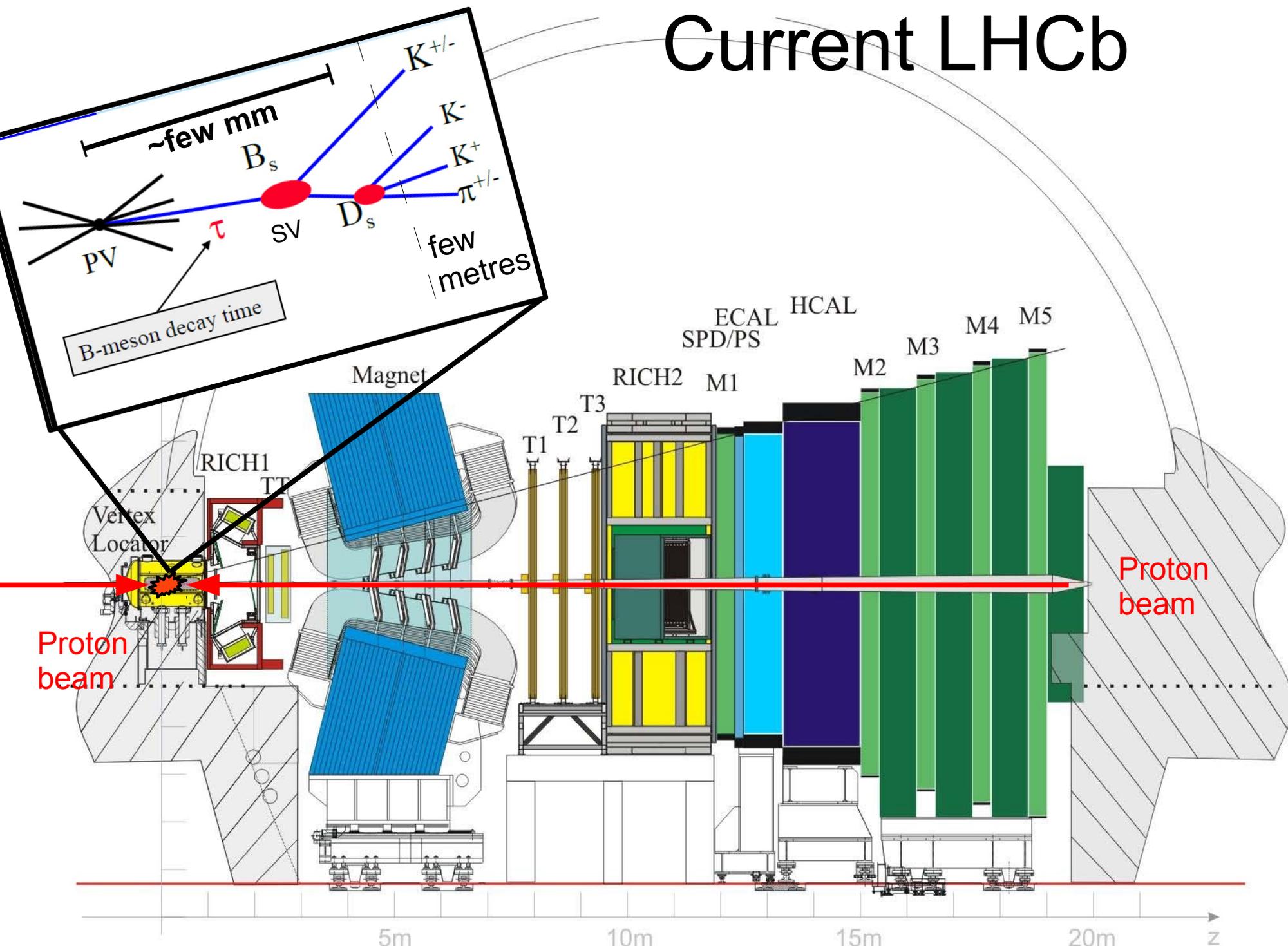
- Built for indirect searches for new physics via precision measurements of quantum loop induced processes in the b- and c-quark systems
 - Rare decays
 - Particle/anti-particle asymmetry



b-mesons are produced in the forward direction at the LHC

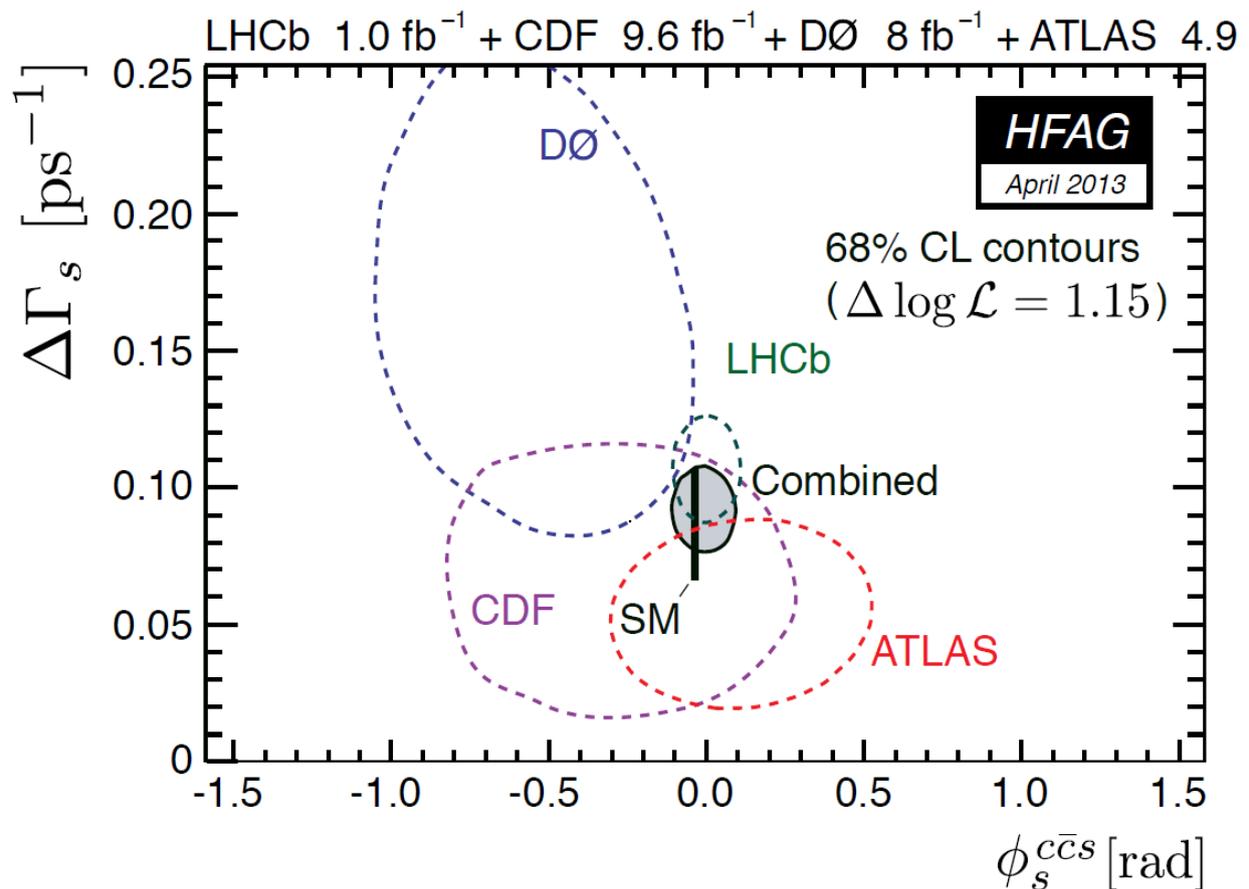


Current LHCb



- LHCb is running at twice its design value ($\sim 2 \times 10^{12}$ $b\bar{b}$ /year), 180+ papers published

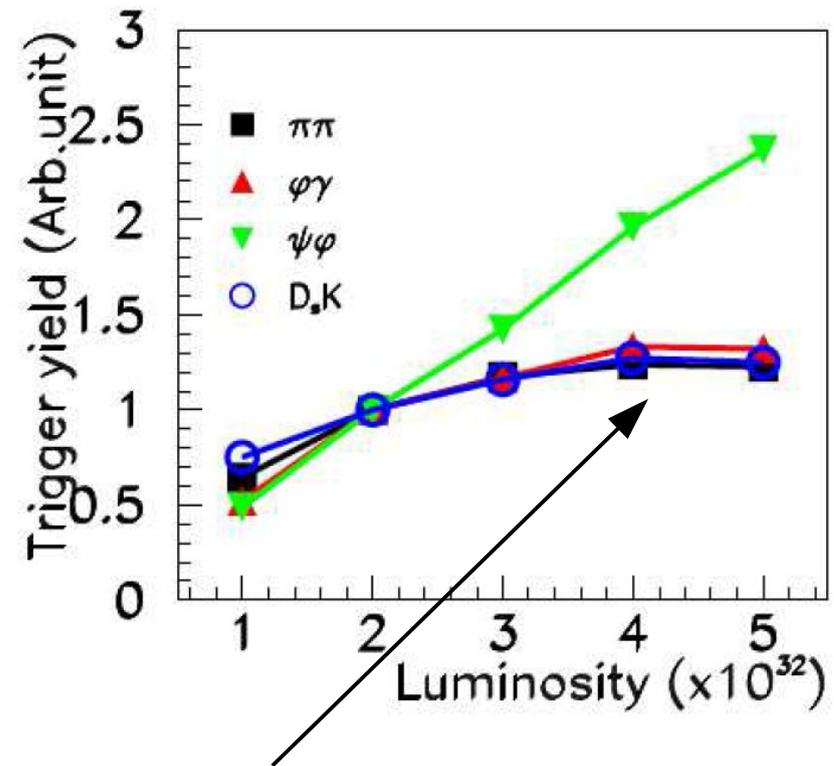
- Almost every physics measurement in LHCb is limited by statistical uncertainties, not systematic



We need more data!!

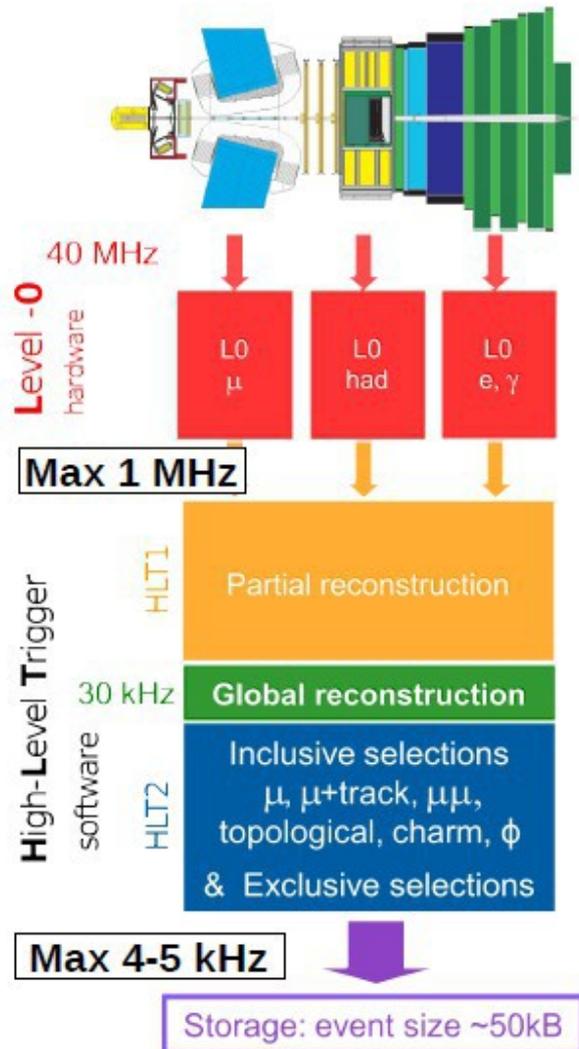
Limitations

- LHCb collision rate is tuned to manage data rate (can be increased), but...
- Statistics are **limited by the 1MHz hardware trigger rate** and then **detector occupancy**

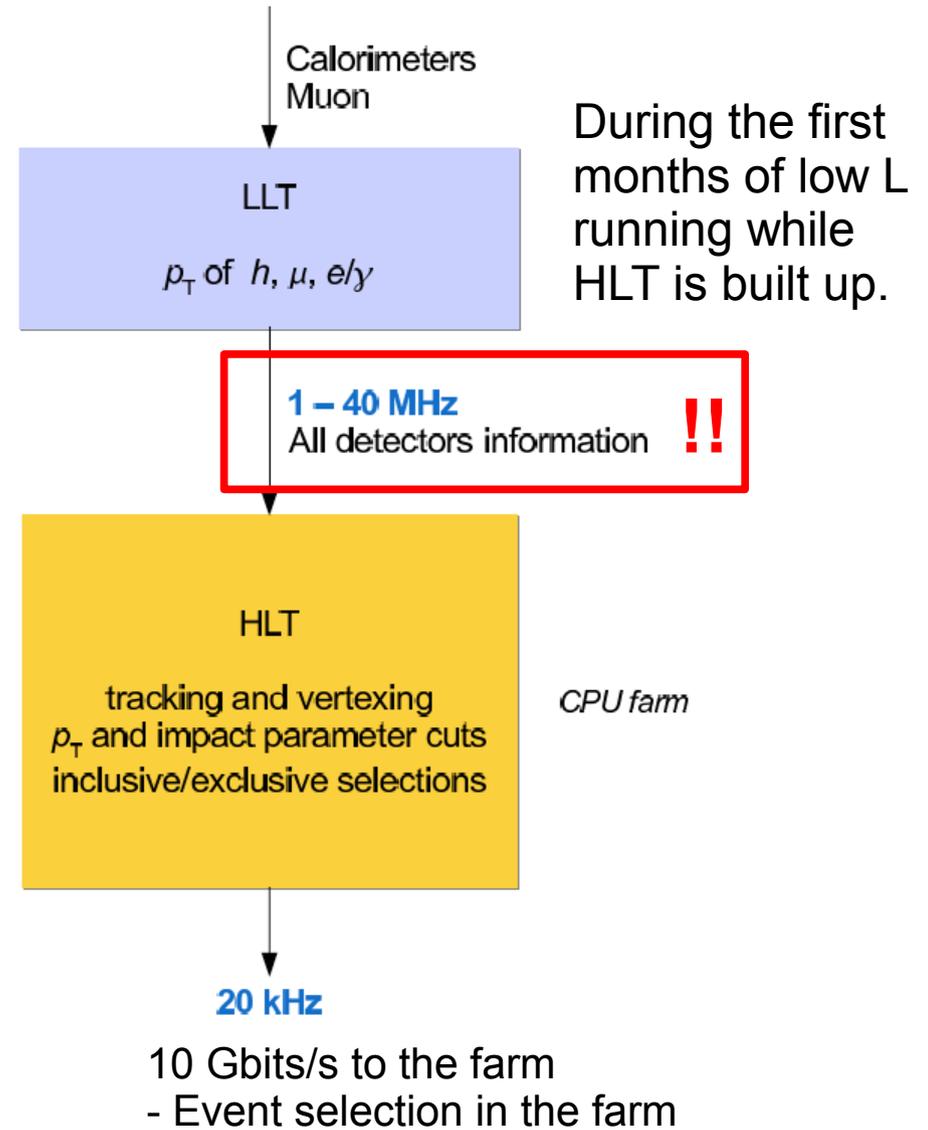


Saturation of hadronic modes with L0-hardware trigger

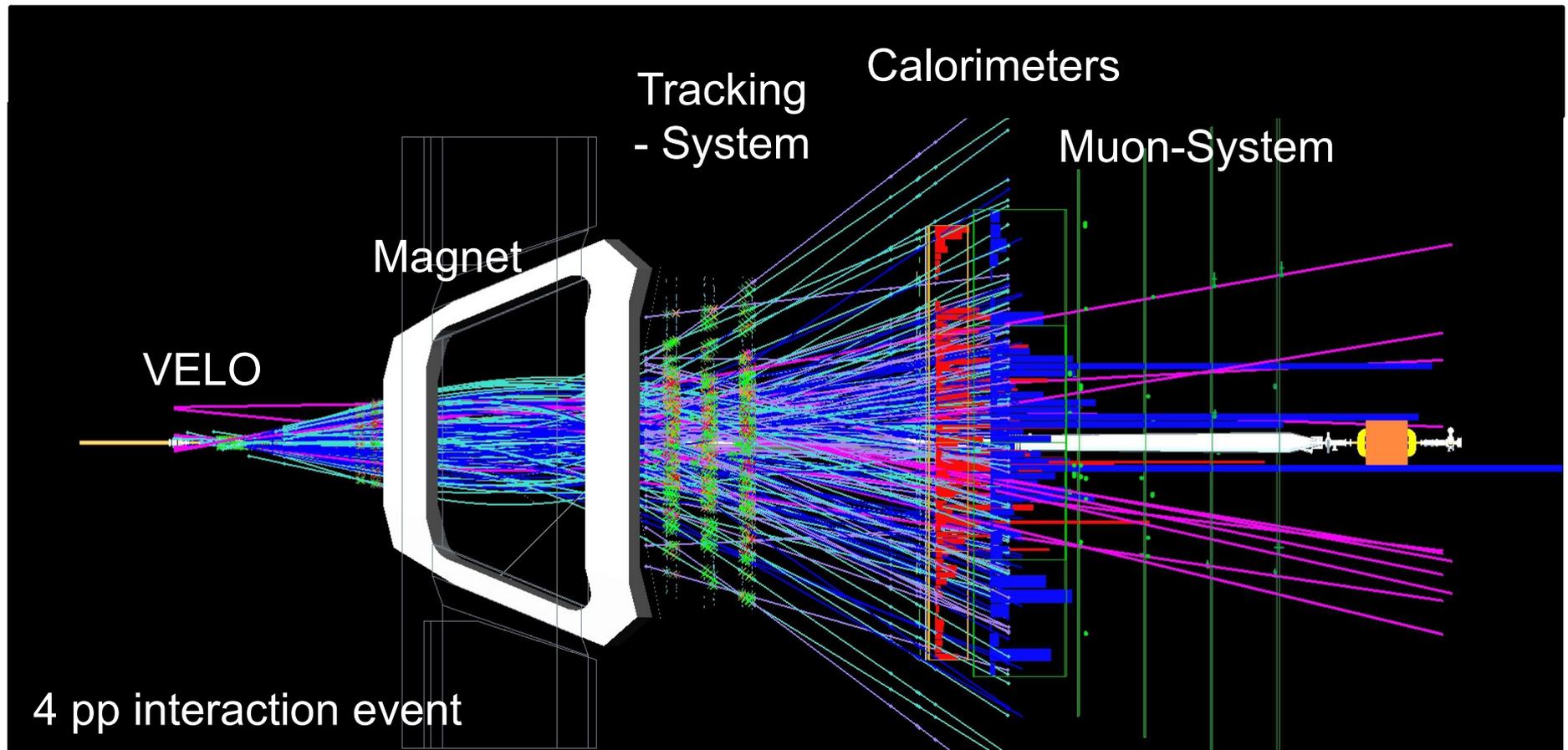
Current Trigger Scheme



Upgrade Trigger Scheme



Detector Occupancy and Efficiency



Current visible pp interactions/event:

Poisson distribution with $\mu \approx 2$;
Upgrade is at $\mu \approx 5$

72 tracks, on average for a B-Bbar event;
180 in upgrade

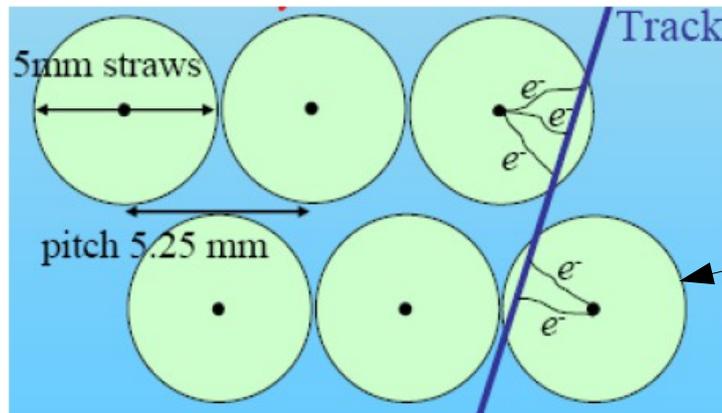
→ We need a high hit detection efficiency (98+%)

Detector Occupancy and Efficiency

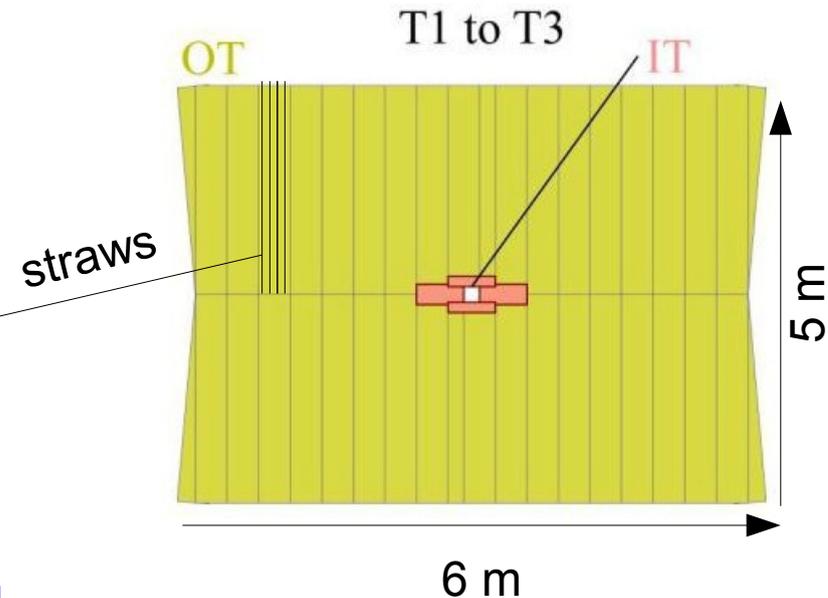
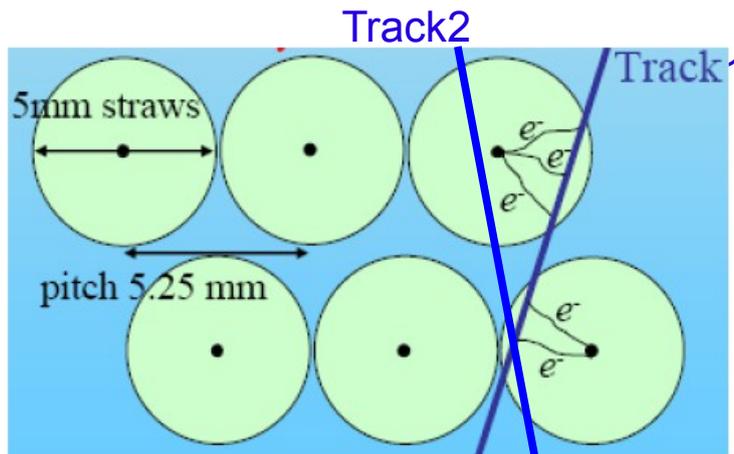
Outer Tracker = 5 mm straw gas drift tubes (2.5m long)

- Detector is insensitive to multiple tracks per tube (35ns drift time)

Good!!



BAD!!

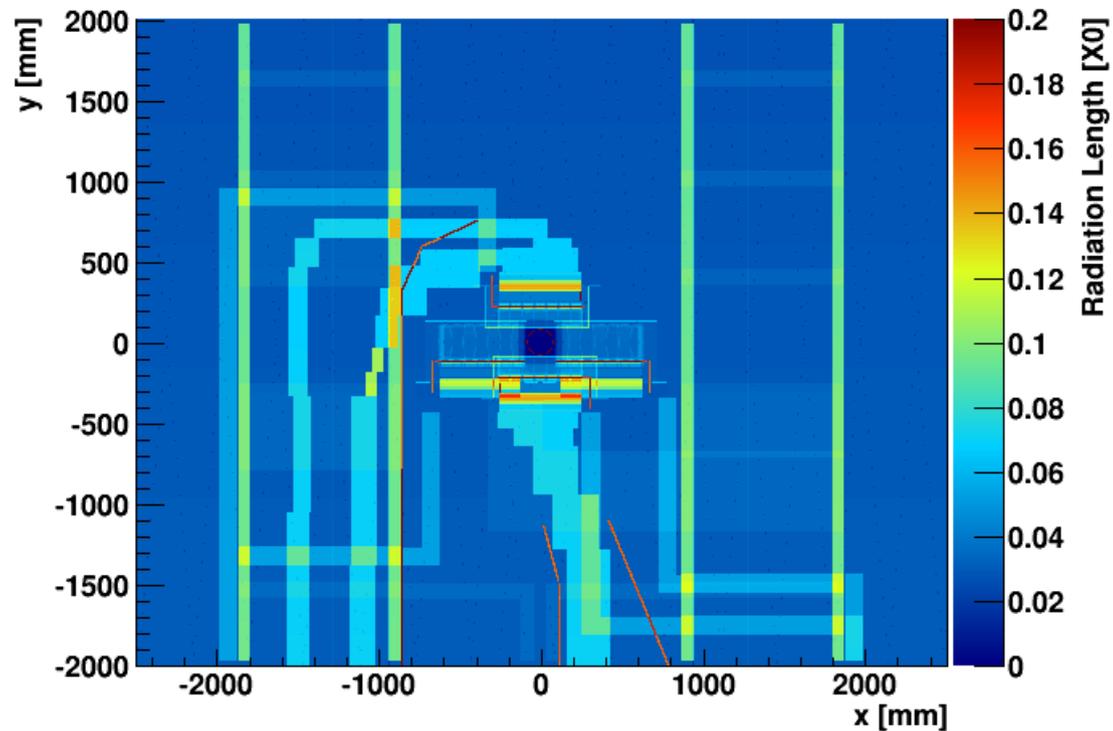


Outer Tracker tracking efficiency decreases above 25% occupancy → 40% expected in the upgrade

Beam bunch spacing will be 25ns in 2015+

Material Budget

- Material in T1



Particles averaged over eta and phi see 17.5% of X0 over 3 stations

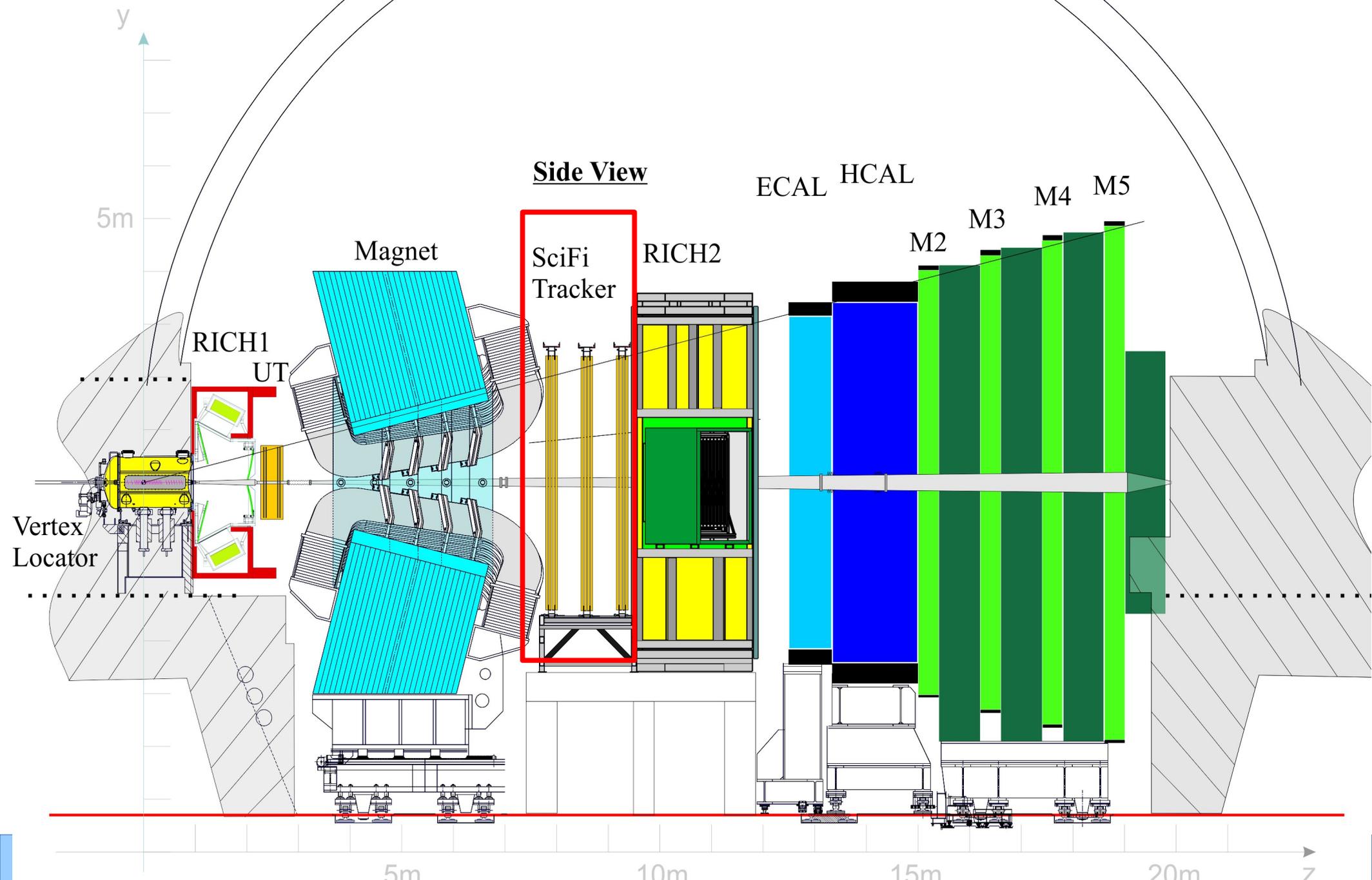
The other Upgrade detectors

- **VeLo** → **VeLoPix** @ 40MHz
 - 55 x 55 μm^2 pixels; a full 3D pattern recognition in HLT.
- **TT** → **UT** (tracker) @ 40MHz
 - silicon strip detectors (X-U-V-X) at 5° ; Improved small angle acceptance, less material ($<4.5\% X_0$).
- **RICH/PID** Upgrade
 - integrated HPD+FEE(pixel @ 1 MHz) → MA-PMT @ 40MHz
 - Remove the aerogel from RICH-1; only C4F10; RICH-2 stays as CF4.
- **Calorimeter** electronics → 40 MHz
 - Scintillating Pad Detector (SPD) and the Preshower (PRS), lead absorber will be removed. Lose some e/gamma PID power
- **Muon Stations**: Front-end is already 40MHz to L0 trigger; switch to LLT;
 - Remove M1

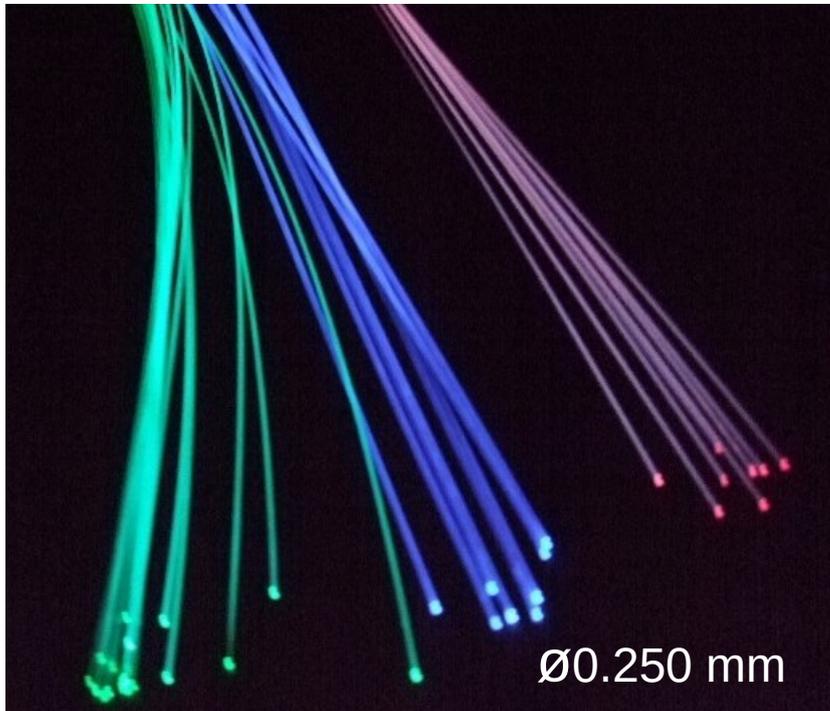
Upgrade summary

- Replace 1 MHz hardware trigger → 40MHz software trigger, all front-end electronics to 40 MHz
- Visible interactions per bunch crossing increase to $\mu = 2.5 - 5$ (from 1.8)
- Expected **annual** physics yields increase (with respect to 2011)
 - 14 Tev cross section (x2), trigger rate (\geq x4), luminosity (\geq x2.5)
 - **x10** in muonic channels
 - more than **x20** in hadronic channels
- **10 times smaller uncertainties after 10 years**

LHCb Upgraded Spectrometer

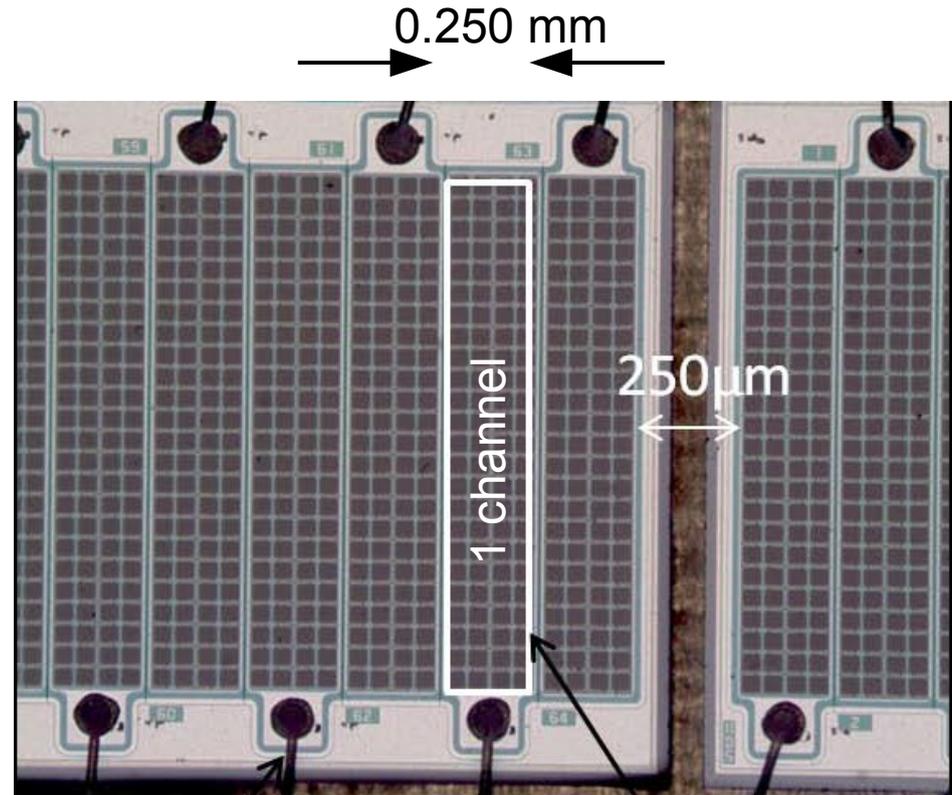


The SciFi Tracker



Scintillating fibres

- fast scintillation decay time (2.8ns)
- good light yield and attenuation length



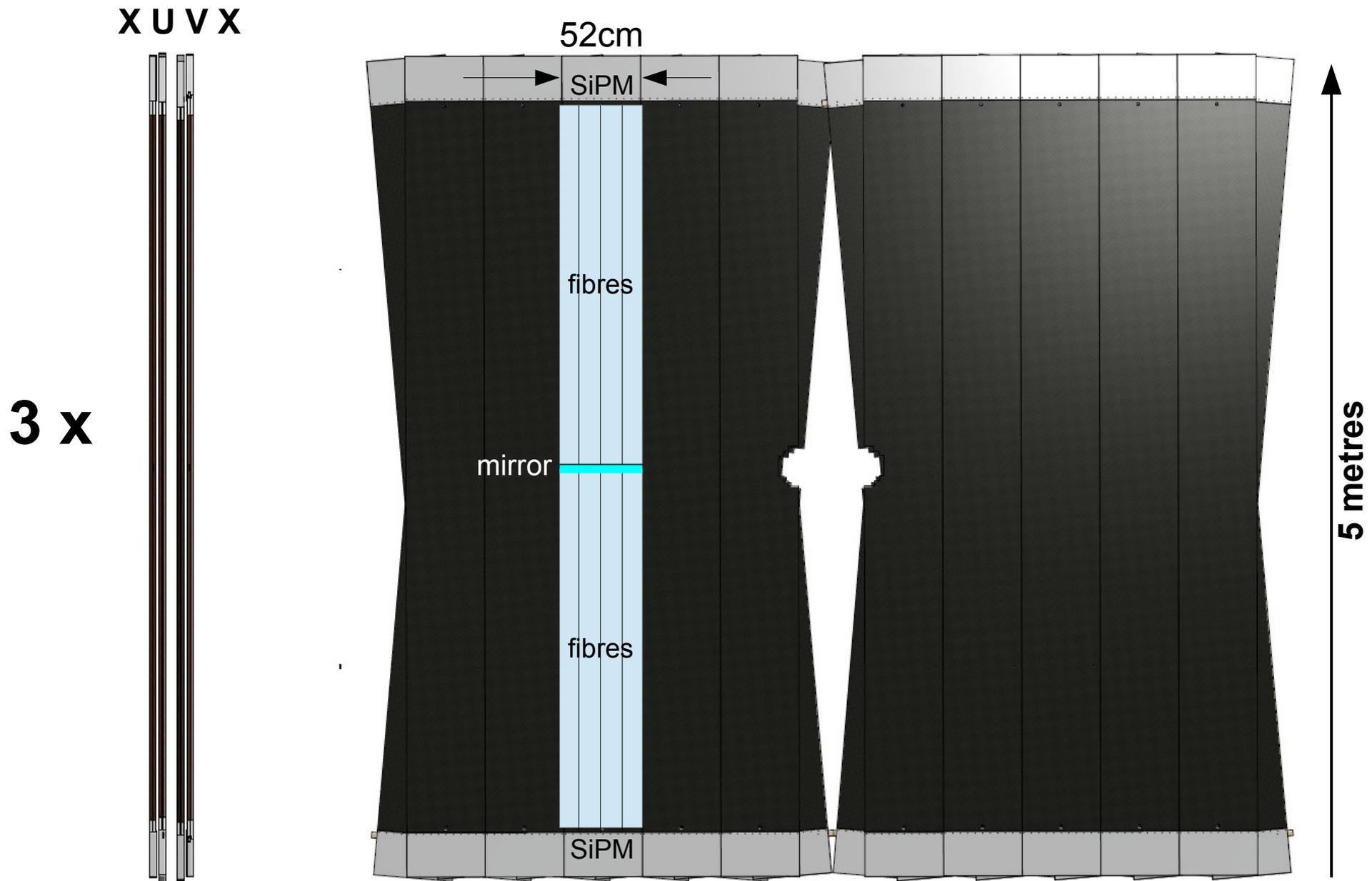
An array of pixelated silicon photomultipliers

- fast signals
- high photon detection efficiency (40+%)
- compact channel size

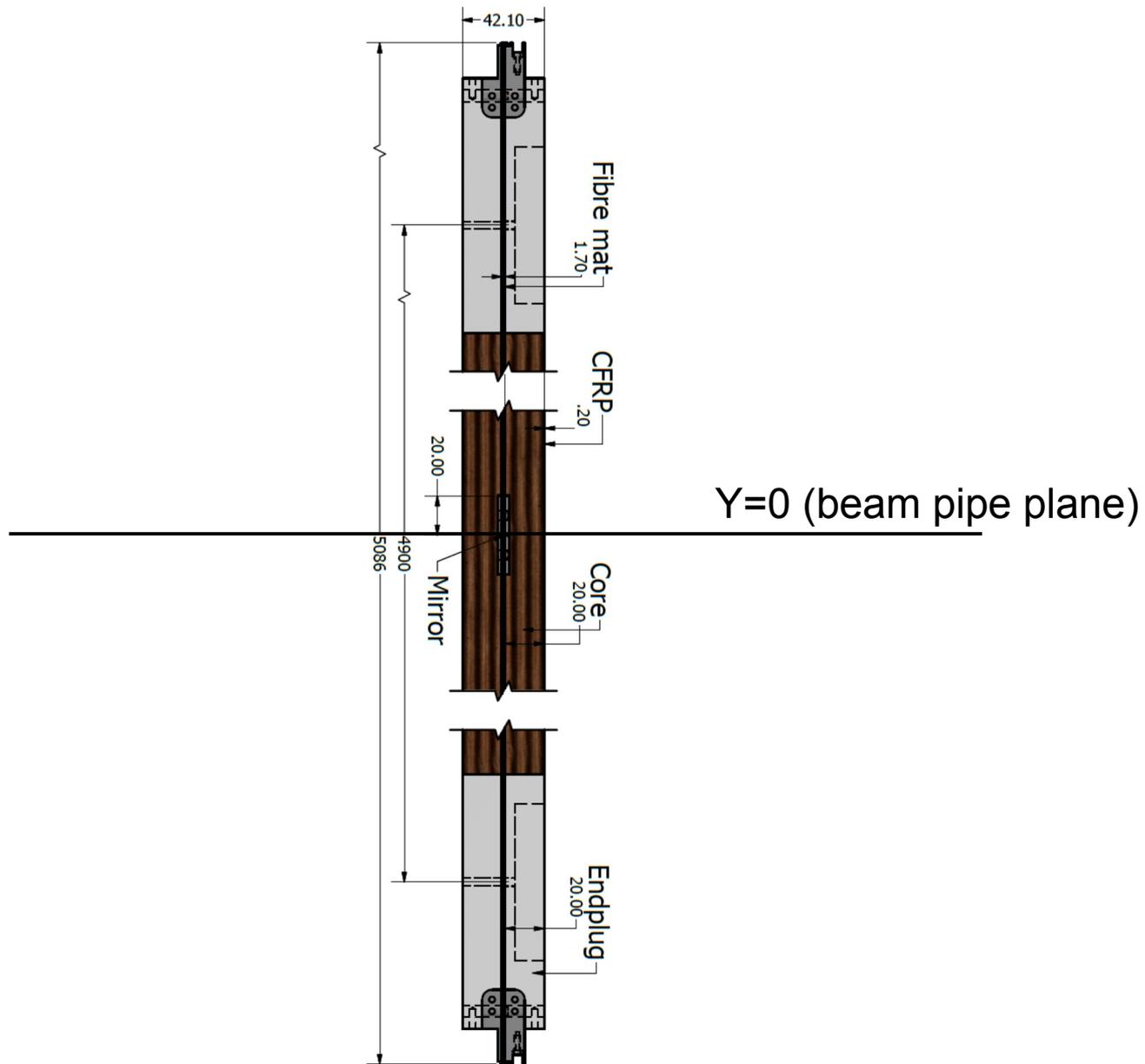
SciFi Collaboration

- 20 Institutions, 10 countries

	Task(s)	Institute(s)
Detector	SiPM assembly SiPM QA Fibre QA Fibre mat production Panel & module construction Read-out box Module testing (including electronics)	EPFL CERN, EPFL, NCBJ CERN, NCBJ, RWTH, TUD, HD Russia, RWTH, TUD, HD Russia, RWTH, TUD, HD CERN, EPFL, LPC, NIKHEF, RWTH CBPF, NIKHEF
Electronics	PACIFIC ASIC Front-end boards Tell40 board software	UB, IFC, LPC, NIKHEF, HD EPFL, LPC, NIKHEF, RWTH, HD LPNHE, TUD
Infrastructure	Frames Cooling	CERN, EPFL, NIKHEF CERN, RWTH

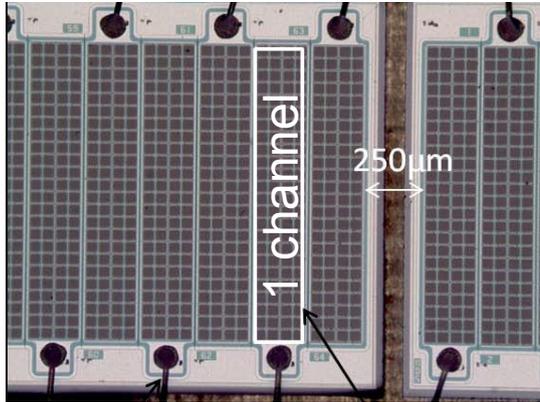


Modules

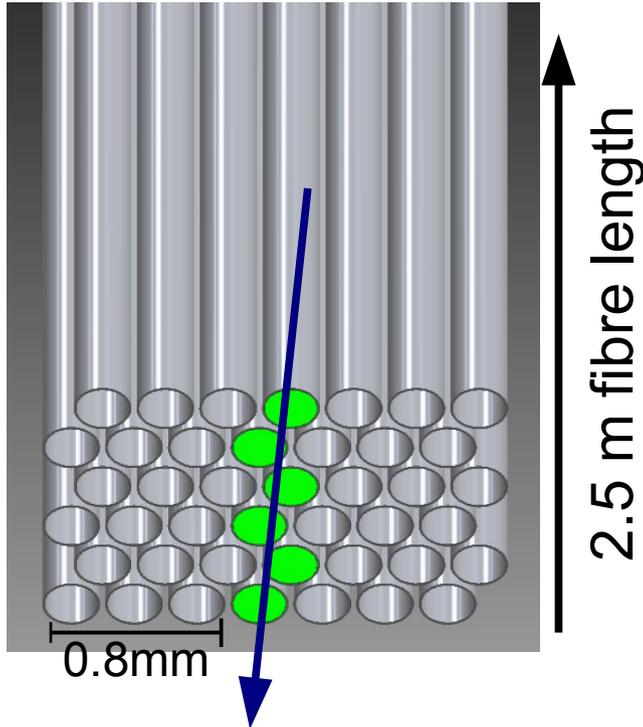


Basic principle

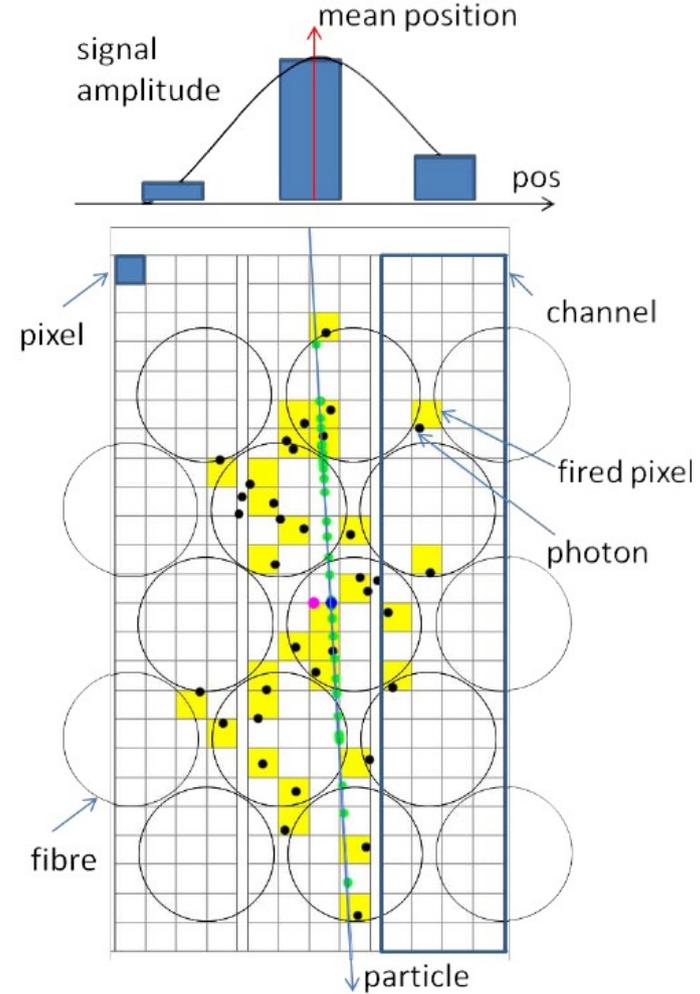
SiPM array



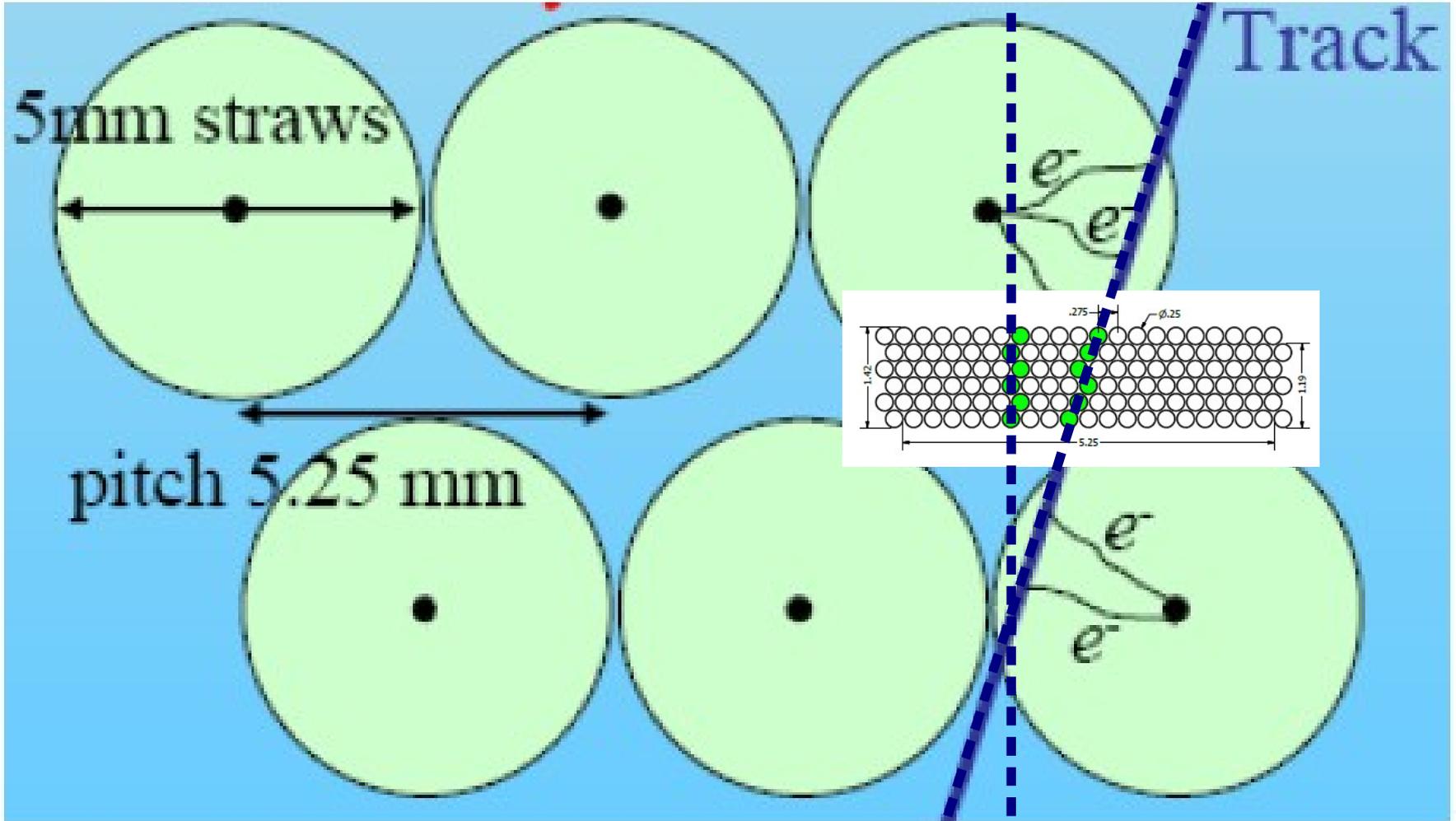
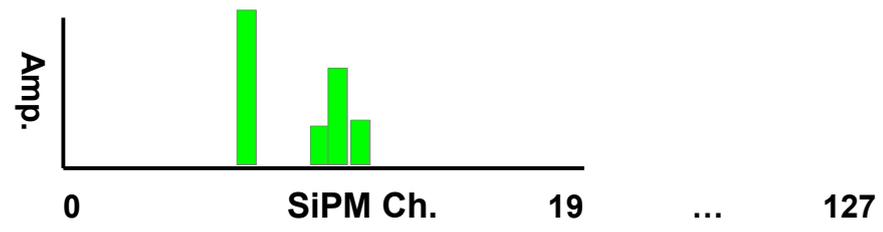
Scintillating
Fibres
(0.250mm
diameter)



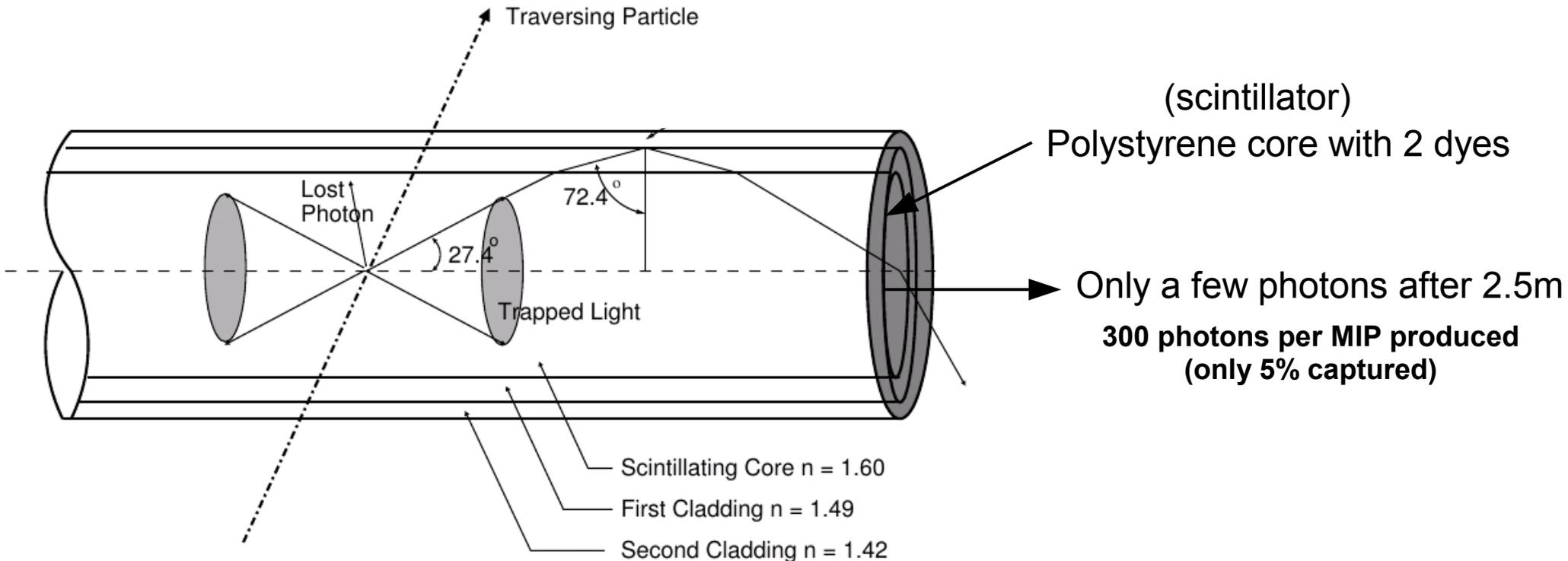
Signal cluster



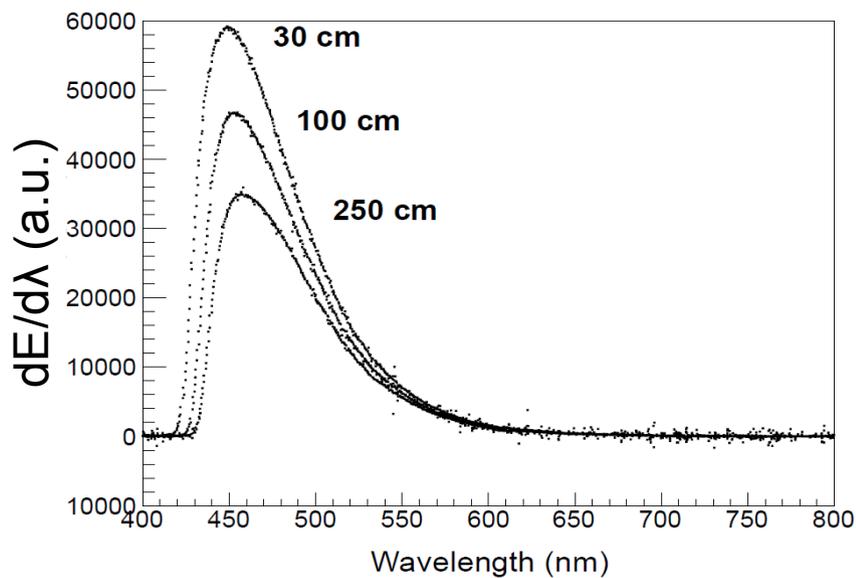
Typically one observes 15-20 photoelectrons for 5 layers of fibre



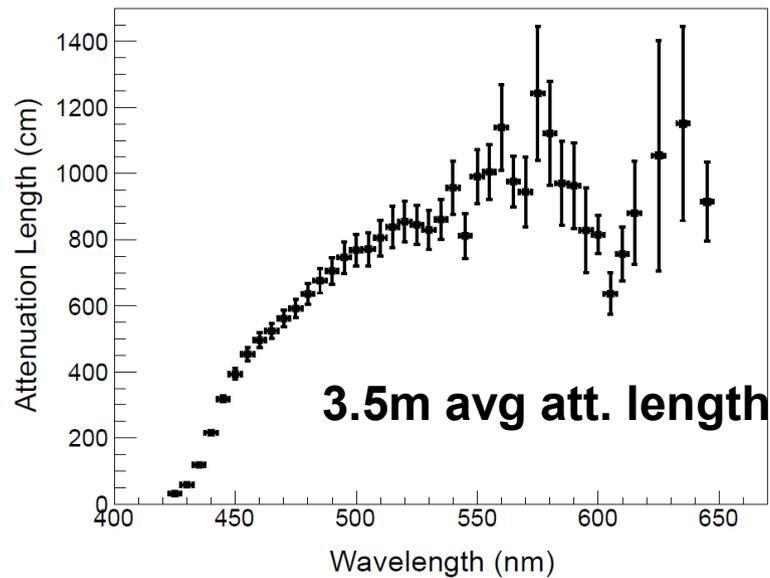
Scintillating Fibres



SCSF-78MJ



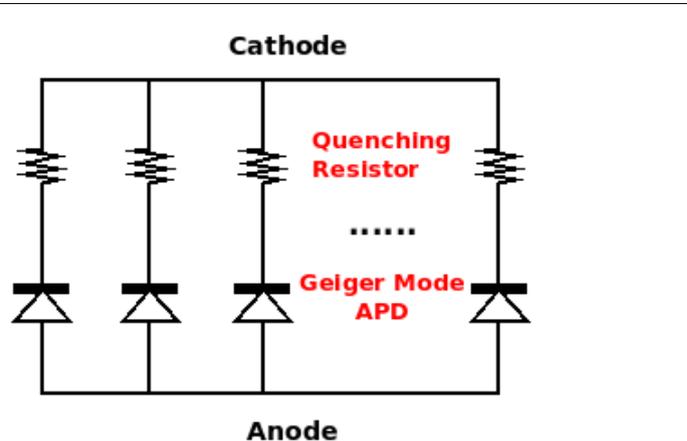
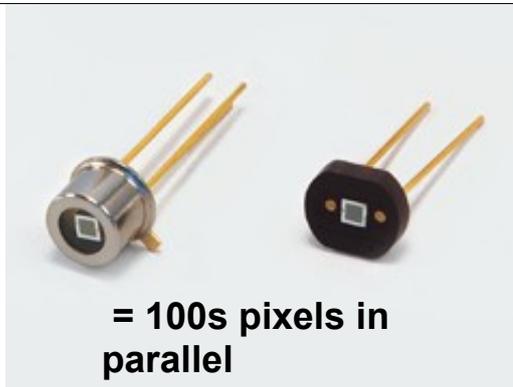
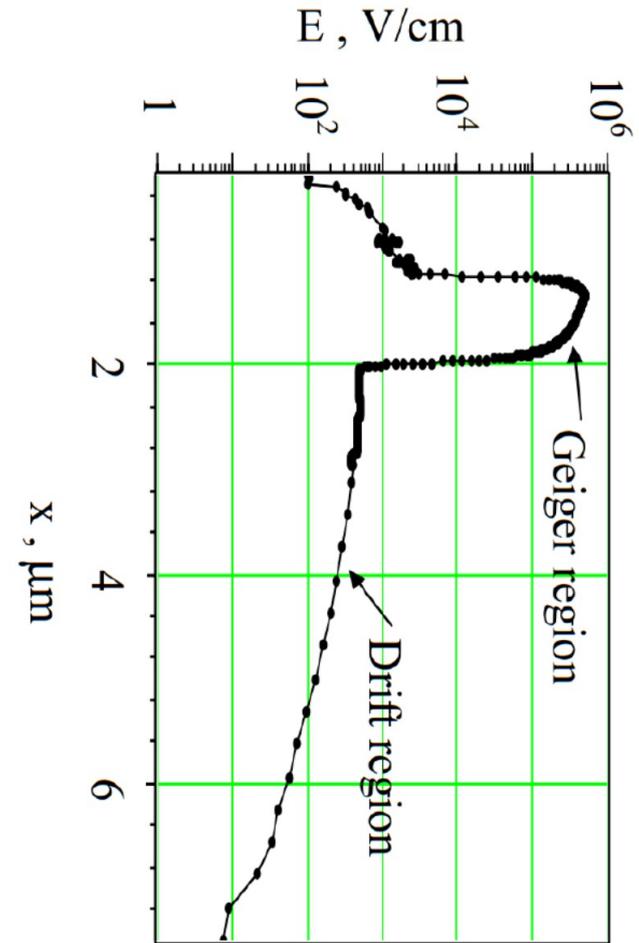
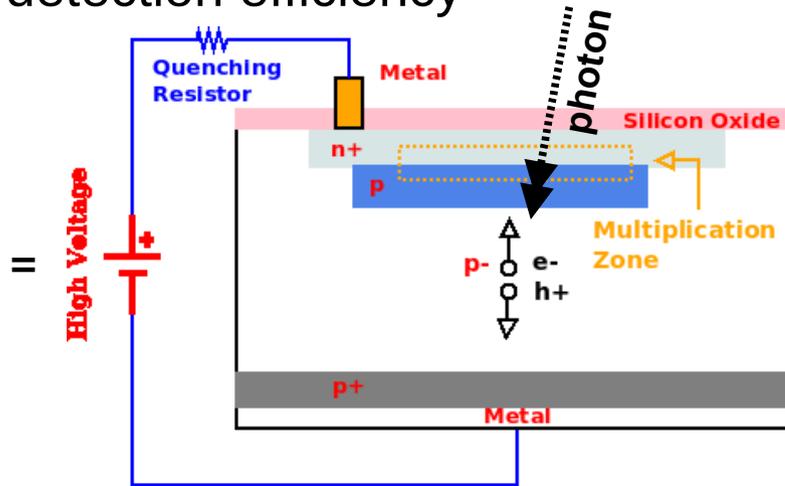
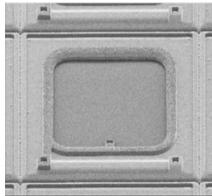
SCSF-78MJ



SiPMs

- The SiPM pixel is a photo-diode (reverse-biased, above breakdown Geiger mode)
- a single free electron/hole-pair can trigger an avalanche of electrons
- 10^6 — 10^7 gain
- 20-50% photon detection efficiency

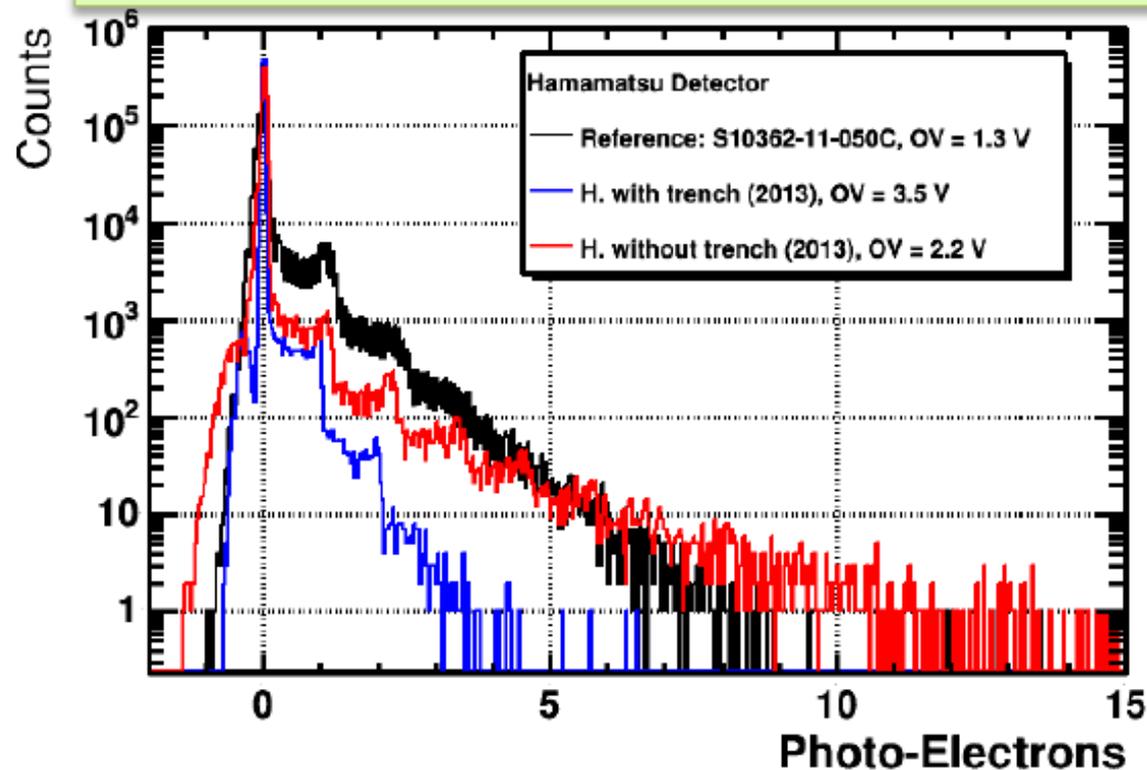
Each pixel



SiPM dark noise

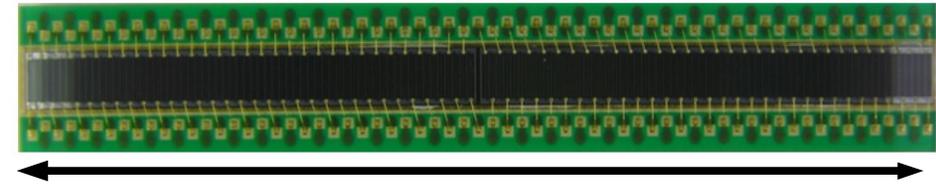
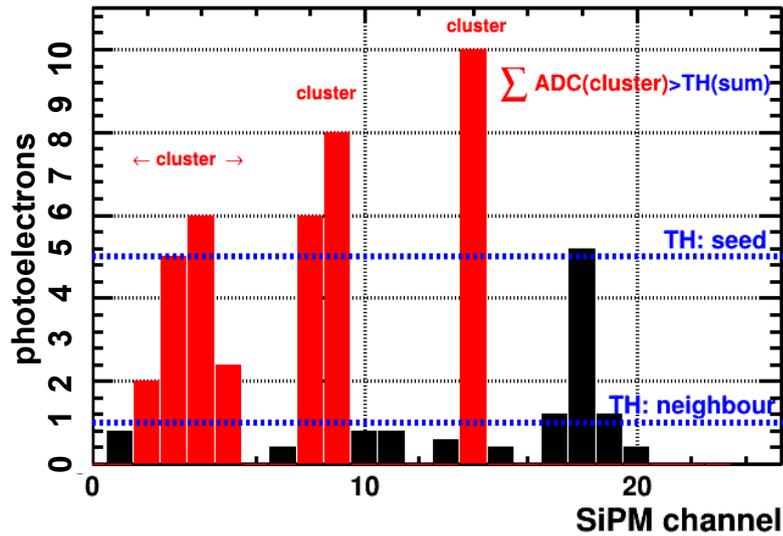
Unirradiated SiPMs produce about 100 kHz of >0.5 photoelectron signals per mm^2 at room temperature from **thermal excitations** and **pixel crosstalk**.

Hamamatsu, 50um pixel, standard (ref.), with trench (2013), without trench (2013), at typical over-voltage, recorded with VATA64.



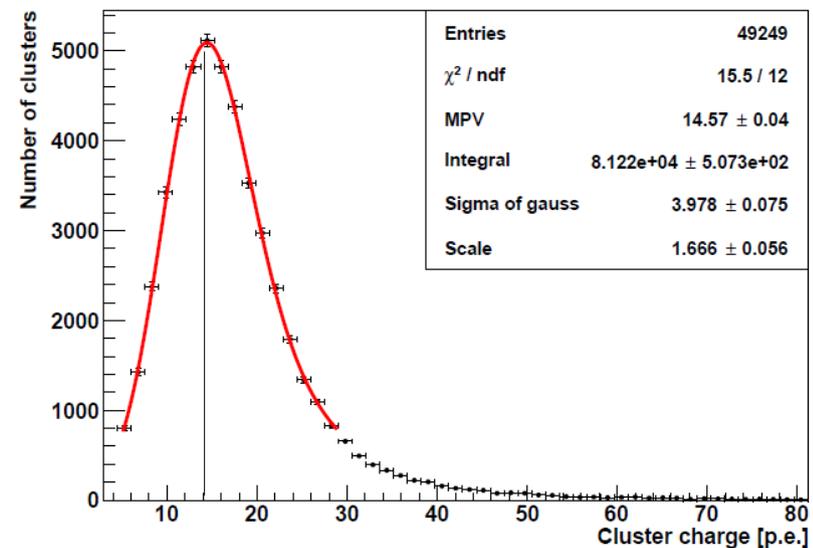
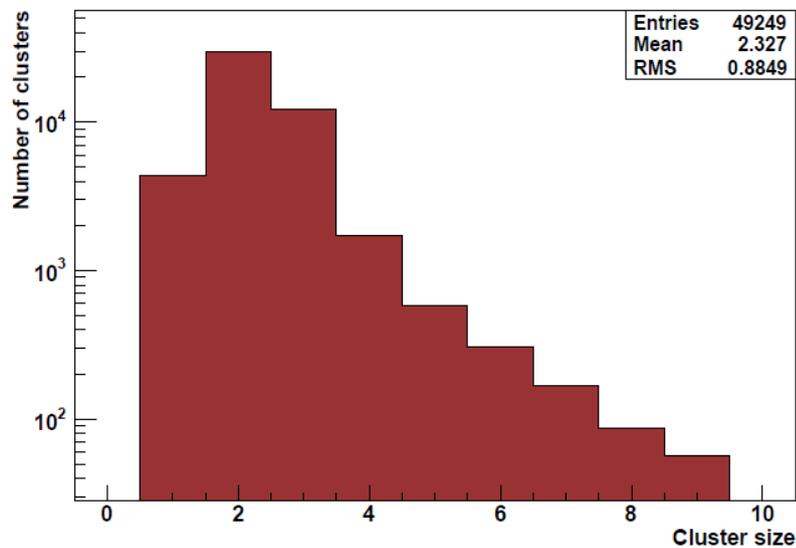
Plot from G. Haefeli (EPFL), presented at TIPP 2014

Clustering algorithm



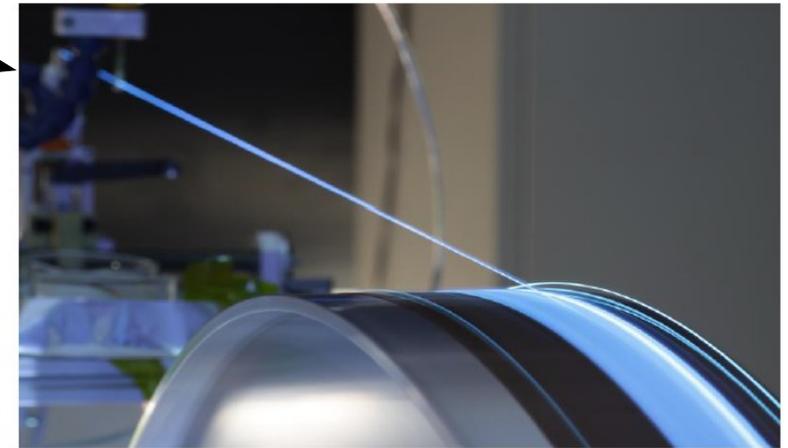
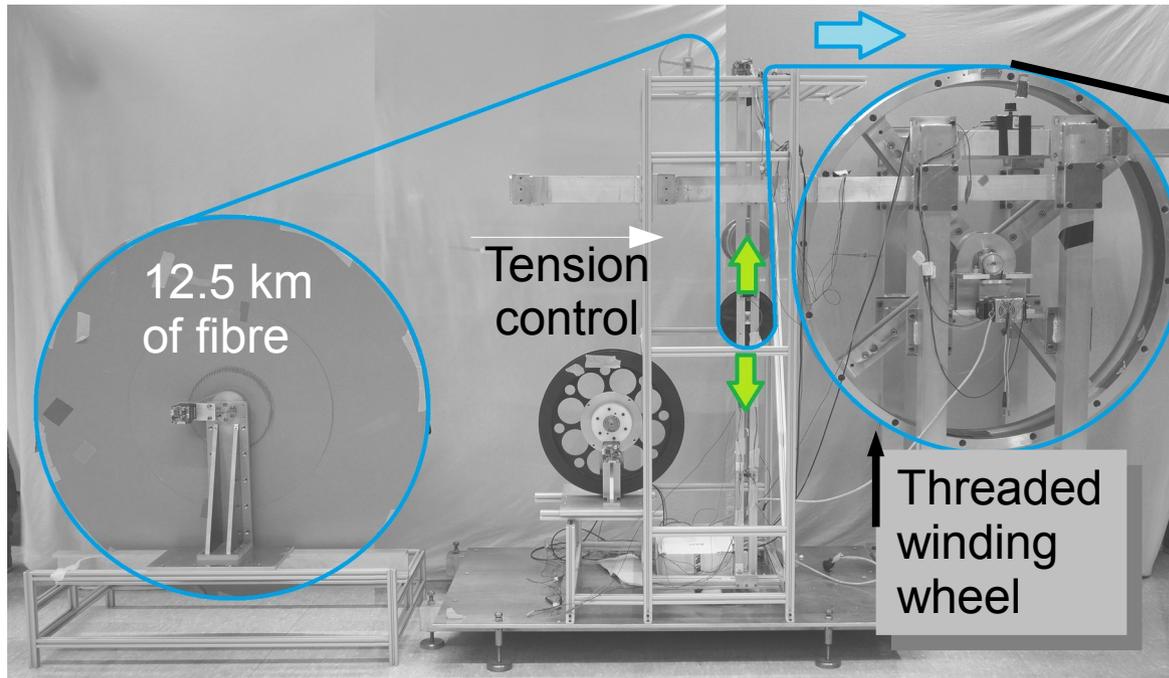
128 Channel SiPM array

Apply clustering and threshold cuts to reject dark noise clusters in the front-end electronics

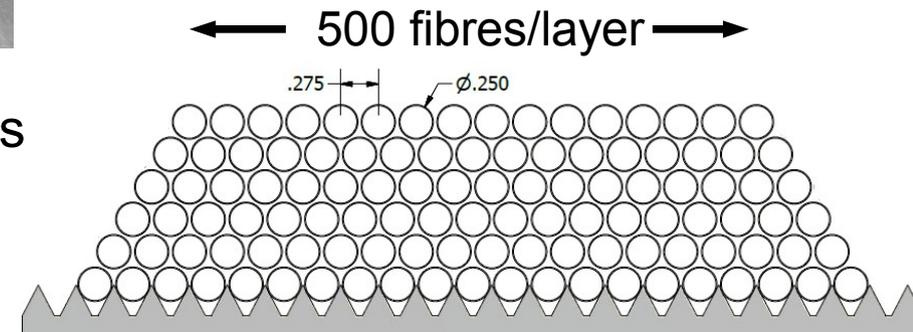


Fibre Mats

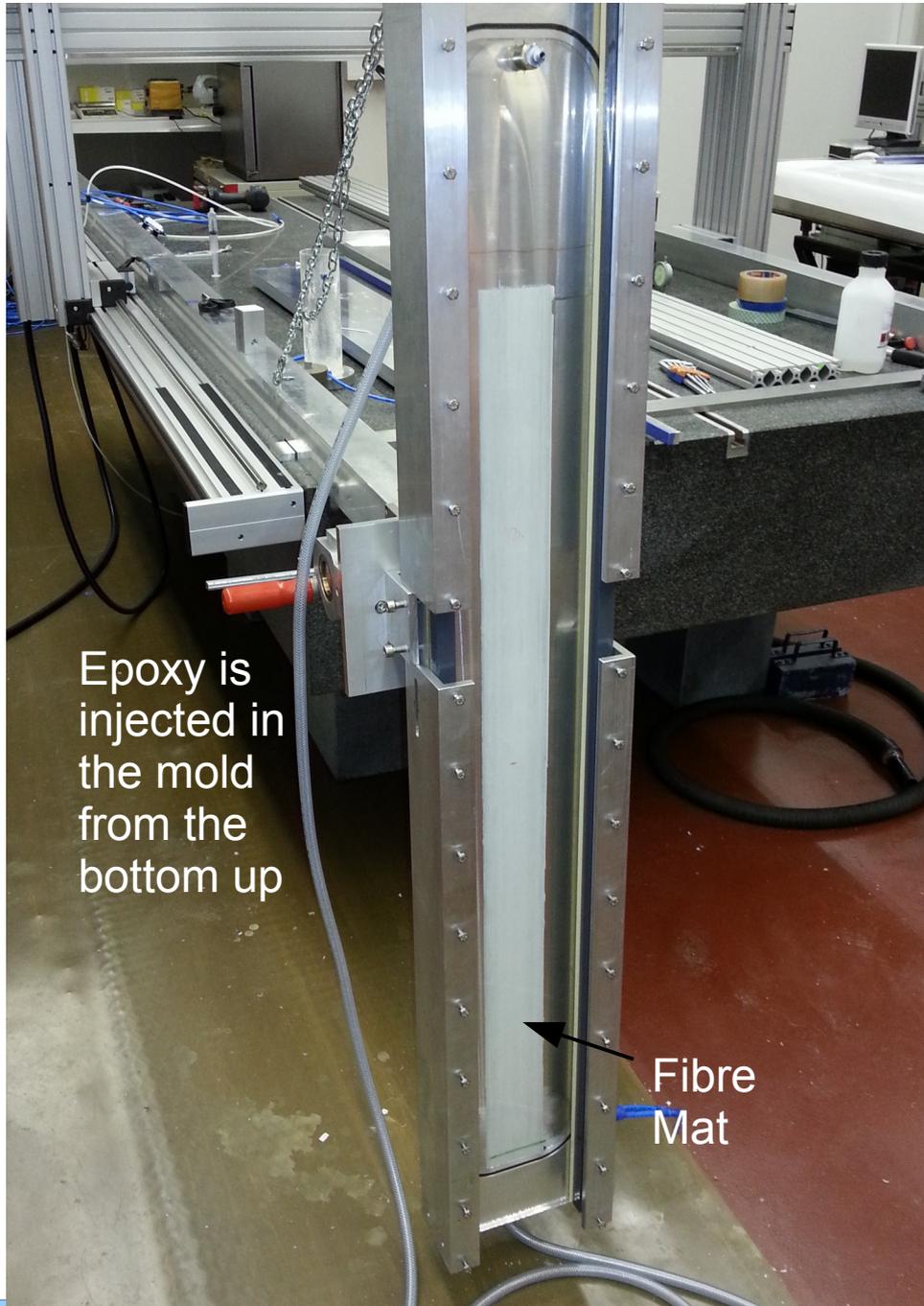
Fibre mats are produced from winding a single fibre onto a threaded wheel.



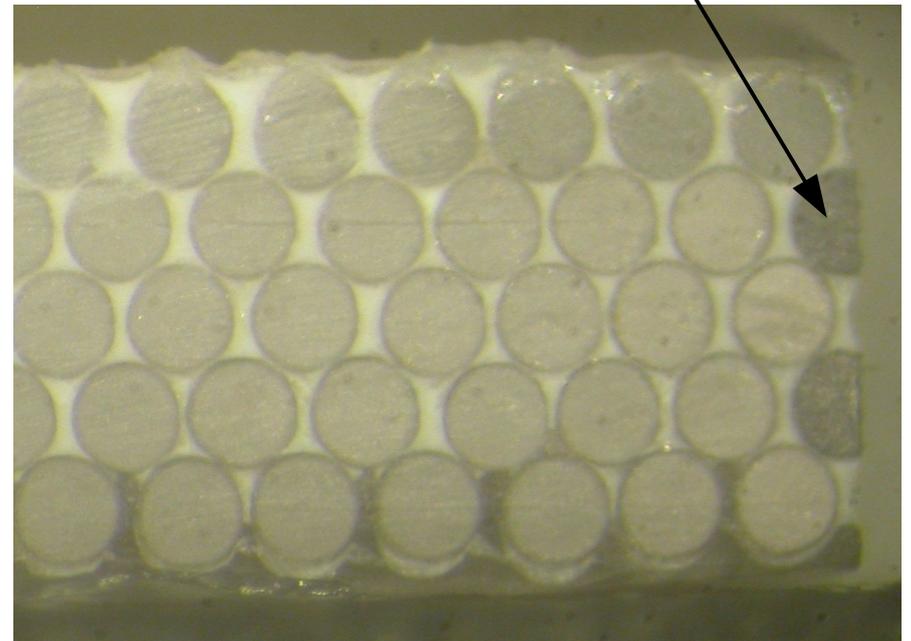
- Need about 8km of fibre for one mat of 6 layers 2.5 metres long
- 10,000 km of fibre in total ...

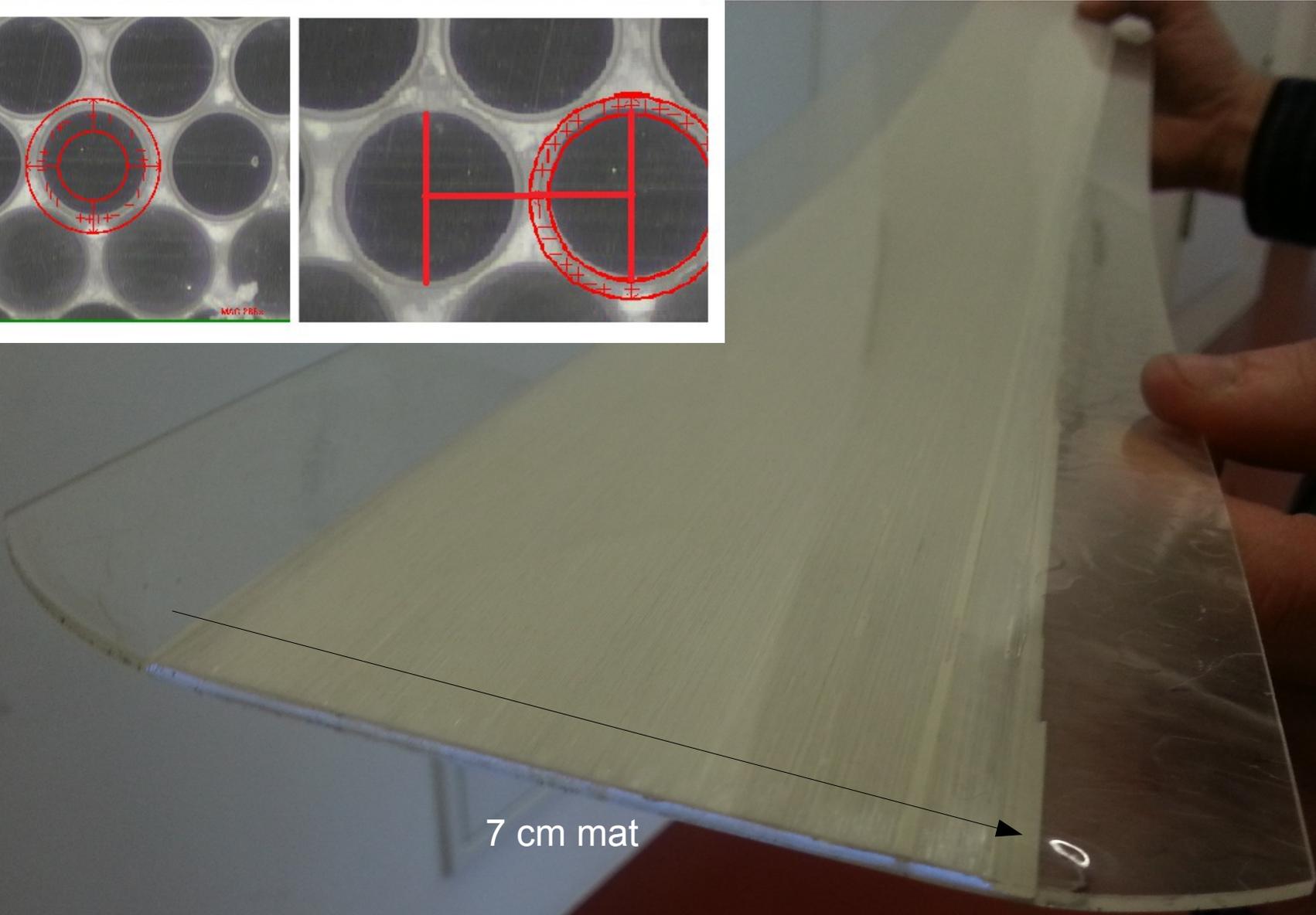
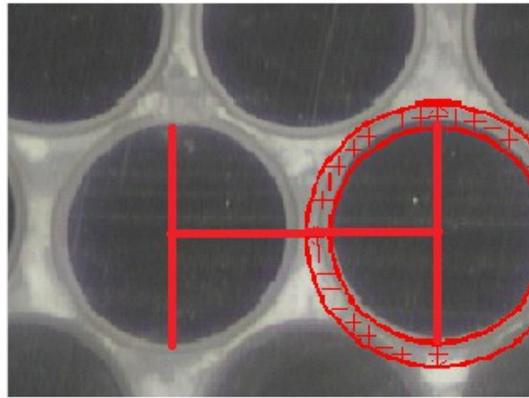
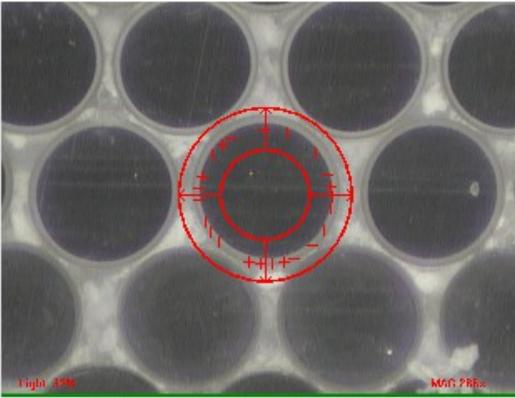
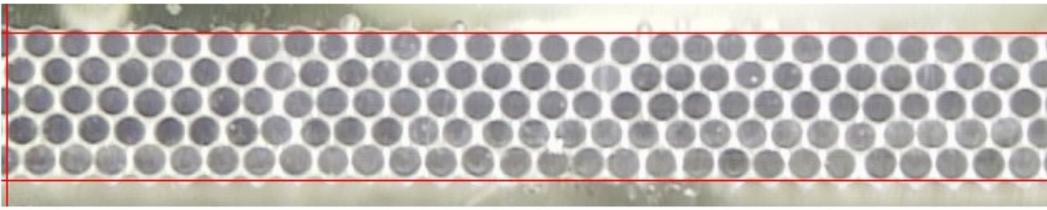


Fibre Mats



Cutting will create dead fibres on the edges





7 cm mat

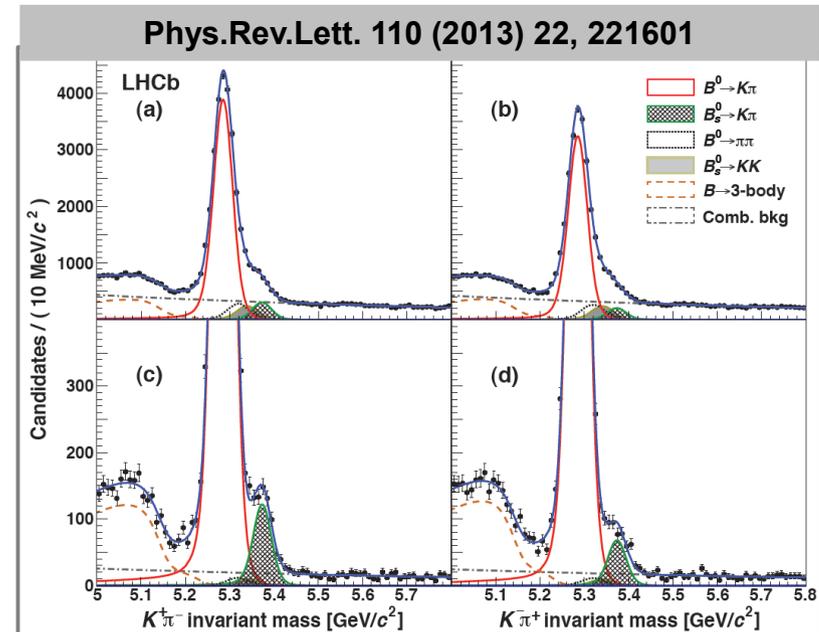
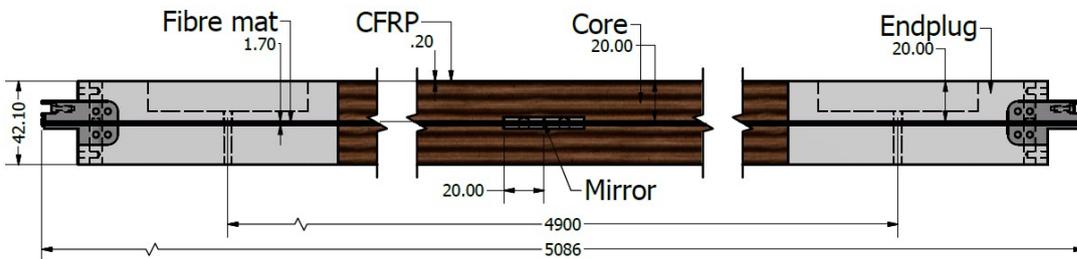
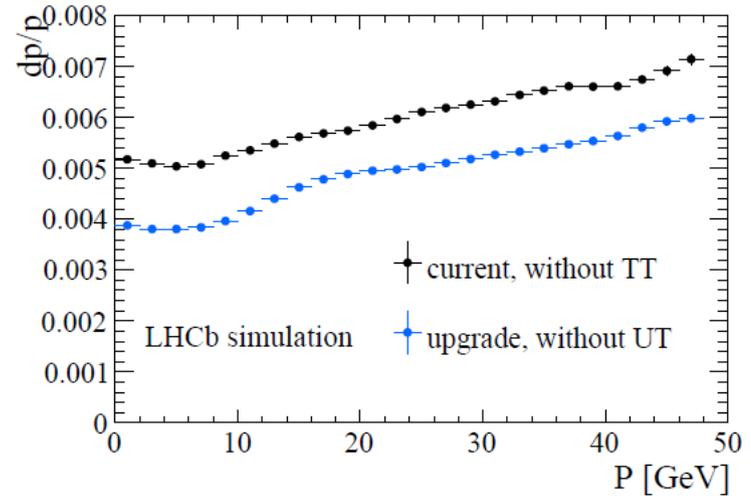
13.5 cm (500 fibres wide) mats are now being produced as well

Challenges: Detector design

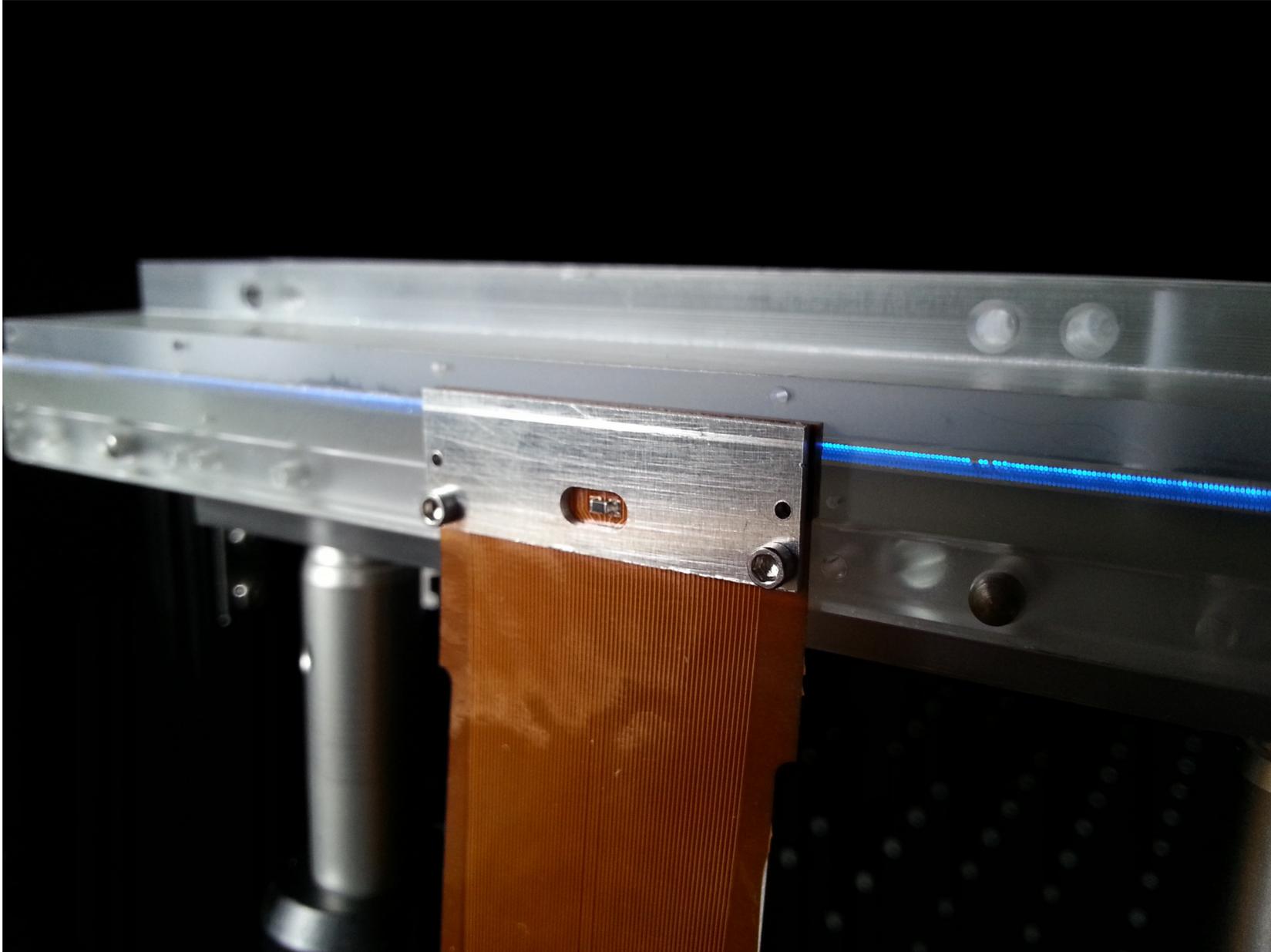
- Stability and alignment of the detector must be $\sim 100\mu\text{m}$
- Must be $< 1\%$ of a radiation length per detector layer (4mm equiv. of plastic)



Less material + stable detector =
improved momentum resolution =
better mass resolution

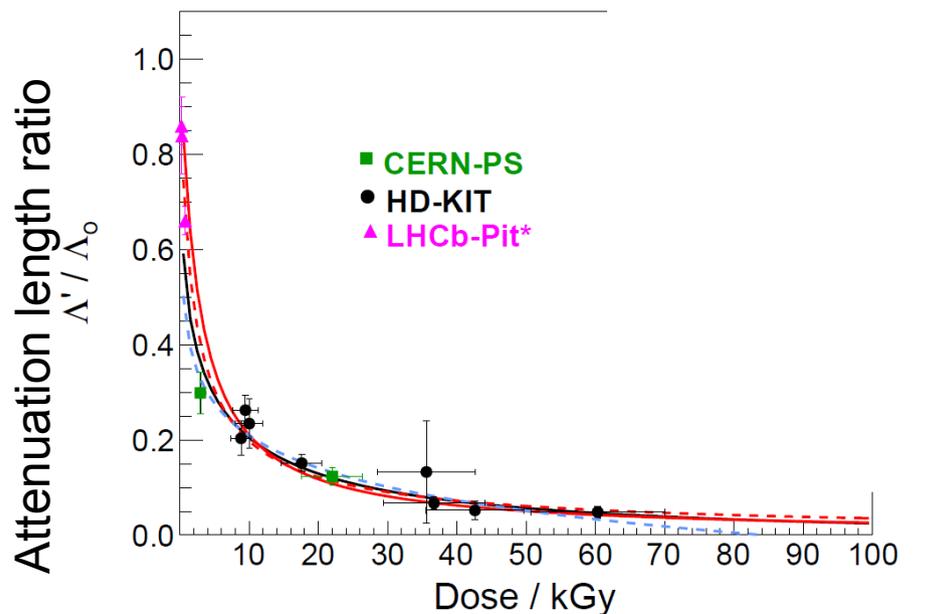


Prototypes and Test-beam in October 2014! :)



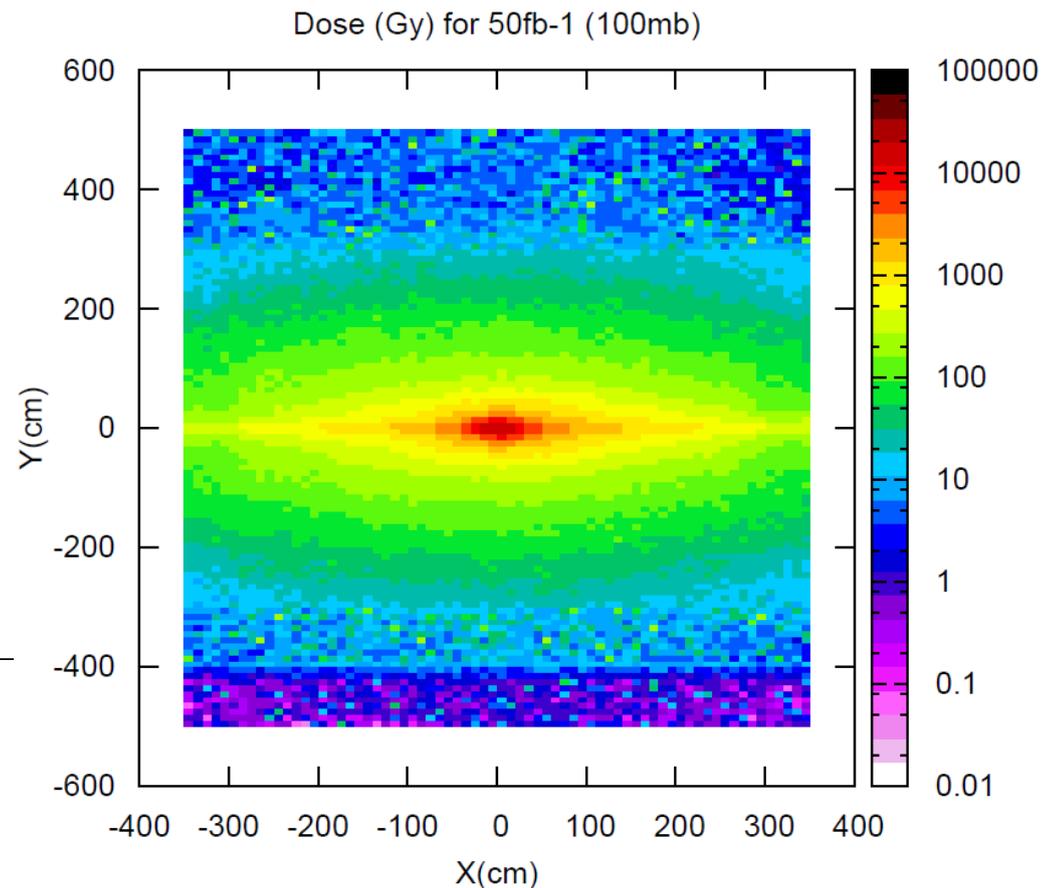
Challenges: Fibre irradiation

- The scintillating **fibres darken with radiation** (up to 35 kGy expected near the beam pipe over the upgrade lifetime)



Expect a 40% loss in signal near the beam pipe after 10 years

Expected ionizing dose for LHCb Upgrade

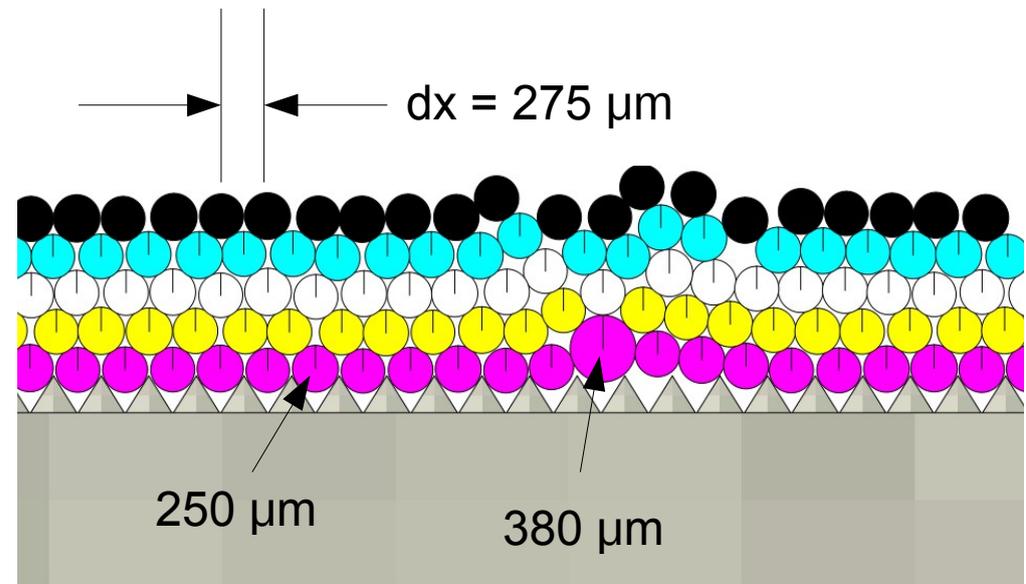
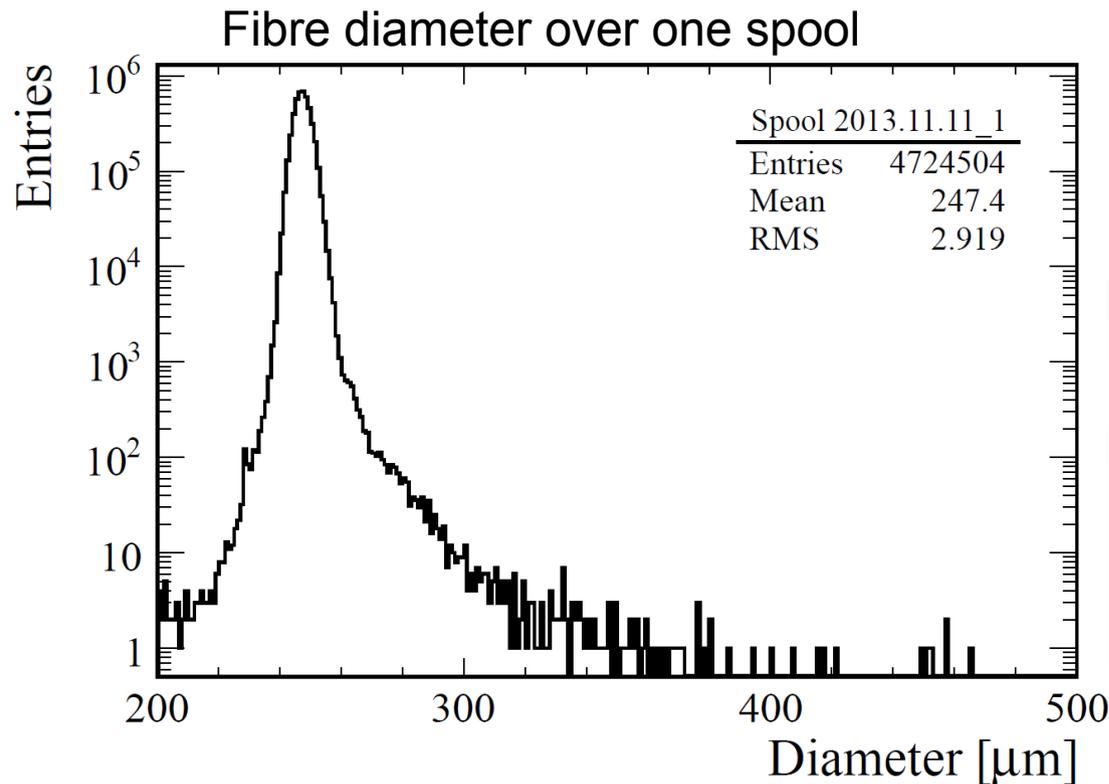


Challenges: Fibre bumps

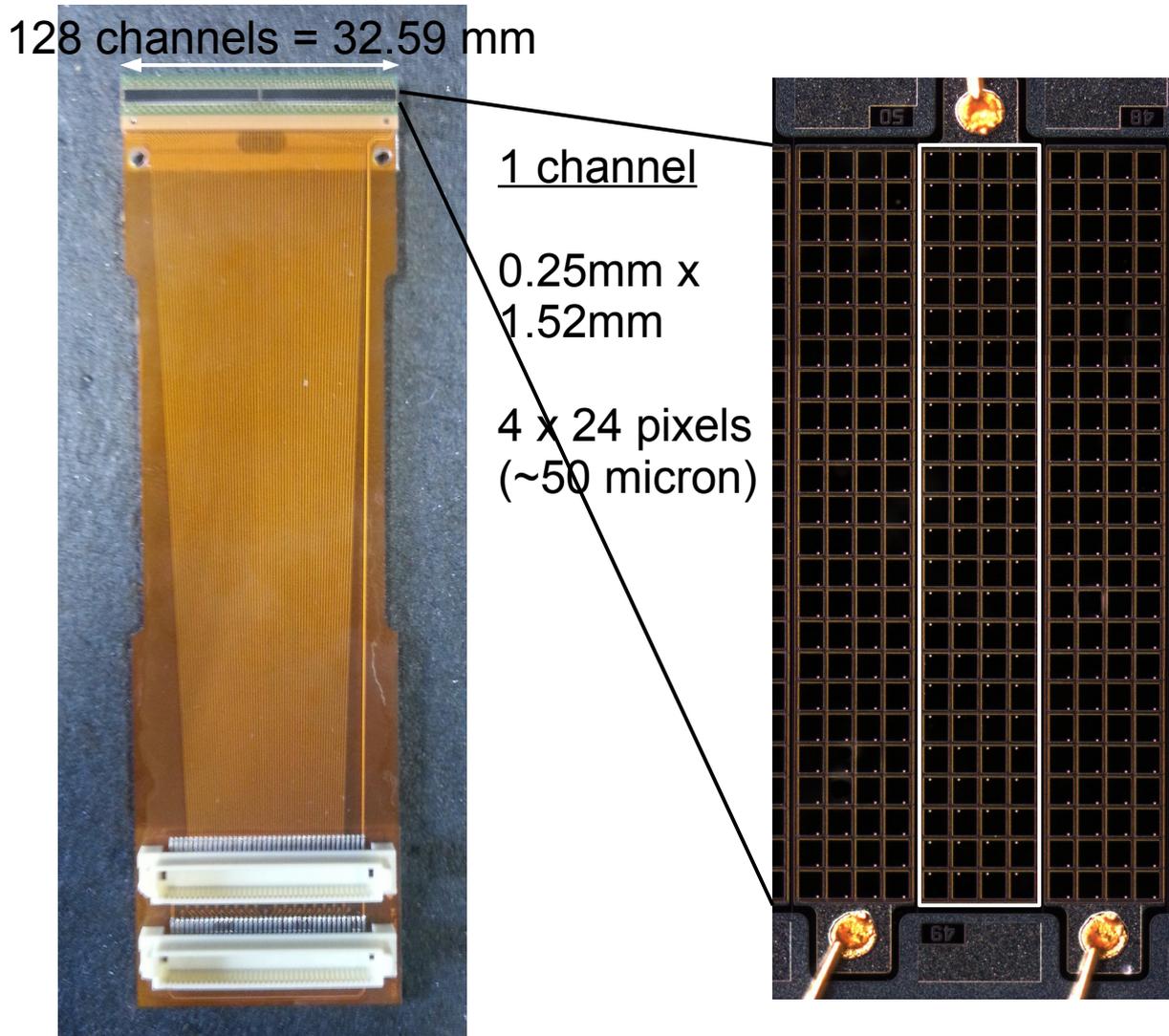
- Defects of the fibre can be created during the extrusion process making “blobs”

For a good mat
 $\text{rms}(dx) = 8-15 \mu\text{m}$

$\text{rms}(dx) = 8-15 \mu\text{m}$



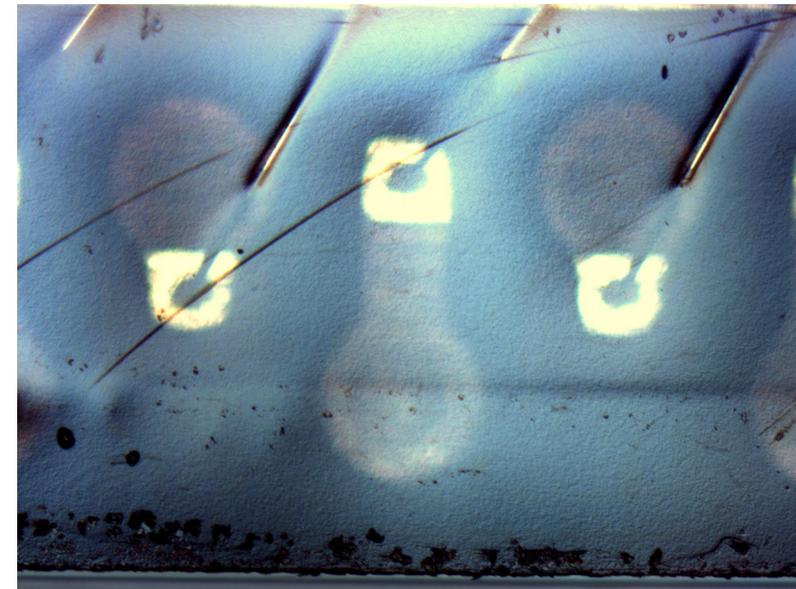
Challenges: SiPM array development



Issues with packaging:

- dead channels
- exposed bondwires

A fix is promised from the companies....

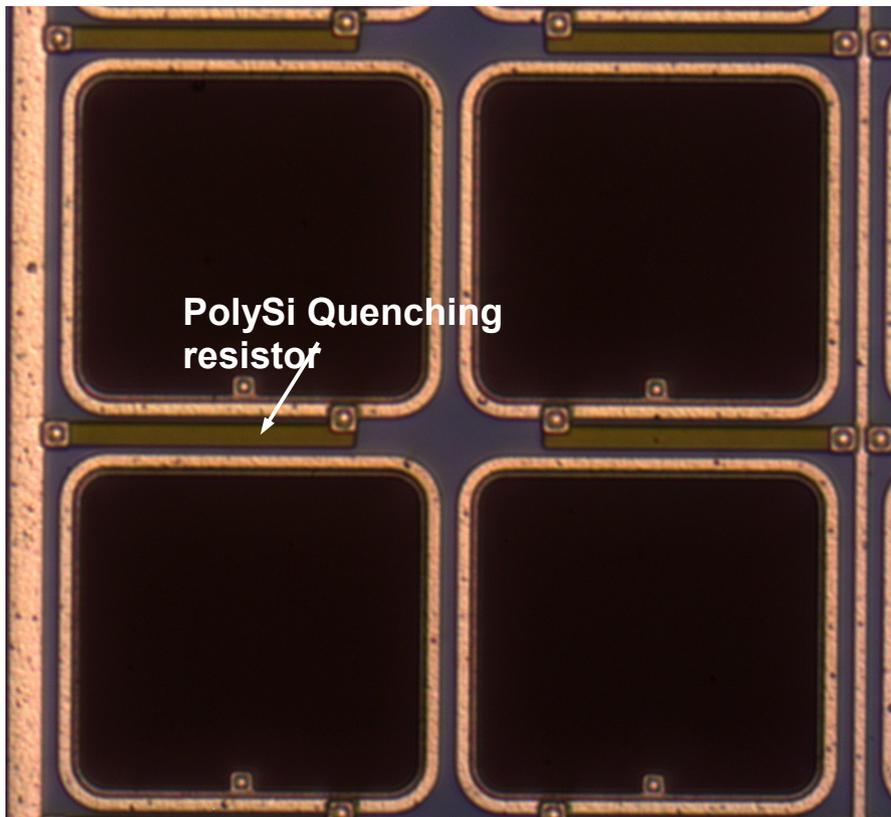


Old Hamamatsu arrays (~2010)

Operating voltage = ~ 72 V

Overvoltage = ~ 1 V

Crosstalk = 21%



New Hamamatsu arrays (2014)

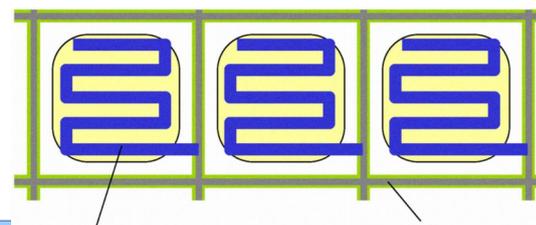
Operating voltage = ~ 52 V

overvoltage = ~ 2.5 V

Crosstalk = $\sim 5\%$

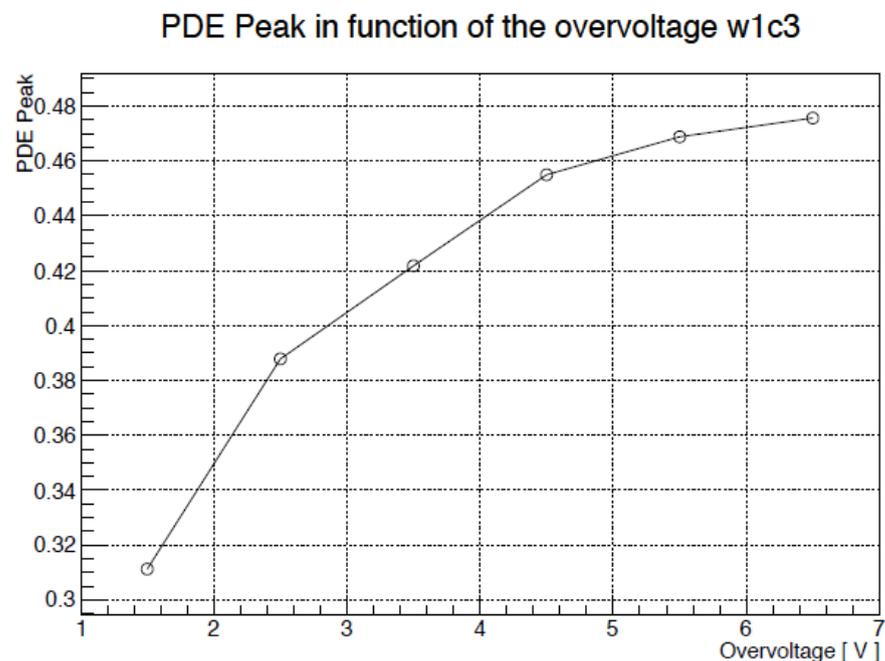
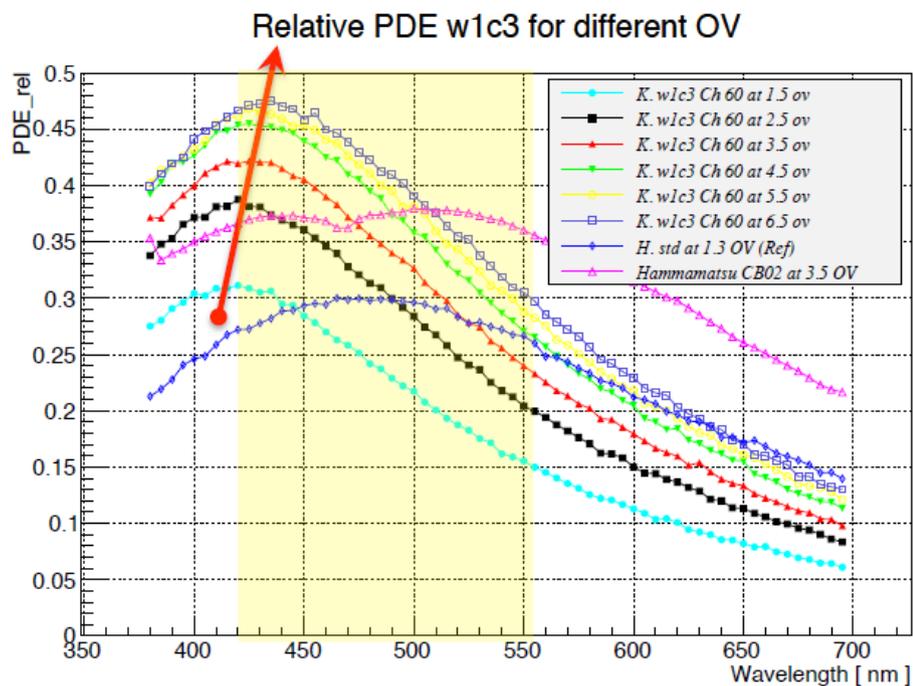


Thin metal film quenching resistor =
higher PDE, lower TCR



(photon detection efficiency)

PDE Hamamatsu and Ketek

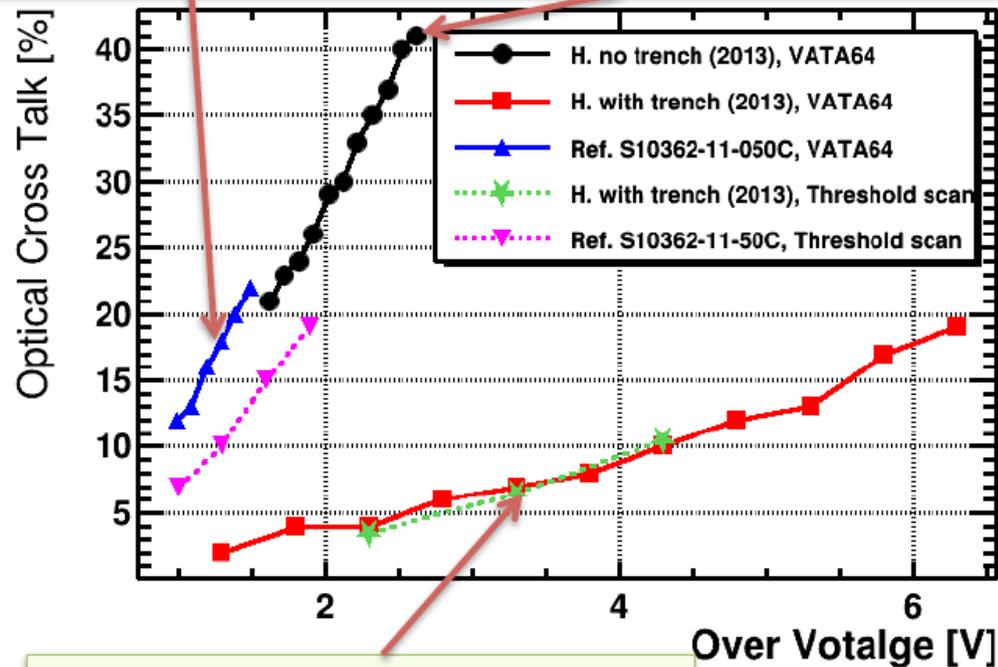


Plots from G. Haefeli, presented at TIPP 2014

(pixel) X-talk Hamamatsu and Ketek

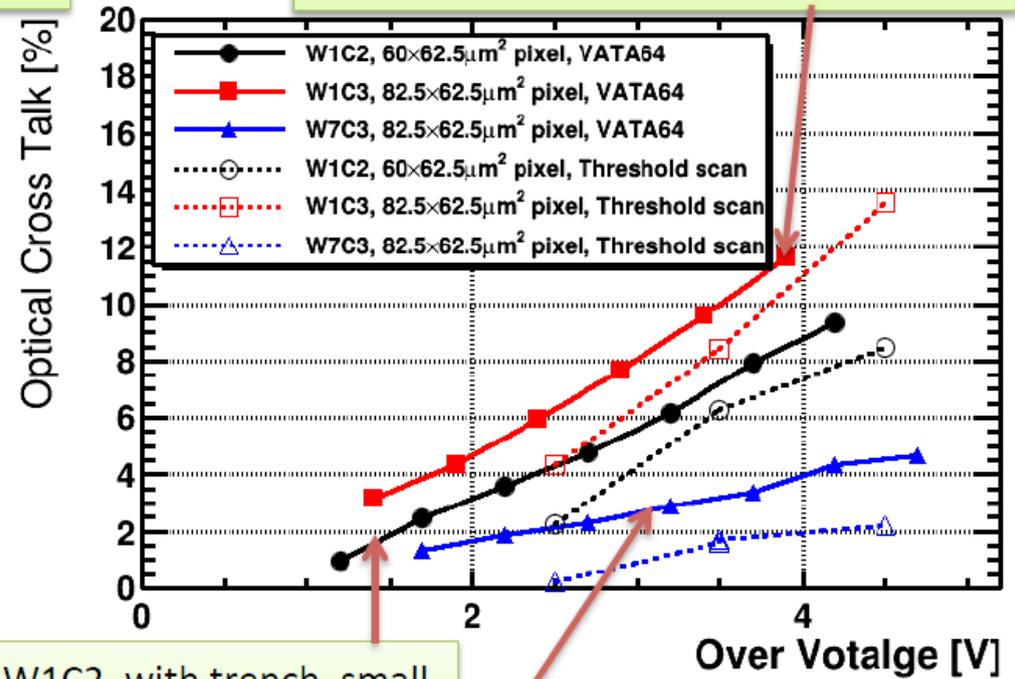
Difference due to after-pulsing
(1.3V over-voltage) (17% vs 10%)

No trench, extremely
high x-talk



With trench, low after-pulsing, 7%
x-talk at 3.5V over-voltage

W1C3, with trench, large pixels (3.5V
over-voltage) (10% vs 8%)



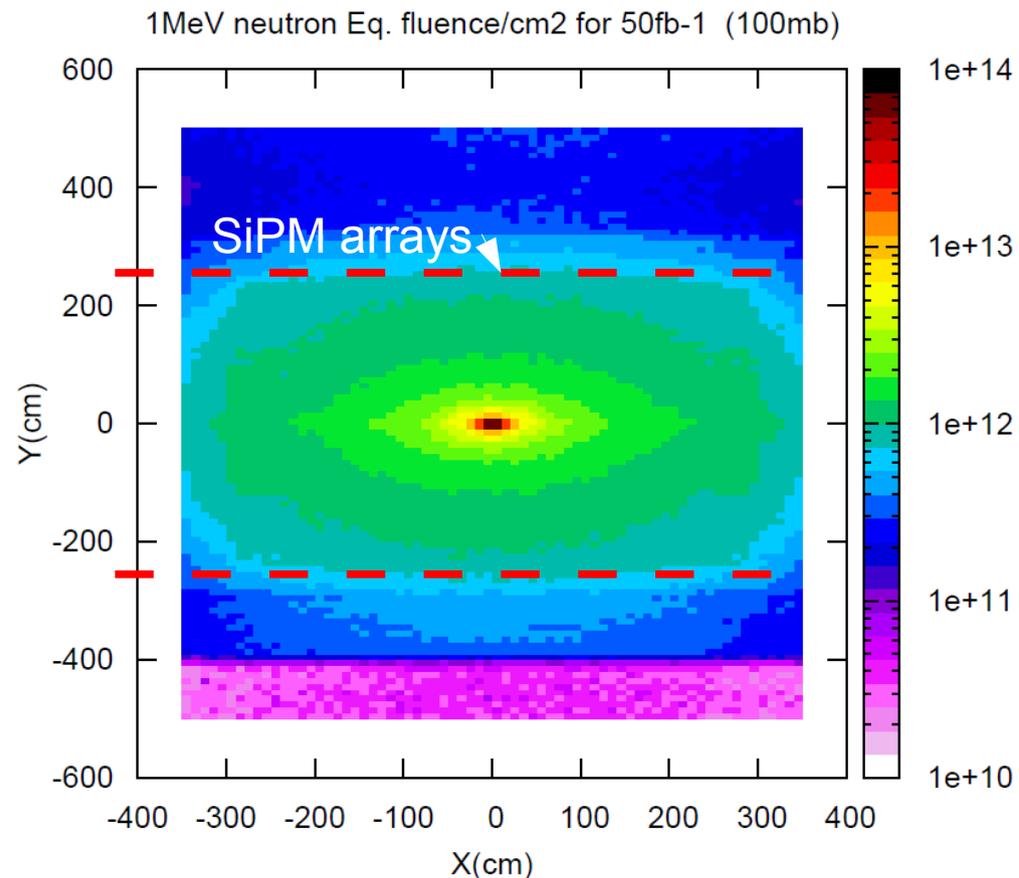
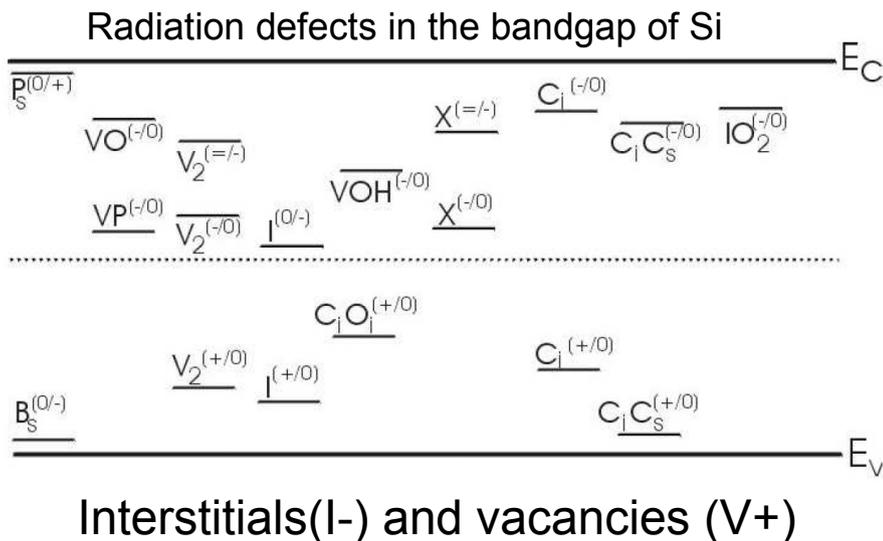
W1C2, with trench, small
pixels (3.5V over-voltage)
(8% vs 7%)

W7C3, double trench, large
pixels (3.5V over-voltage) (3%
vs 2%)

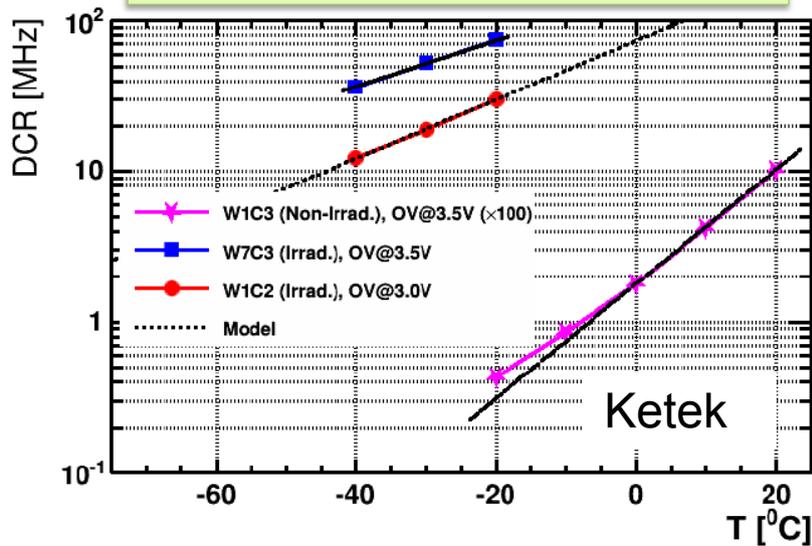
Plots from G. Haefeli (EPFL), presented at
TIPP 2014

Challenge: Radiation damage to SiPMs

- SiPMs create **single photo-electron signals from thermal electrons**, cross-talk between pixels makes 1 photo-electron look like 2+

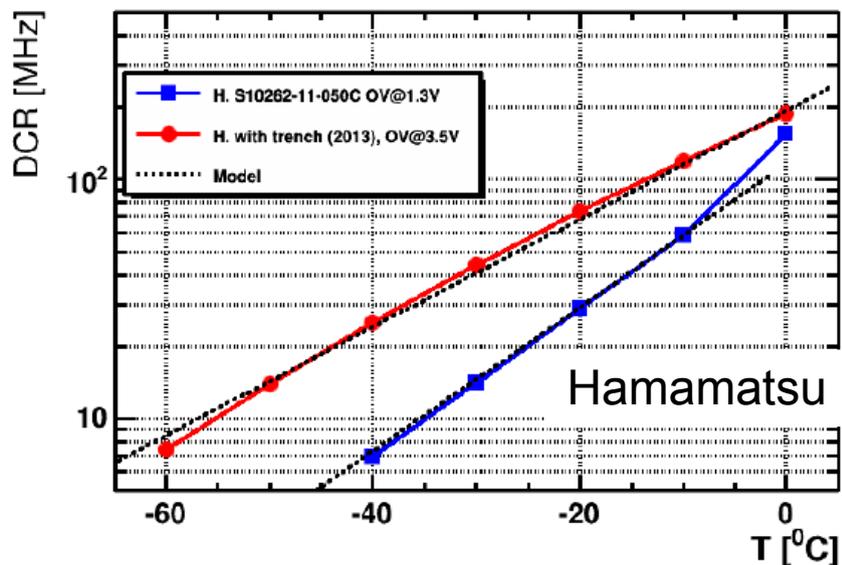


1/3mm², 2×10¹¹ n_{eqv}/cm² fluence at different temperatures

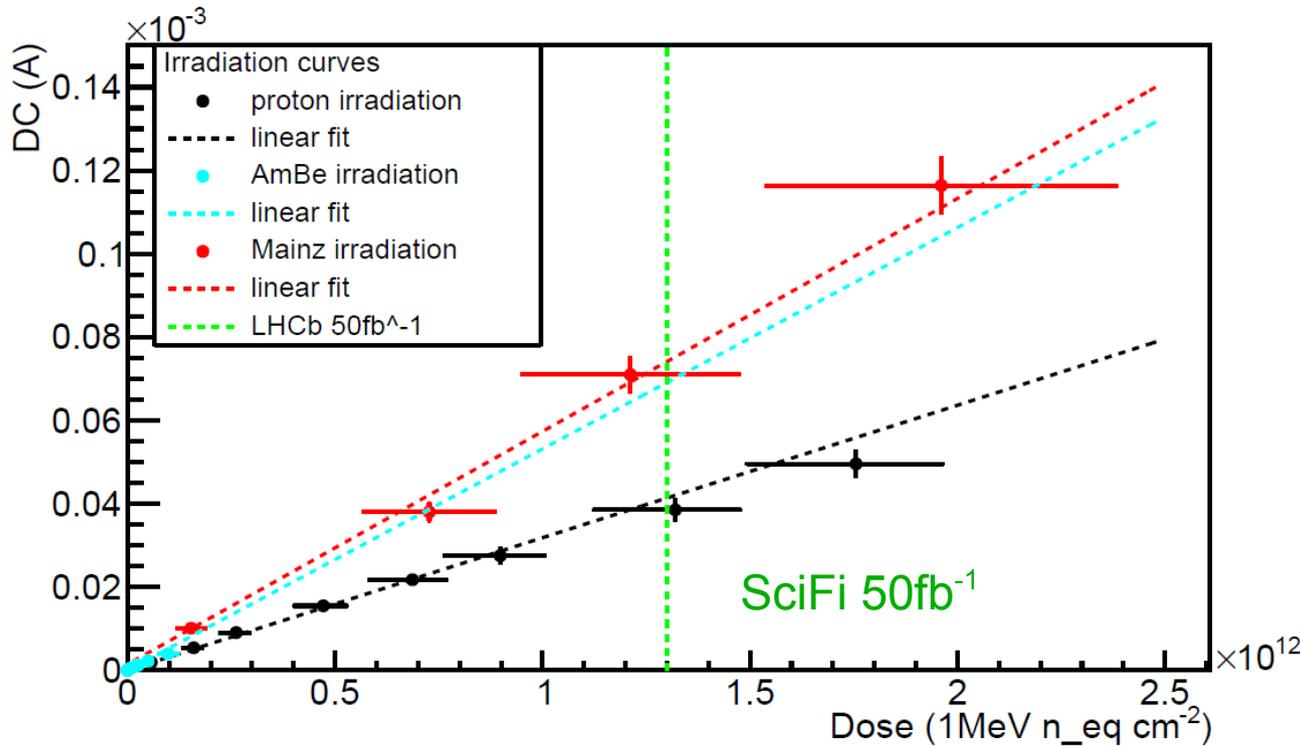


- We expect 1.3 x 10¹² n_{eq}/cm²
- Requires cooling to -40C
- So far, Ketek shows 2-3 times worse noise due to irradiation compared to Hamamatsu

1mm² single channel, 3.1×10¹¹ n_{eqv}/cm² fluence at different temperatures



For a Hamamatsu S12571-50C 1mm² SiPM
 (@-30C, no trenches)

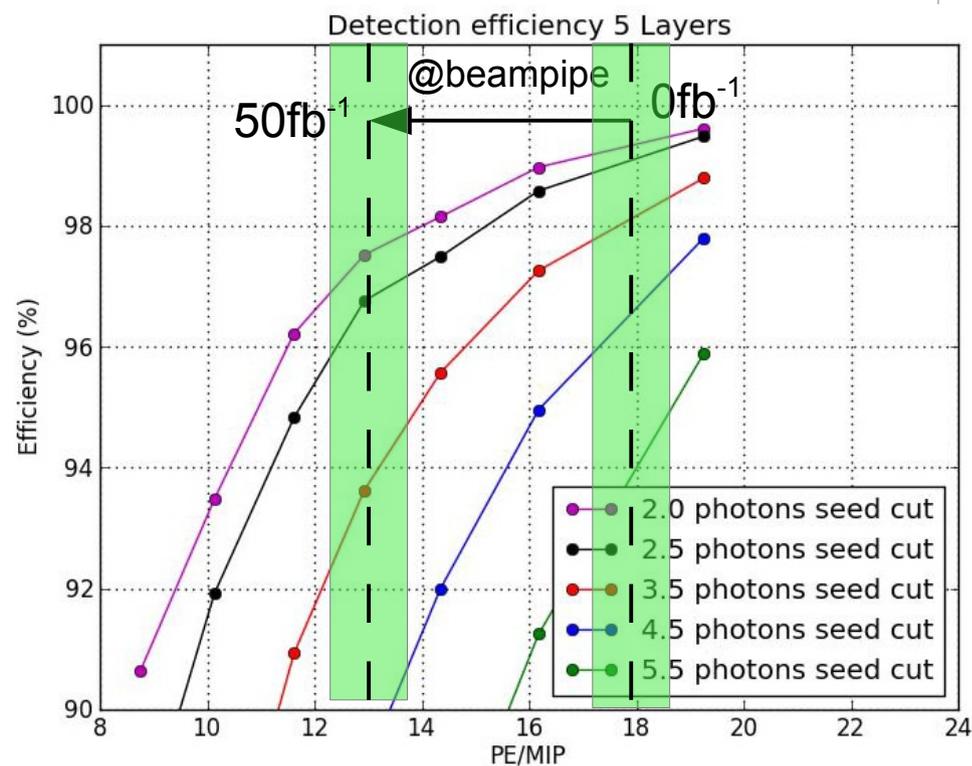
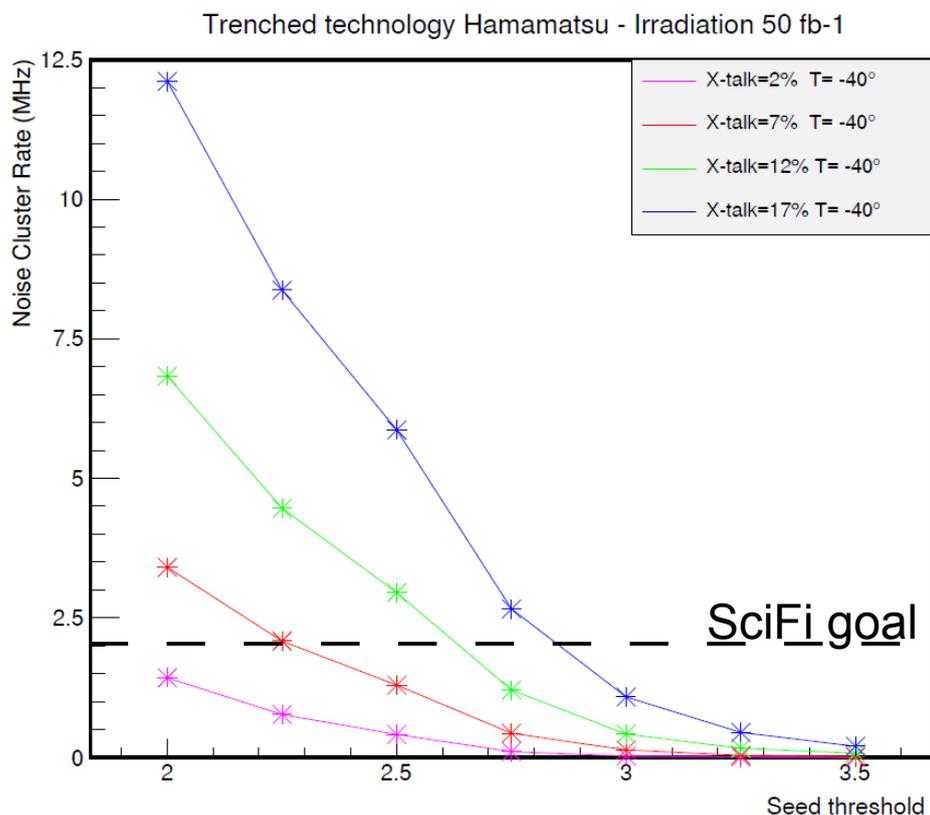


Expect about **20 MHz** of
 >0.5 p.e. noise per
 channel after 1.3×10^{12}
 n_{eq}/cm^2

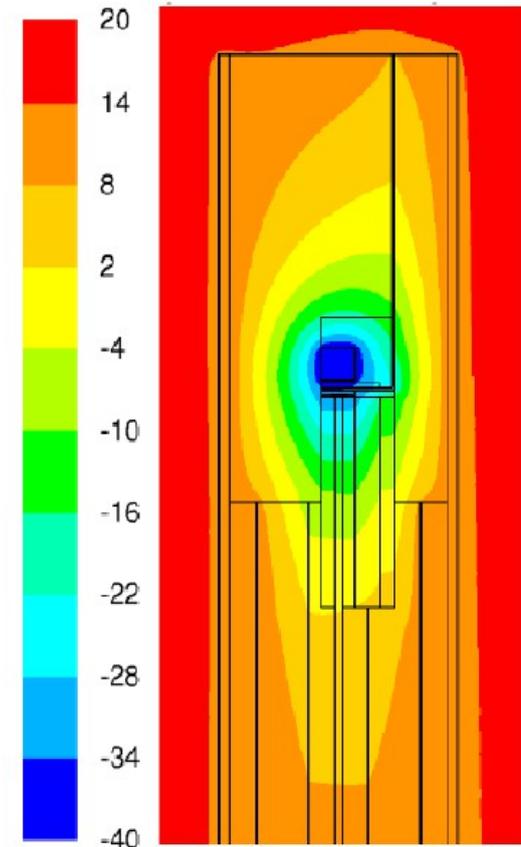
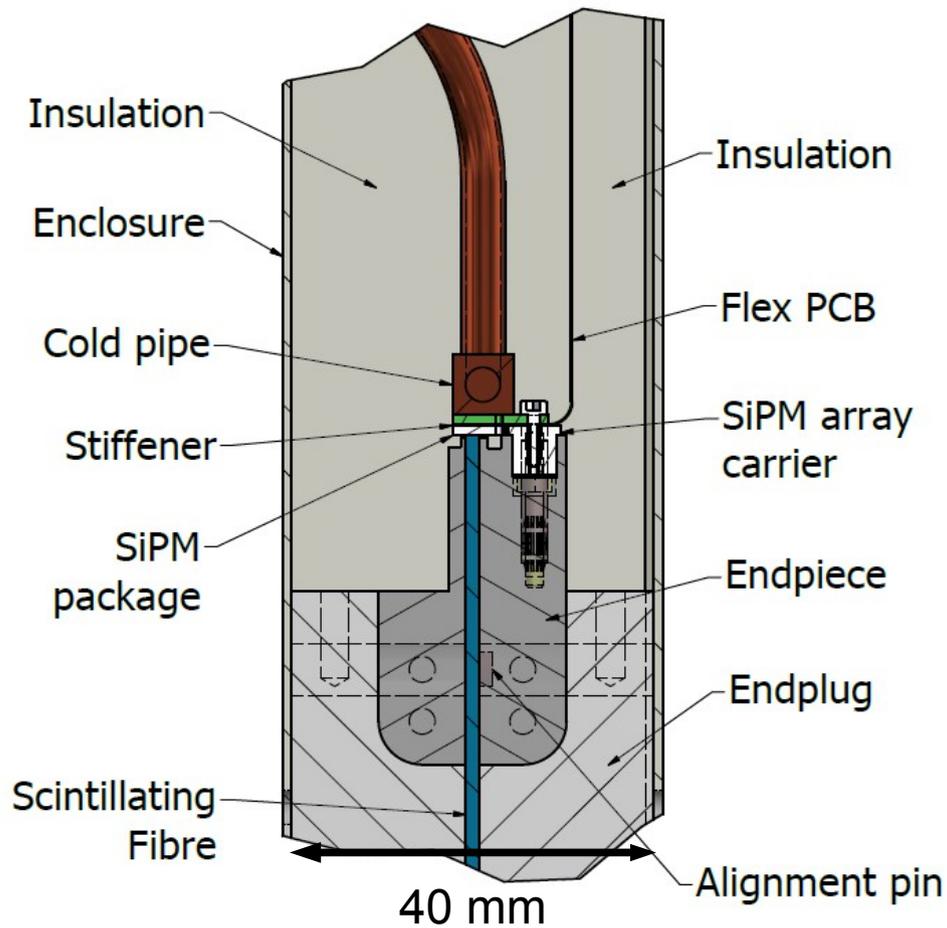
Need to have **clusters** per
 128ch array less than 2
 MHz for efficient tracking =
 no crosstalk!!

Thermal annealing @ +40C is planned during week long stops.

- We want a higher threshold to exclude noise contributions but a low threshold to retain signal, resolution and hit efficiency
- Excessive noise clusters will degrade tracking



Challenges: Cooling



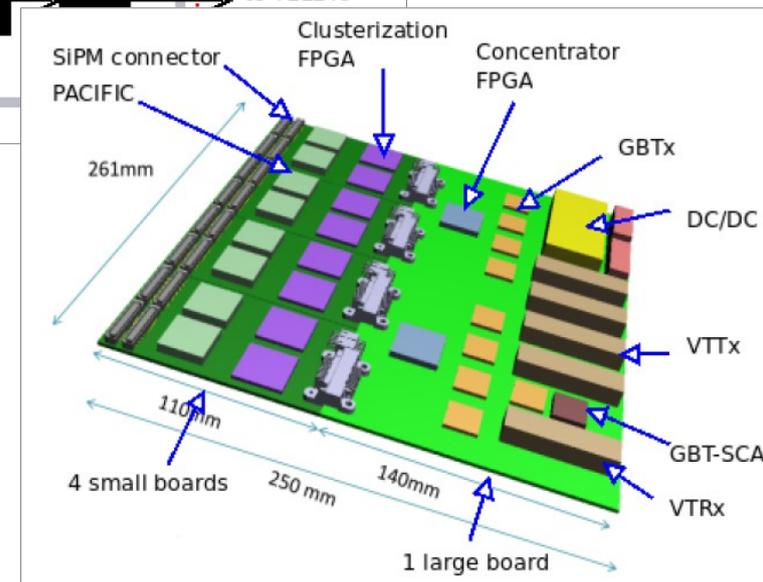
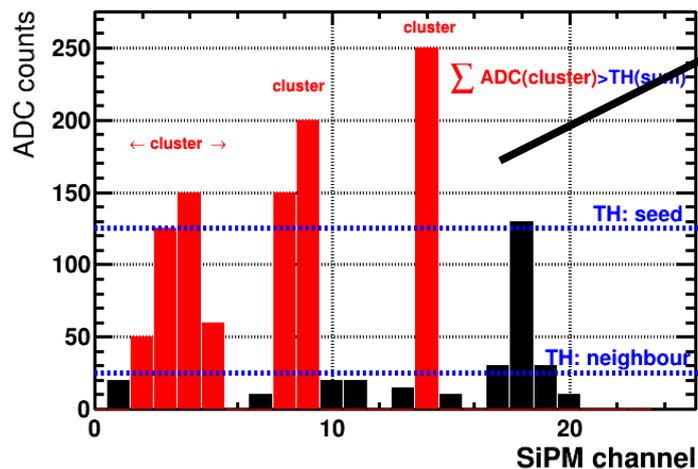
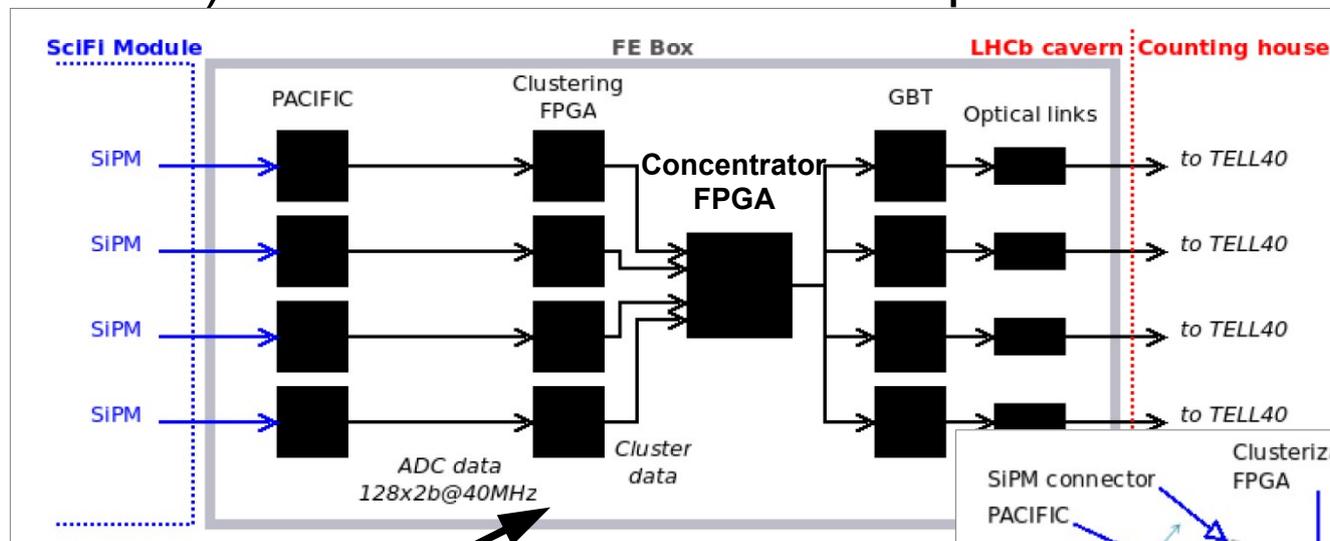
- Acceptable cluster rates require -40C cooling and +40C annealing

$$\text{dark noise} \propto T^2 \exp\left(\frac{-E_g}{2k_B T}\right)$$

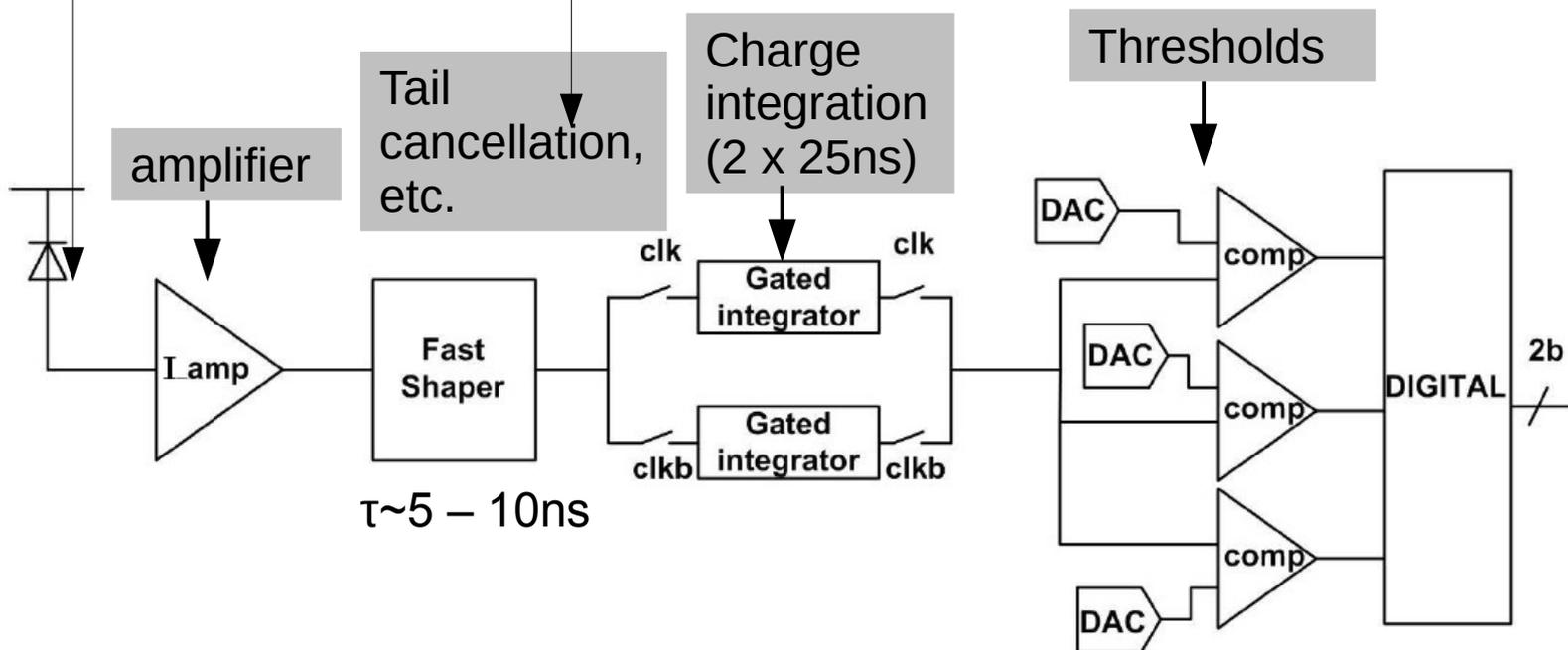
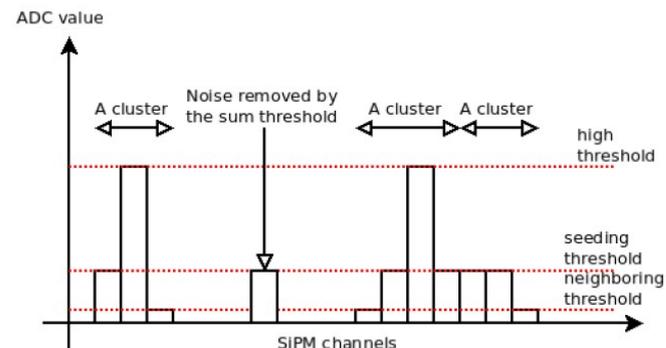
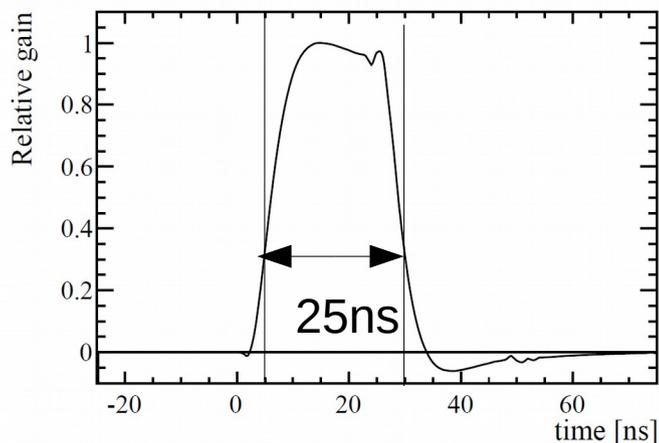
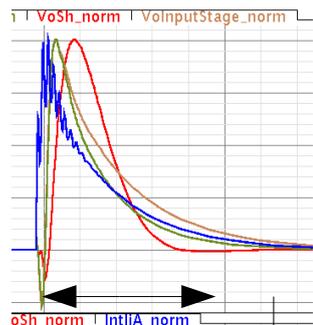
A graph showing the relationship between dark noise (n) and temperature (T(K)). The y-axis is labeled 'n' and the x-axis is labeled 'T(K)'. The curve shows an exponential increase in dark noise as temperature increases, consistent with the equation above.

Challenges: Electronics

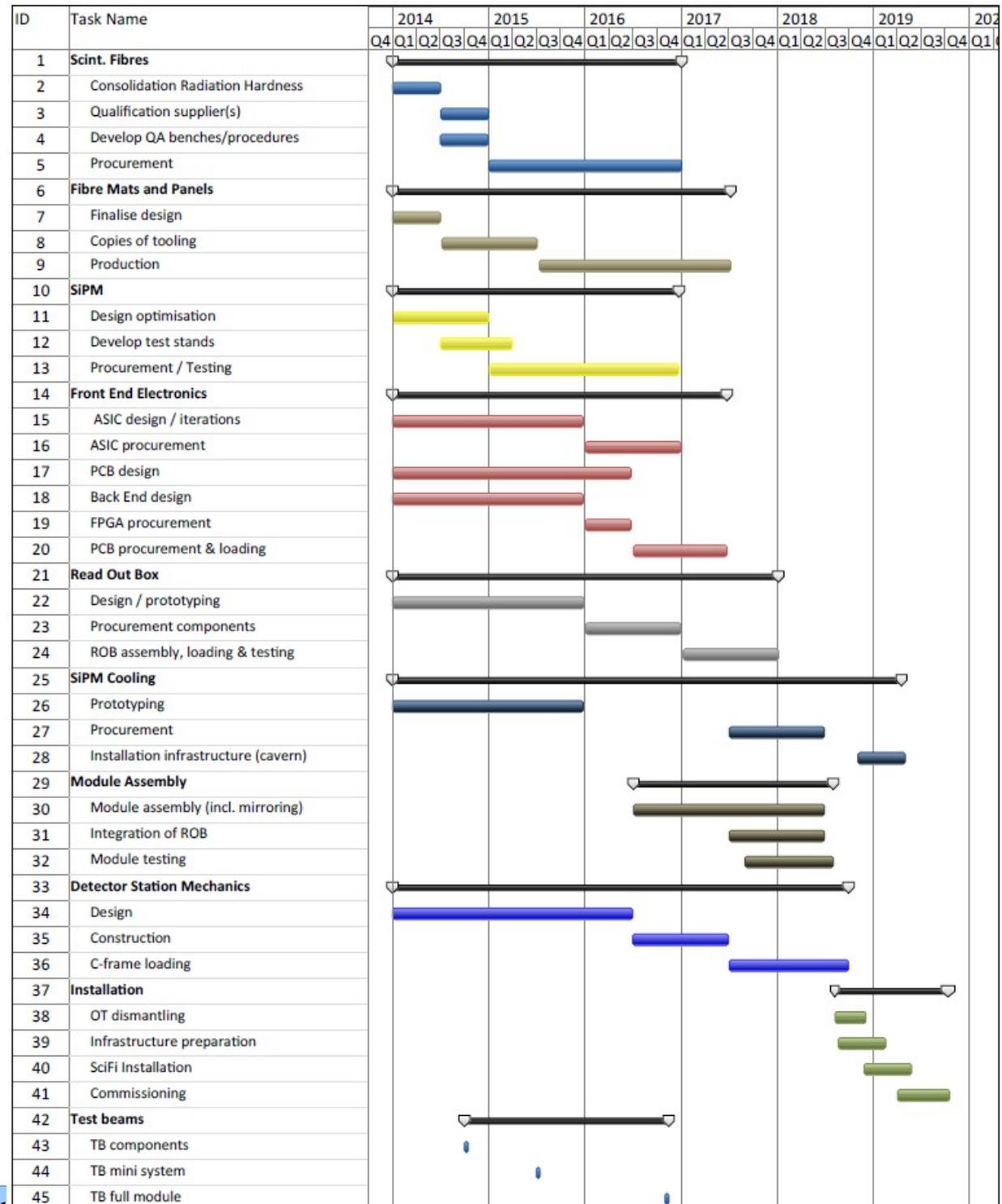
- Digitizes the 560,000 SiPM signals and forms the clusters and hit positions
- ASIC (PACIFIC) and front-end board development



The PACIFIC



- Research is finished.
- Engineering has started.
- SiPM improvements under development yet.
- PACIFIC-2 (8ch at the foundry now)
- Everything must be working and in the LHCb pit mid-2018.



Summary

- The order of magnitude increases in precision will allow new physics searches down to Standard Model theoretical uncertainties
- The SciFi tracker is crucial to scope with the upgrade requirements
- SciFi collaboration with **10 countries in 20 institutions**
- Begin construction in end of 2015; Ready for installation in 2018

backup

