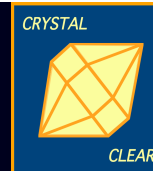


Inorganic Scintillators - An overview

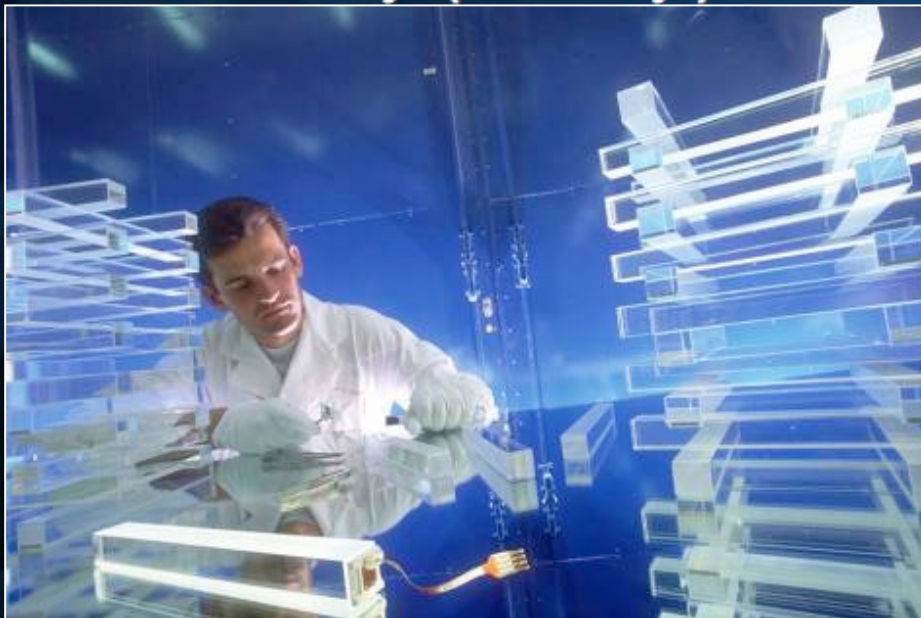
Paul Lecoq
CERN, Geneva



HEP and PET use the largest volume of scintillators



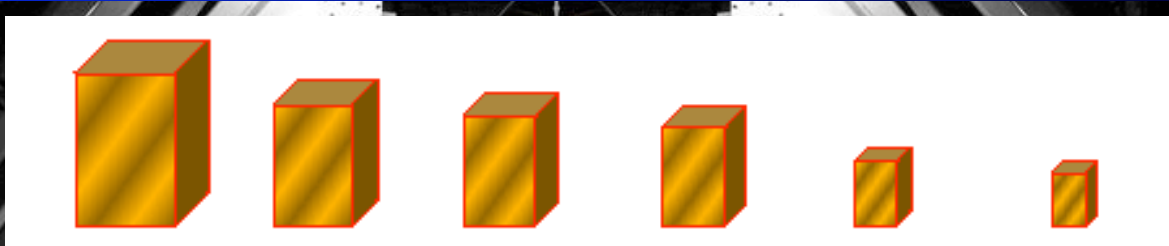
High energy physics
(e.g. CMS)
80,000 crystals; 12,000
liters; highest production rate in 2005
4100 liters/yr (34 tons/yr)



Positron Emission Tomography
in 2003, 450 sc/yr x 10 liters/sc
= 4500 liters/yr (33 tons/yr)



Some popular crystals in HEP



	NaI(Tl)	BaF ₂	CsI(Tl)	CeF ₃	BGO Bi ₄ Ge ₃ O ₁₂	PWO PbWO ₄
Xo [cm]	2.59 😞	2.03 😞	1.86 😊	1.66 😊	1.12 😊	0.92 😊
ρ [g/cm ³]	3.67 😞	4.89 😞	4.53 😞	6.16 😊	7.13 😊	8.2 😊
τ [ns]	230 😞	0.6 😊 620 😞	1050 😞	30 😊	340 😊	15 😊
λ [nm]	415 😊	230 😊 310 😊	550 😊	310 😊 340 😊	480 😊	420 😊
n@λ _{max}	1.85 😊	1.56 😊	1.80 😊	1.68 😊	2.15 😞	2.3 😞
LY [%NaI]	100 😊	5 😞 16 😞	85 😊	5 😊	10 😊	0.5 😞



Scintillators for PET

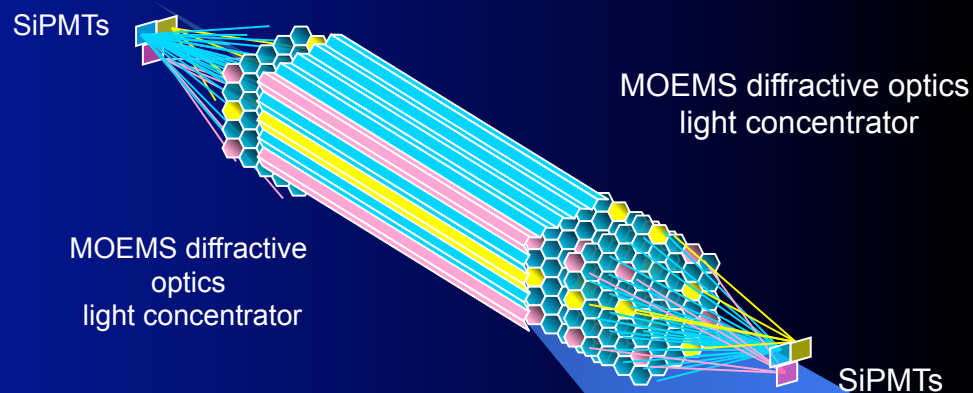


Scintillators for PET

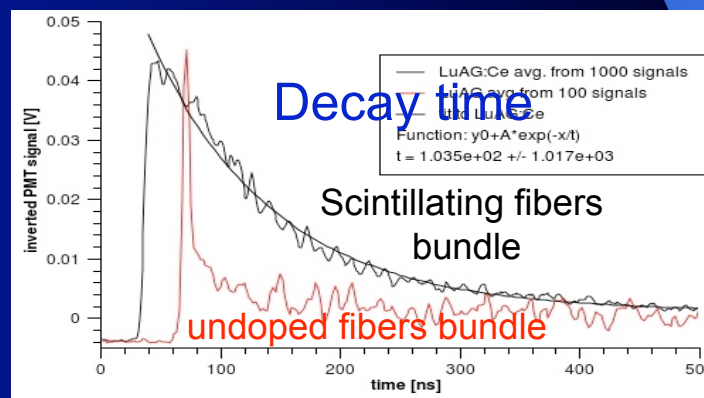
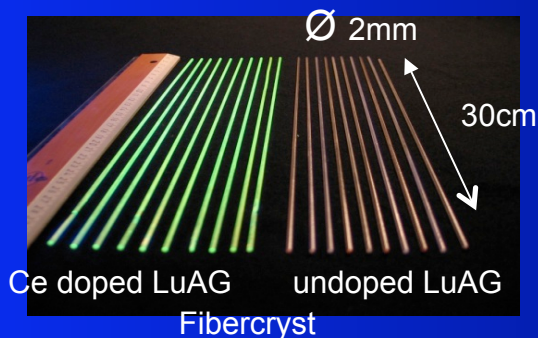
	1962	1977	1995	1999	2001	2003	2007
	NaI	BGO	GSO:Ce	LSO:Ce	LuAP:Ce	LaBr ₃ :Ce	LuAG:Ce
Density (g/cm ³)	3.67	7.13	6.71	7.40	8.34	5.29	6.73
Atomic number	51	75	59	66	65	47	63
Photofraction	0.17	0.35	0.25	0.32	0.30	0.13	0.30
Decay time (ns)	230	300	30-60	35-45	17	18	60
Light output (hv/MeV)	43000	8200	12500	27000	11400	70000	>25000
Peak emission (nm)	415	480		420			
Refraction index	1.85	2.15	430	1.82	365	356	535
			1.85		1.97	1.88	1.84

Dual readout based on Metamaterials

- Select a non-intrinsic scintillating material (unlike BGO or PWO) with high bandgap for low UV absorption
- The undoped host will behave as an efficient Cerenkov: heavy material, high refraction index n , high UV transmission
- Cerium or Praesodinium doped host will act as an efficient and fast scintillator
 - $\approx 60\text{ns}$ decay for Ce
 - $\approx 20\text{ns}$ decay for Pr



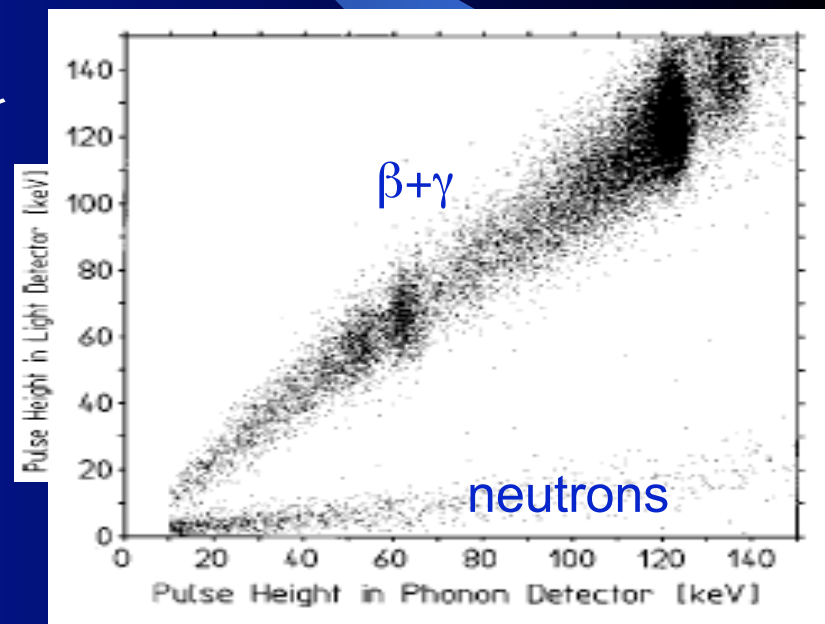
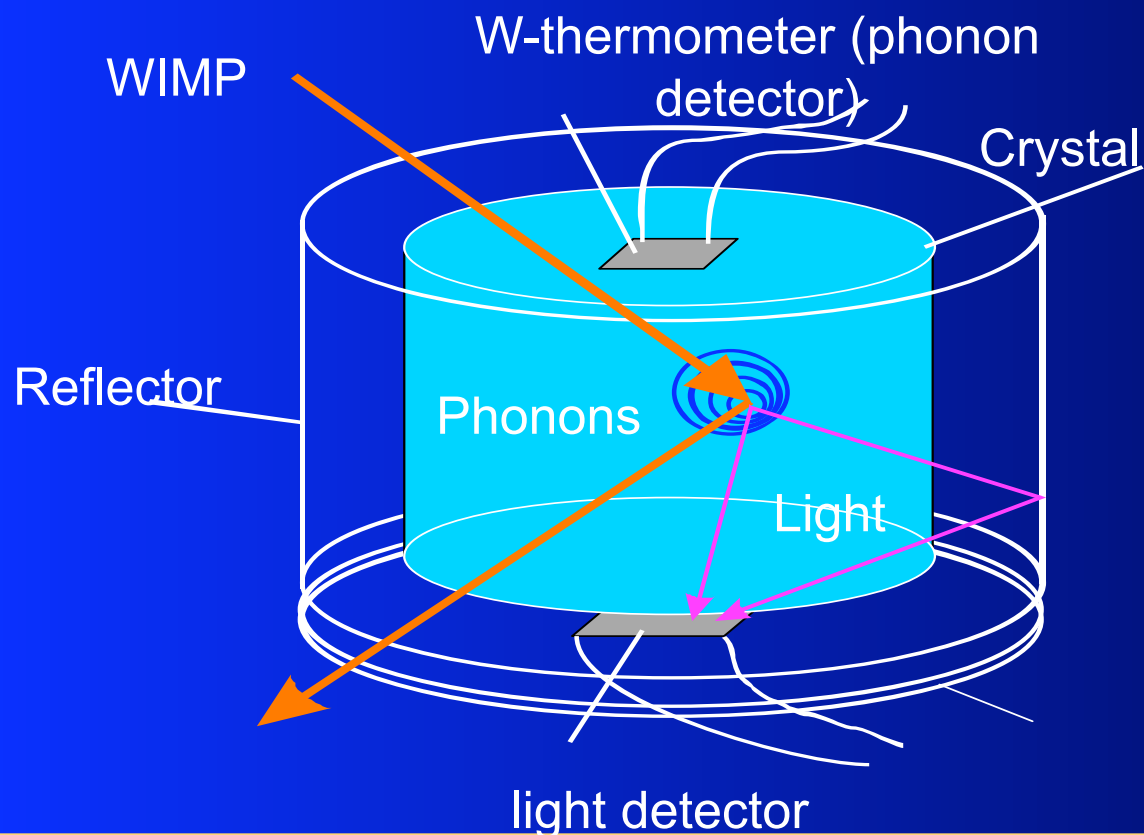
LuAG is an excellent candidate



Dark matter search

- **Sensitivity + Selectivity**

high-energy resolution of cryogenic phonon detector + discrimination of events with low detection threshold (≤ 10 keV)



P.Meunier et al, Appl.Phys.Let 75 (1999) 1335



Scintillator requirements



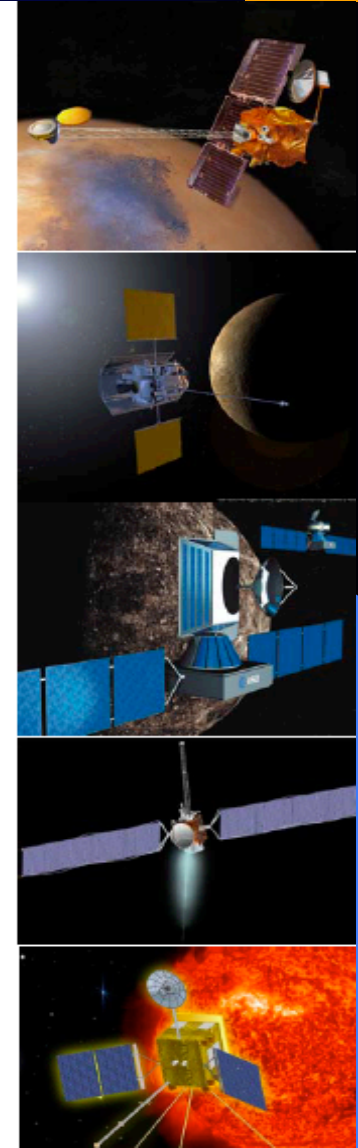
- high light yield at low temperatures
 - ✓ low threshold,
 - ✓ good energy resolution
- Radiopurity (ex Lu, Rb, K, U, Th)
- Possible candidates
 - ✓ CaWO_4 – satisfactory choice, currently in use, large ongoing effort to improve the material
 - ✓ ZnWO_4 – scintillator under development for cryogenic application
 - ✓ CaMoO_4 and CdMoO_4 – material under investigation

Scintillation Detectors on Space Missions



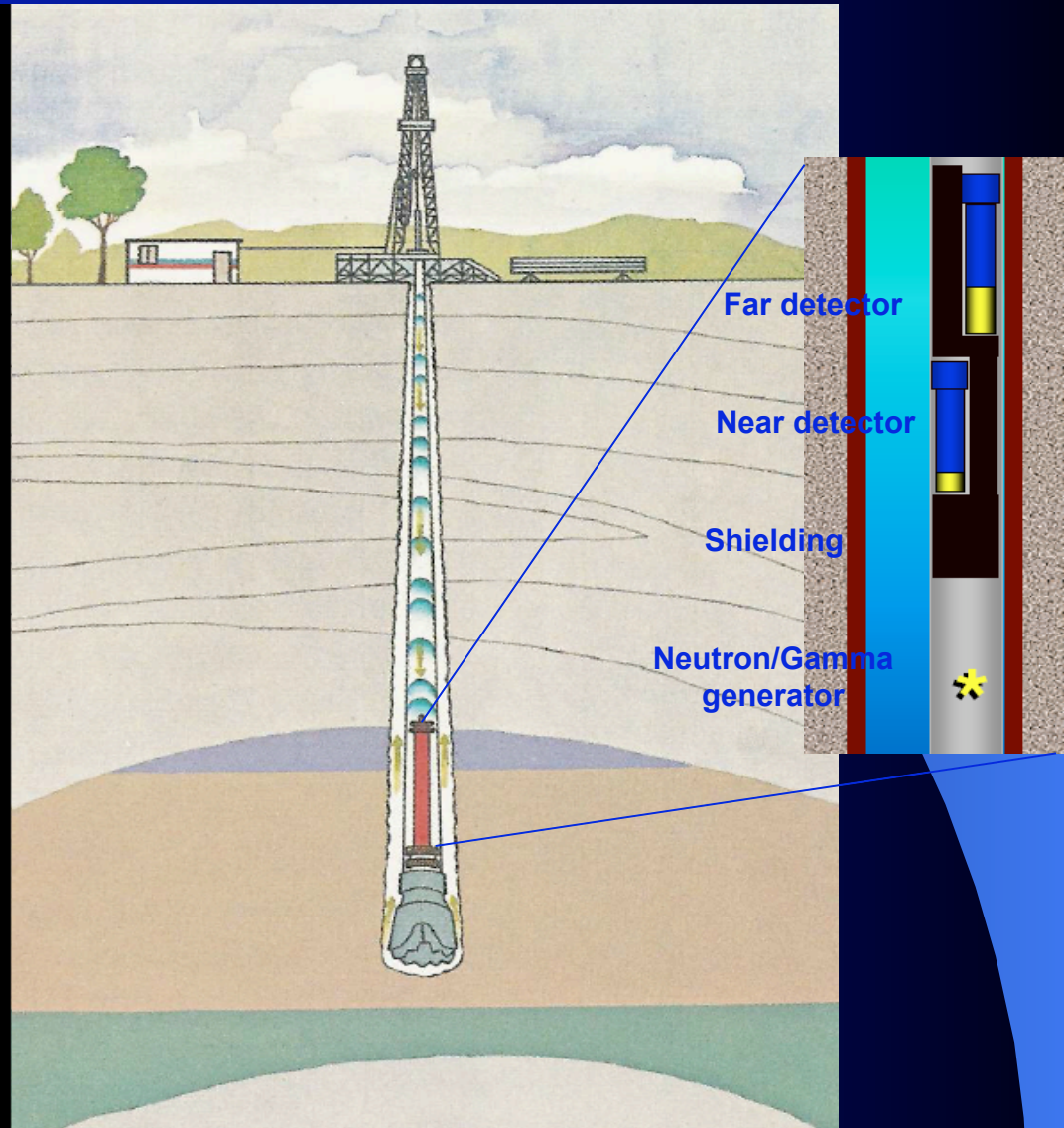
	GRS	NS	PS
• Past missions			
– Phobos	CsI	Stilben, plastic	yes
– Lunar Prospector	BGO/BC454	BC454	
– Near	NaI/BGO		
– Mars Observer	HPGe/BC454	BC454	
– Mars Odyssey	HPGe	BC454, Stilben, plastic	CsI
• Current missions			
– Ulysses	CsI/GRB		Plastic
– Messenger	HPGe/BGO	GS20, BC454	
• Missions in implementation			
– Dawn	CZT/BGO	BGO, BC454, G20	
– Phobos Grunt	LaBr ₃	BC454, Stilben, plastic	
– Solar Orbiter	LaBr ₃		
– BepiColombo	LaBr ₃		CsI

GRS=gamma-ray spectrometer
NS=neutron spectrometer
PS=particle spectrometer



Measurement Issues

- Source and sensor both in borehole
- Usually want to measure Formation
- Need to make measurement with 1-3 seconds of data





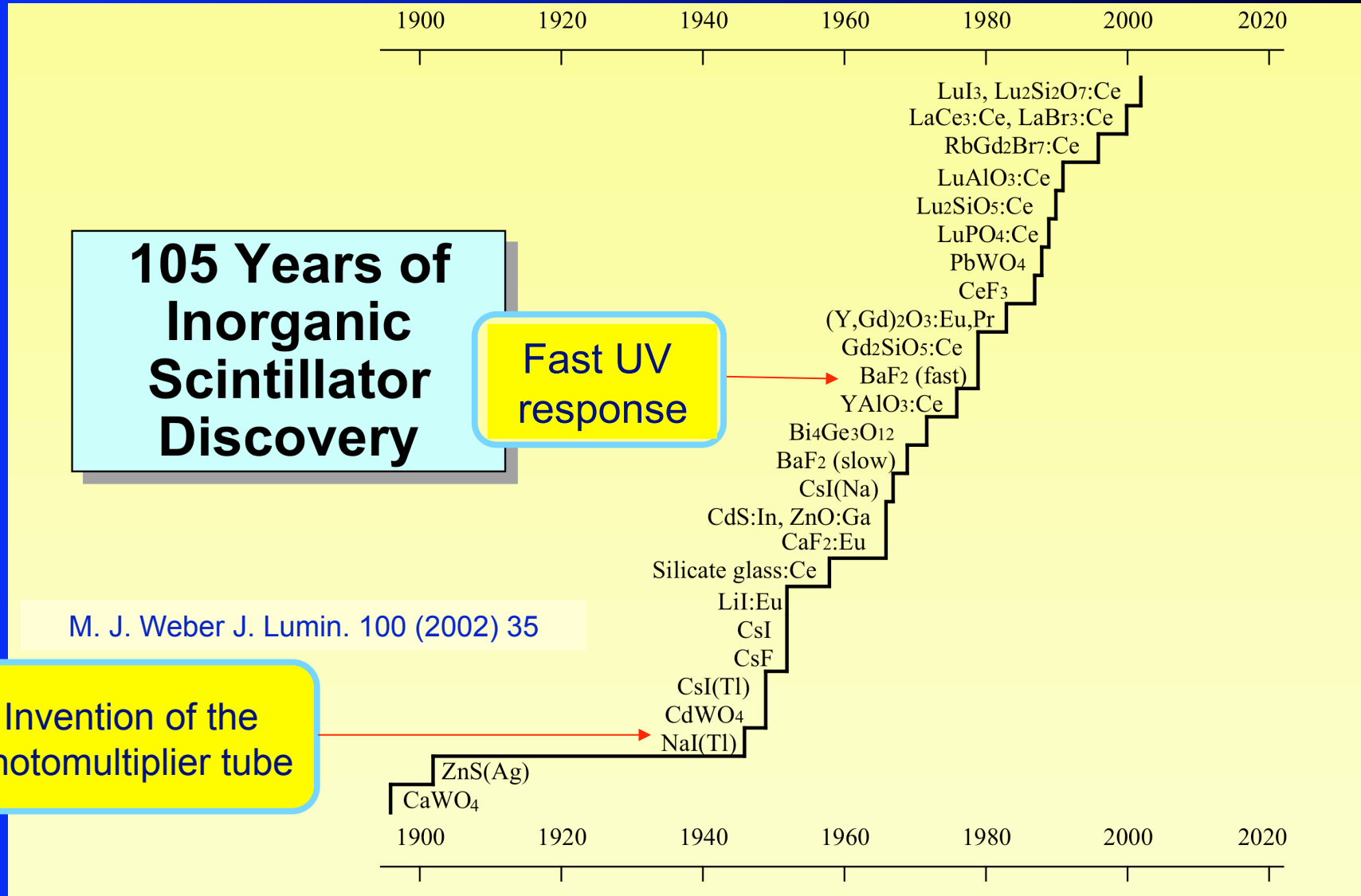
**Mobile and fixed position; X ray, ^{60}Co , ^{137}Cs
 NaI , CdWO_4 , BGO
Spectrometers, counters, imagers**

8 Managed by UT-Battelle
for the Department of Energy

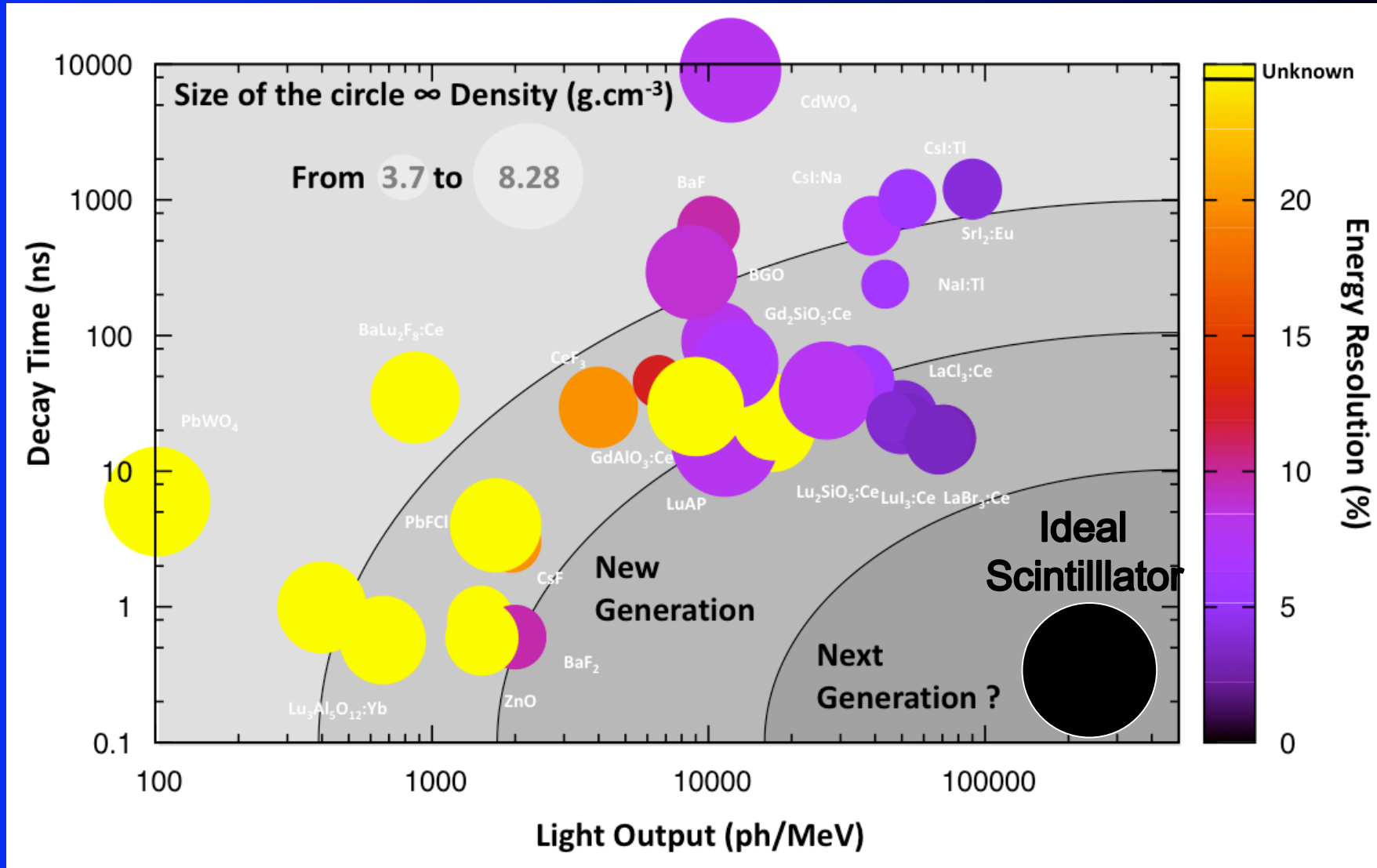
Scintillator Applications in Homeland Security



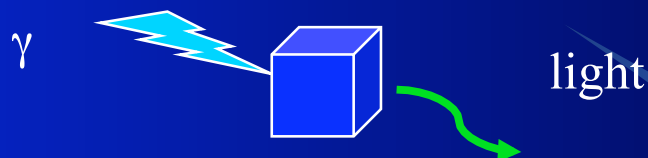
History of scintillator discovery



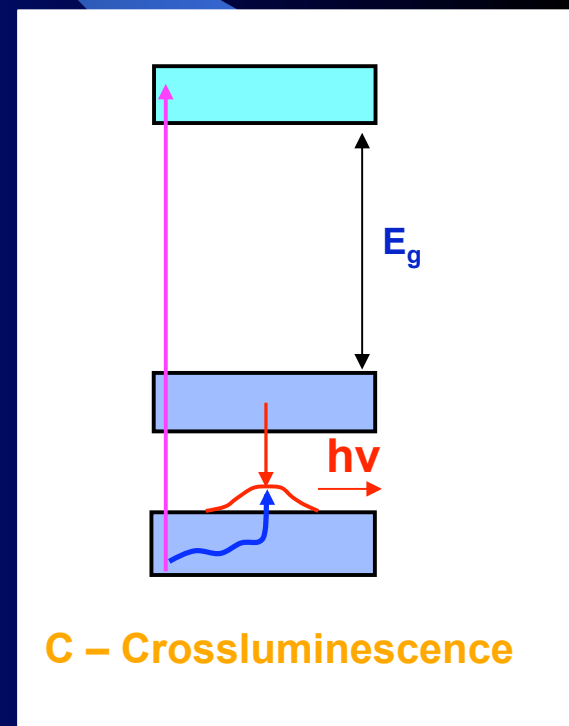
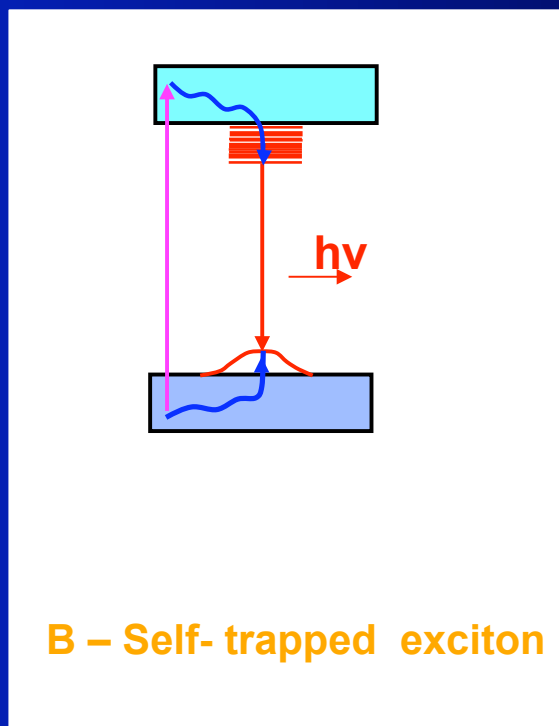
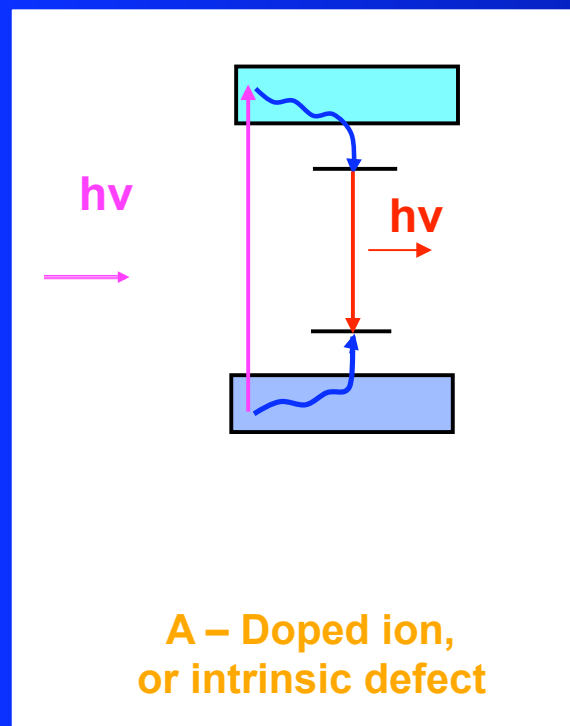
Classification of scintillators



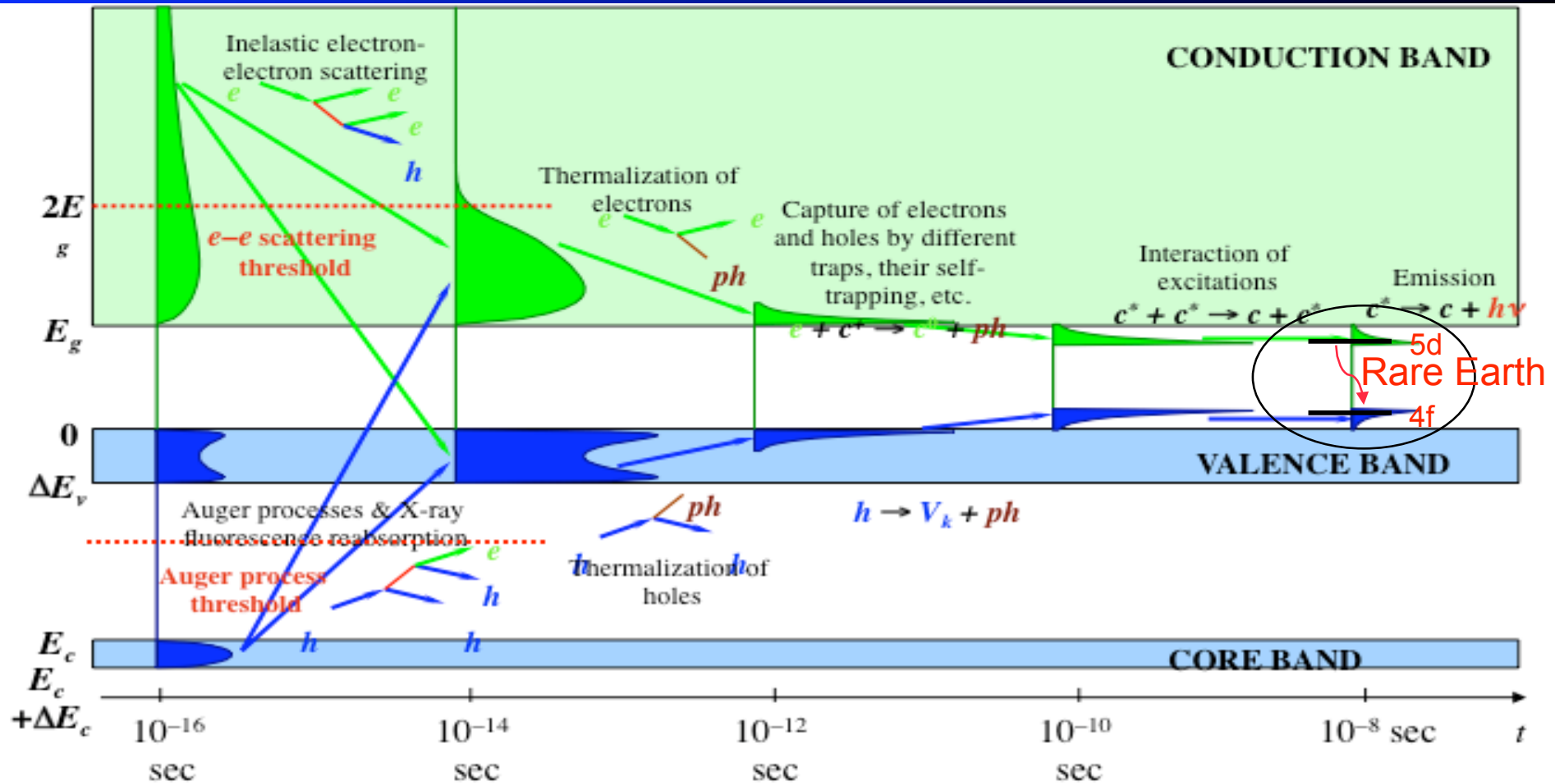
Fundamental aspects of Scintillation



Different scintillation mechanisms

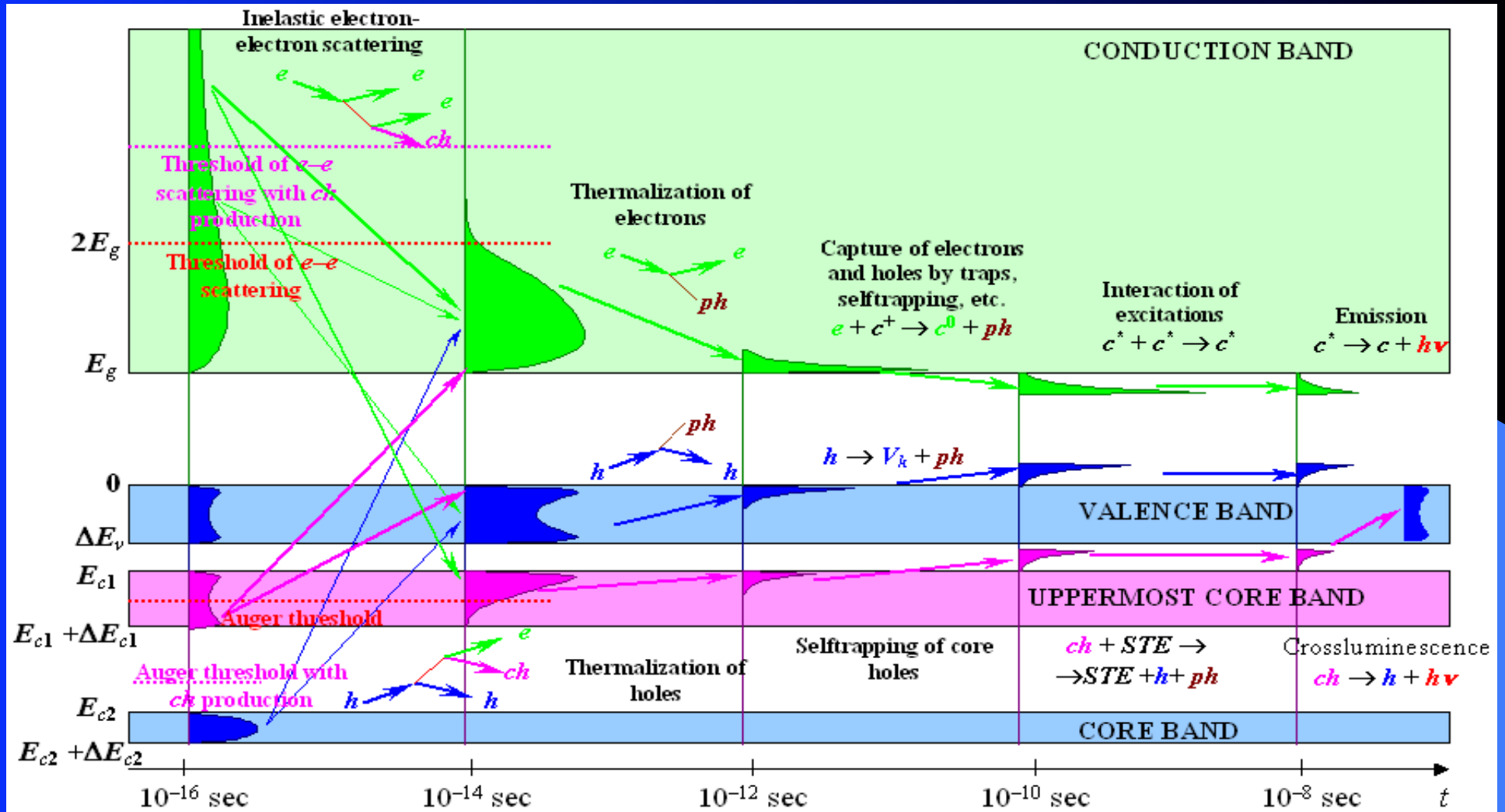


Relaxation of electronic excitations *intrinsic luminescence*

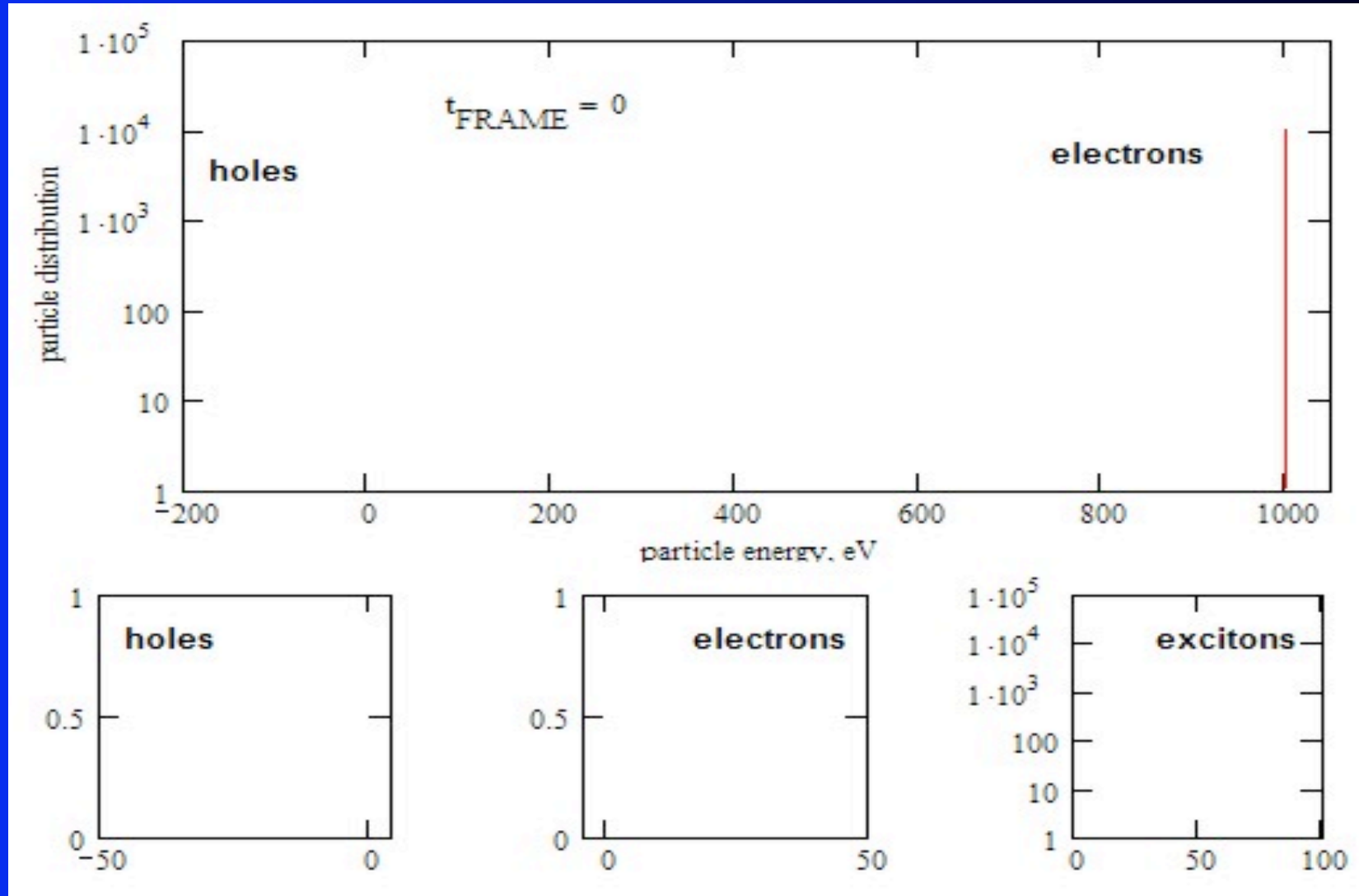


A. Vasil'ev, SCINT2001 proceedings, NIMA 486 (2002) 367

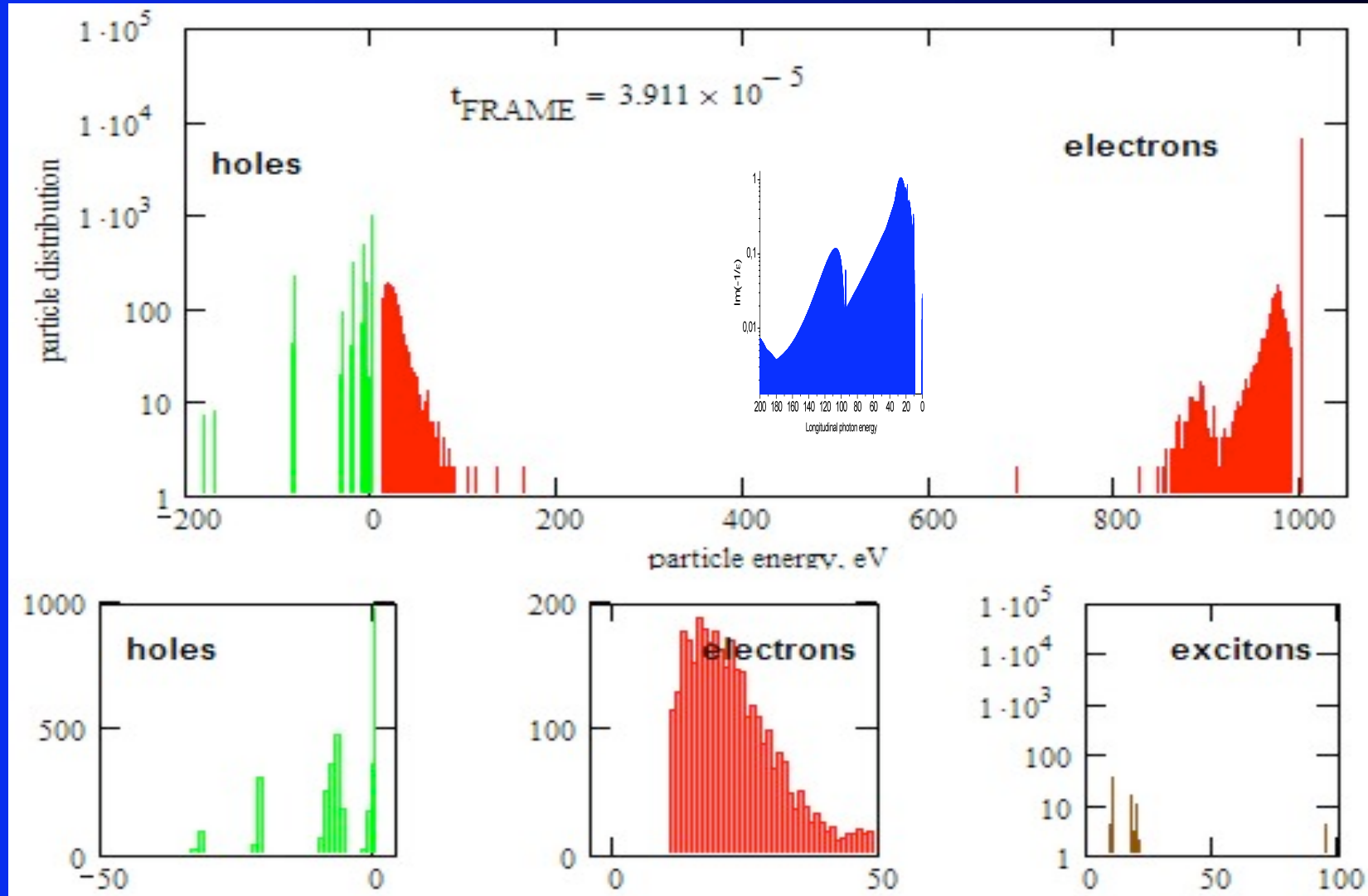
Relaxation of electronic excitations Cross-luminescence



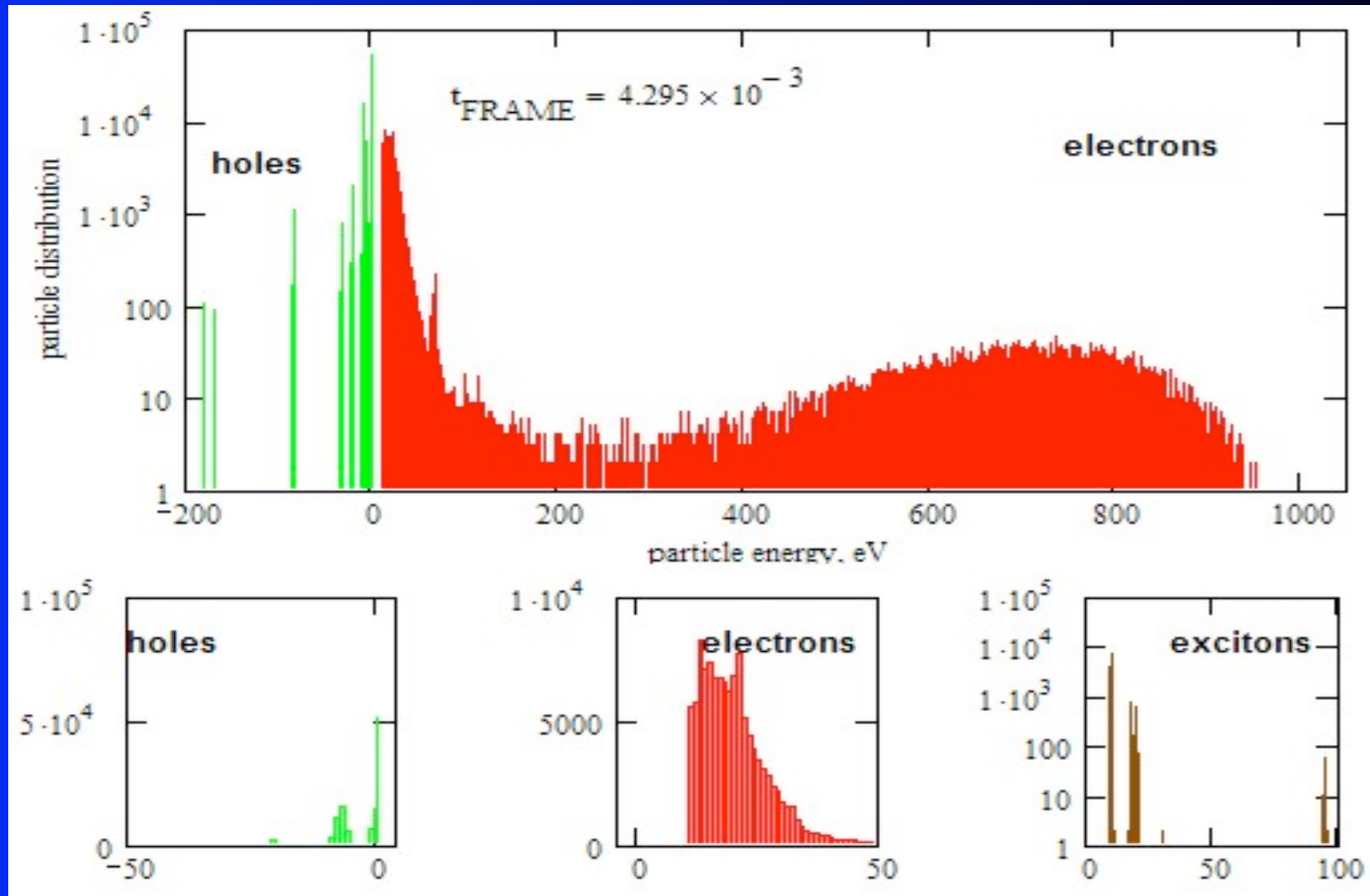
Evolution of energy distribution for 1000 eV electrons



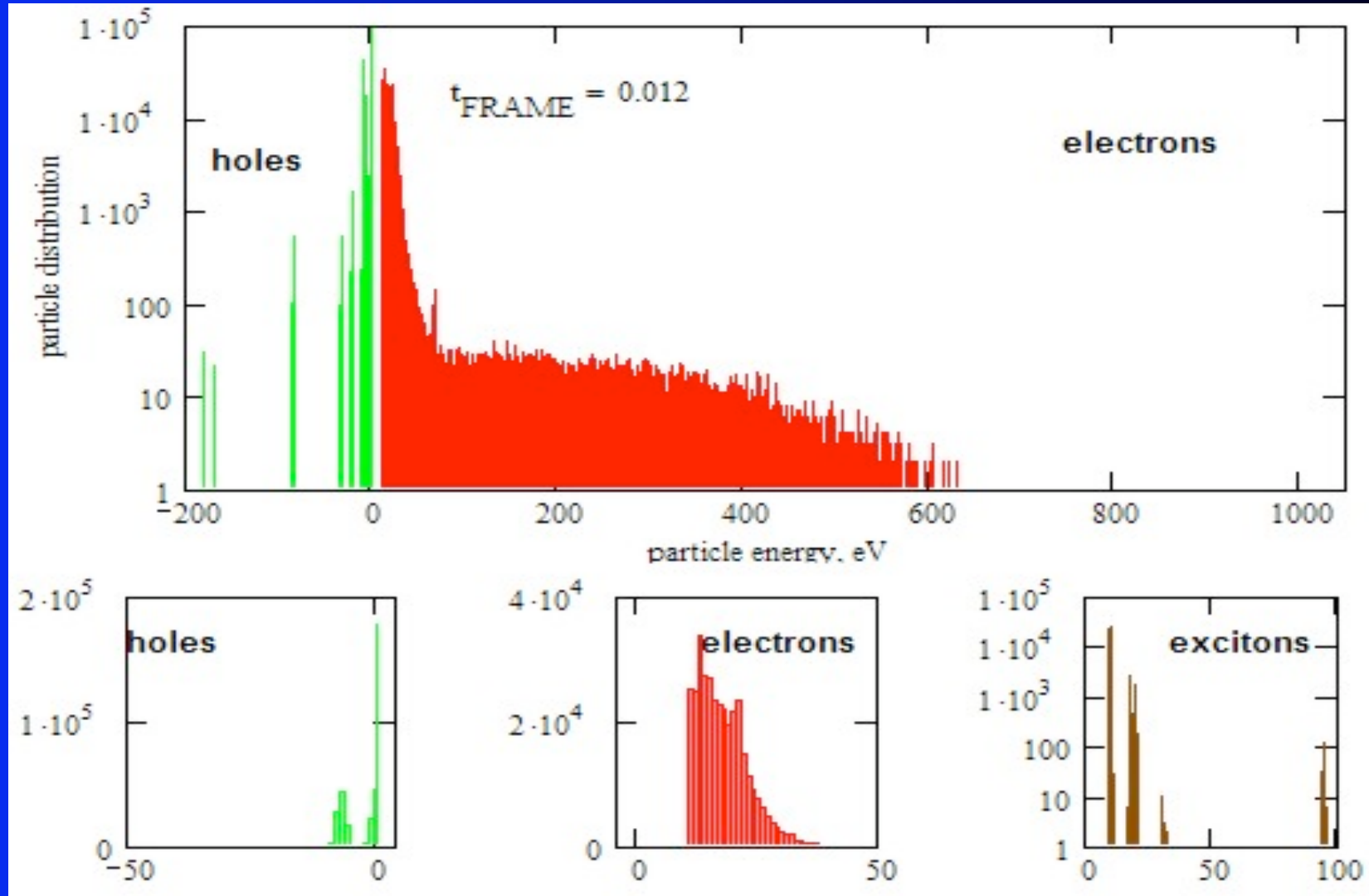
Evolution of energy distribution for 1000 eV electrons



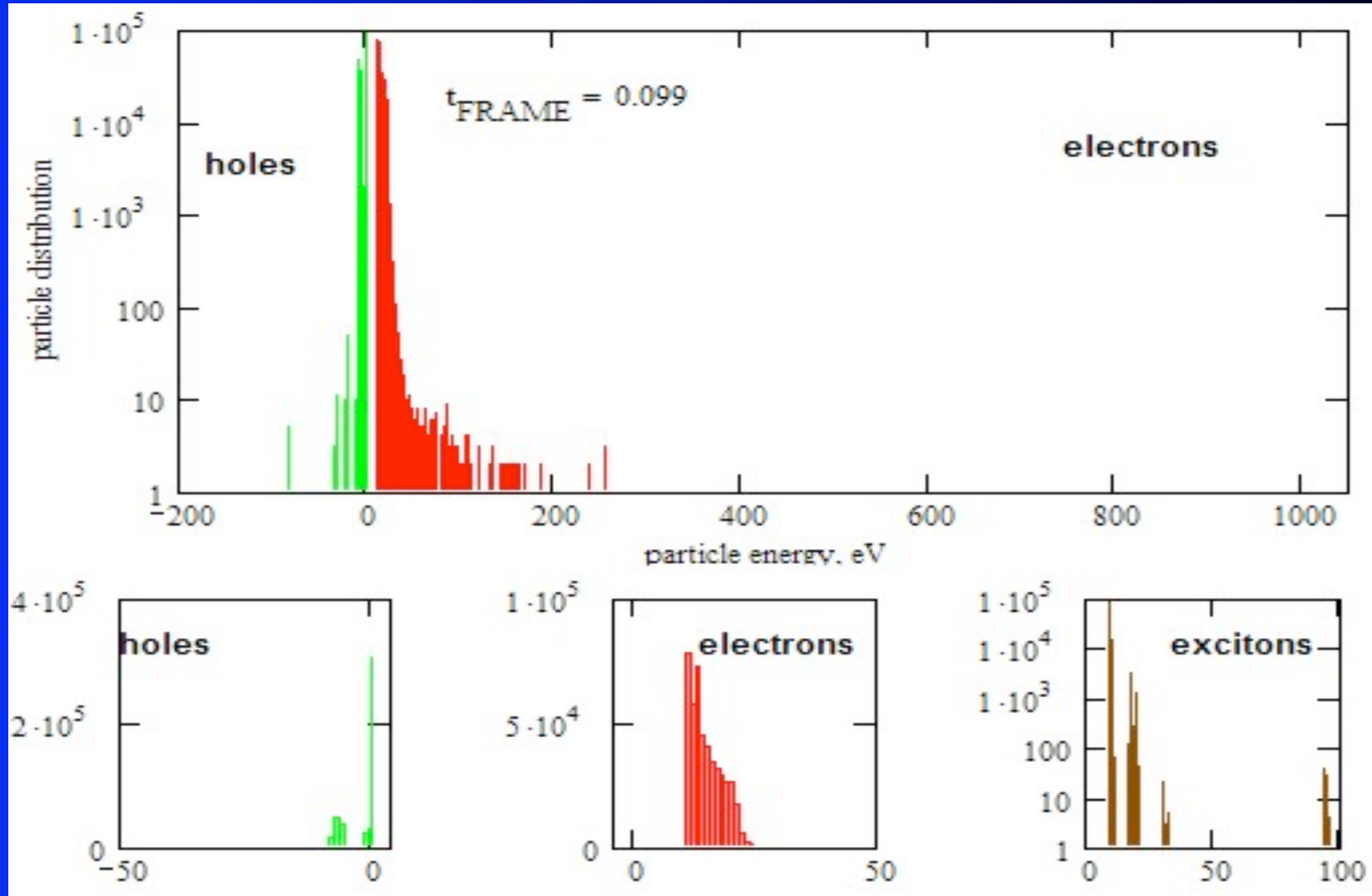
Evolution of energy distribution for 1000 eV electrons

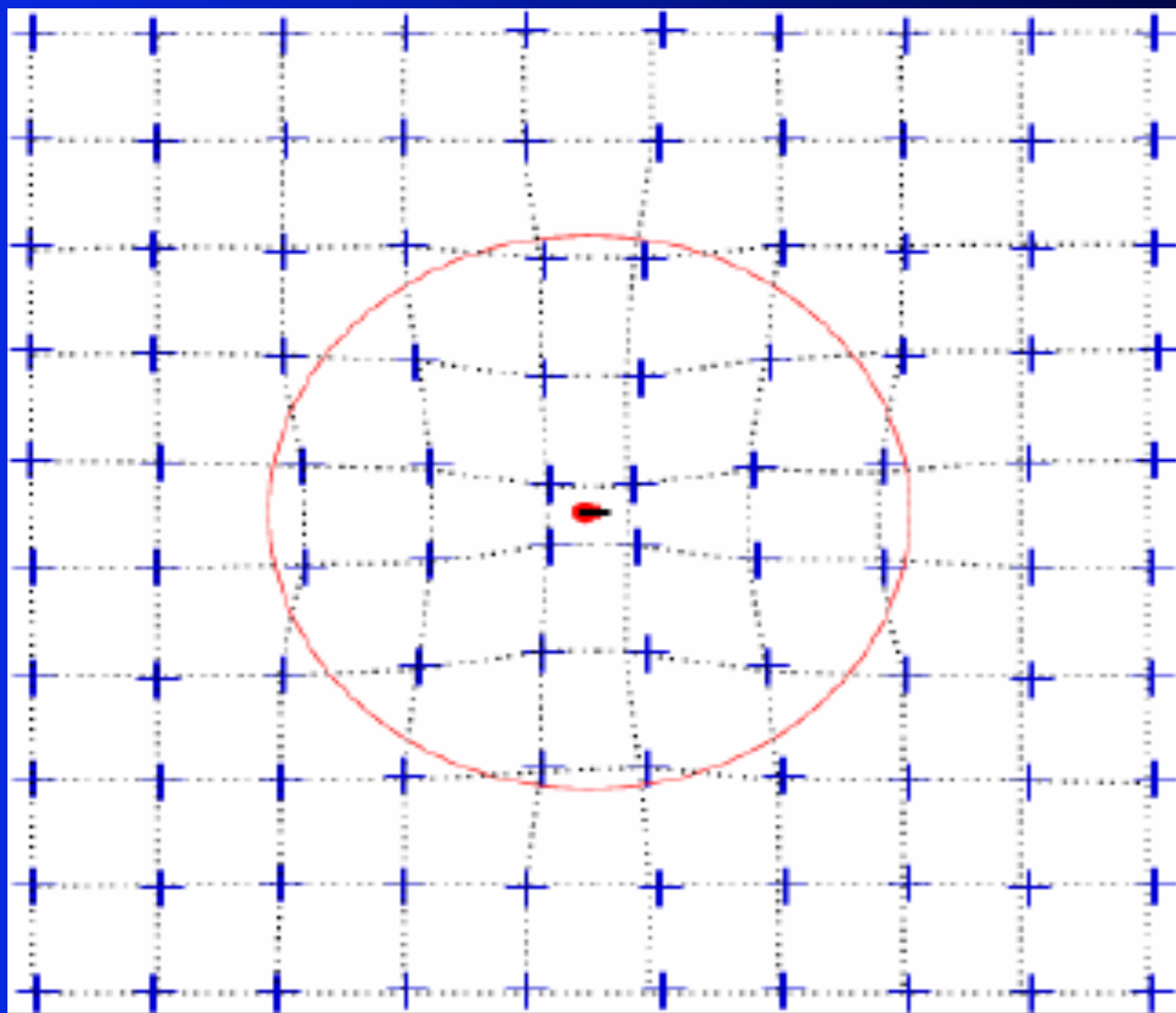


Evolution of energy distribution for 1000 eV electrons



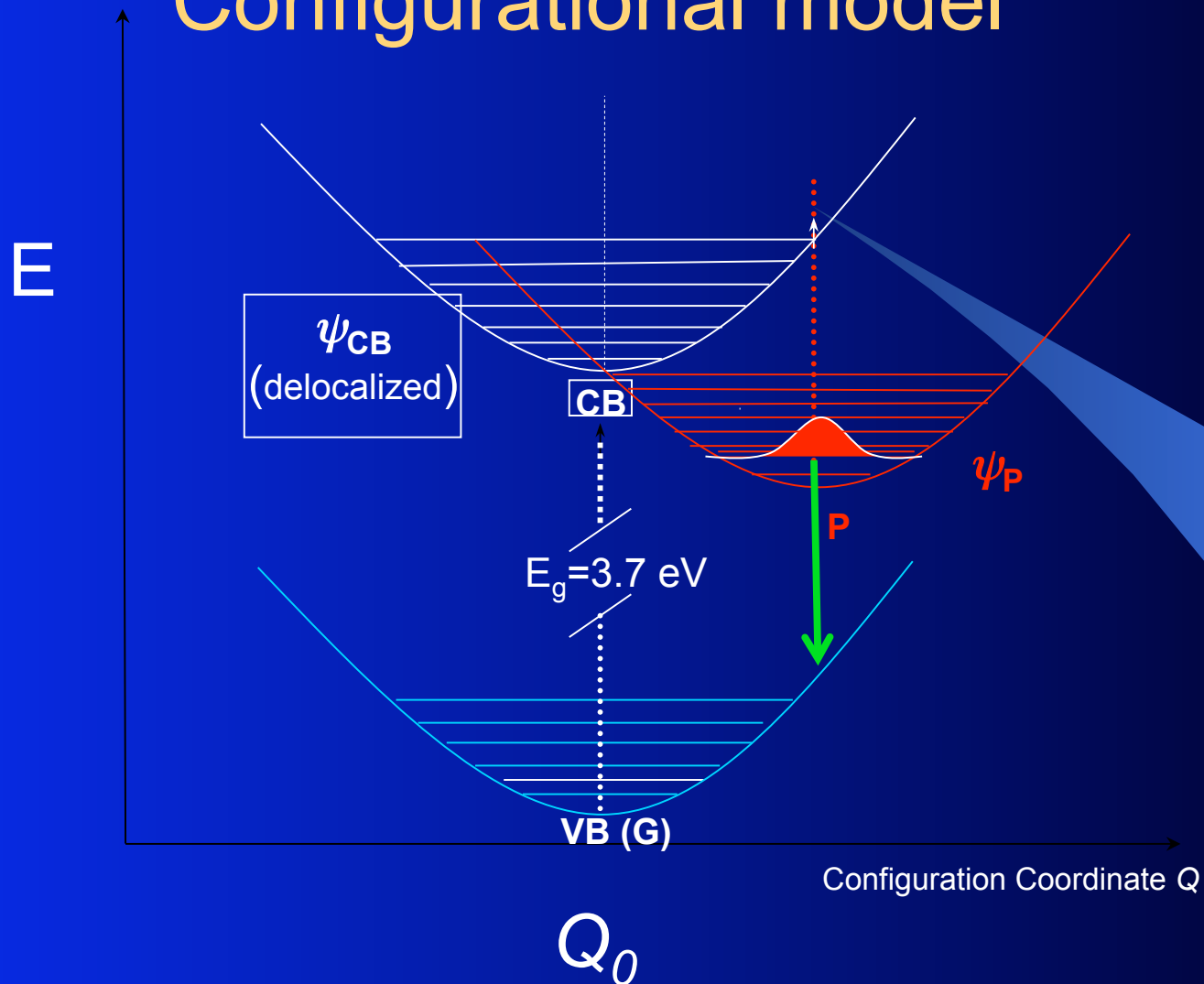
Evolution of energy distribution for 1000 eV electrons





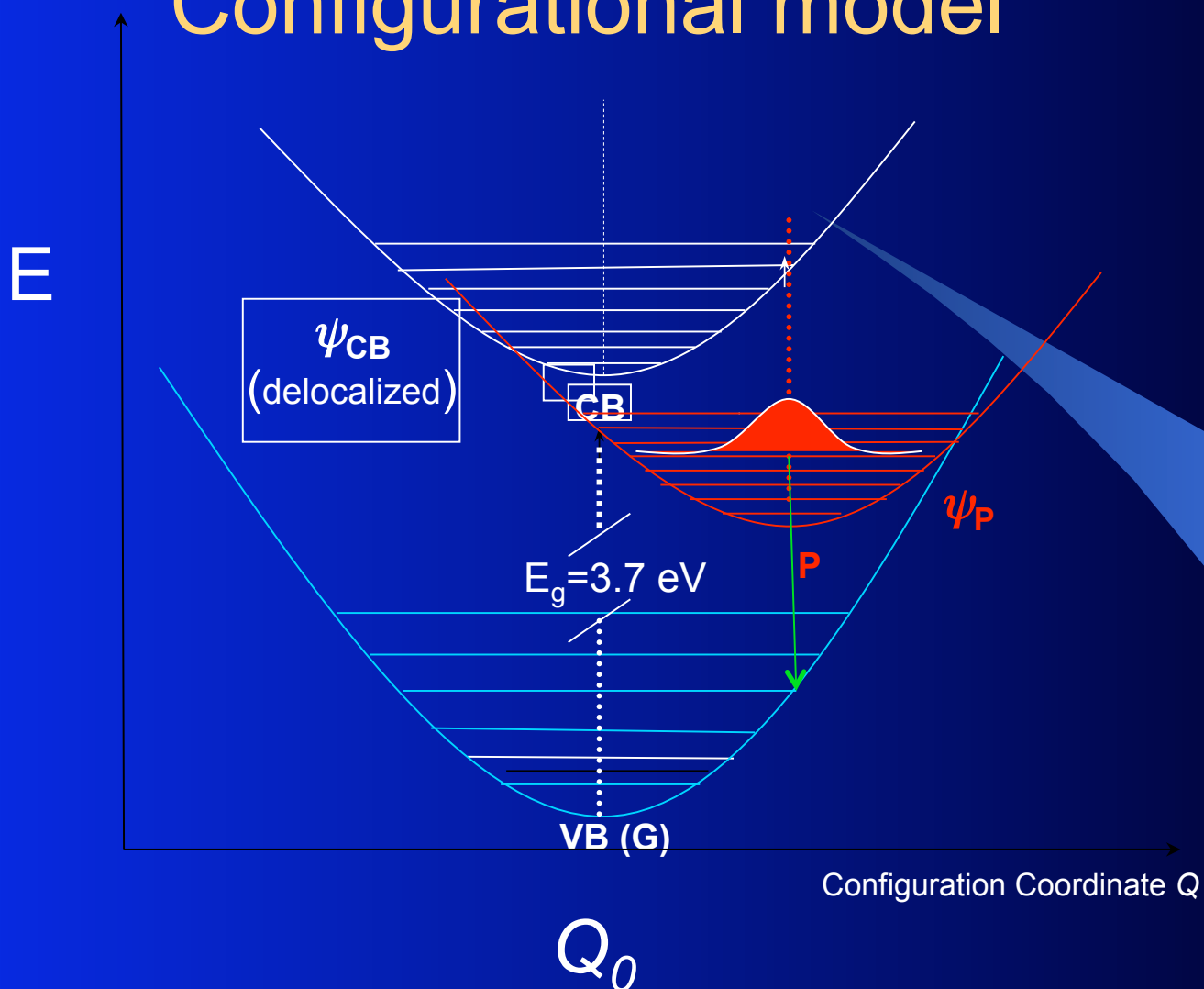
polaron - electron + distorted lattice

Configurational model



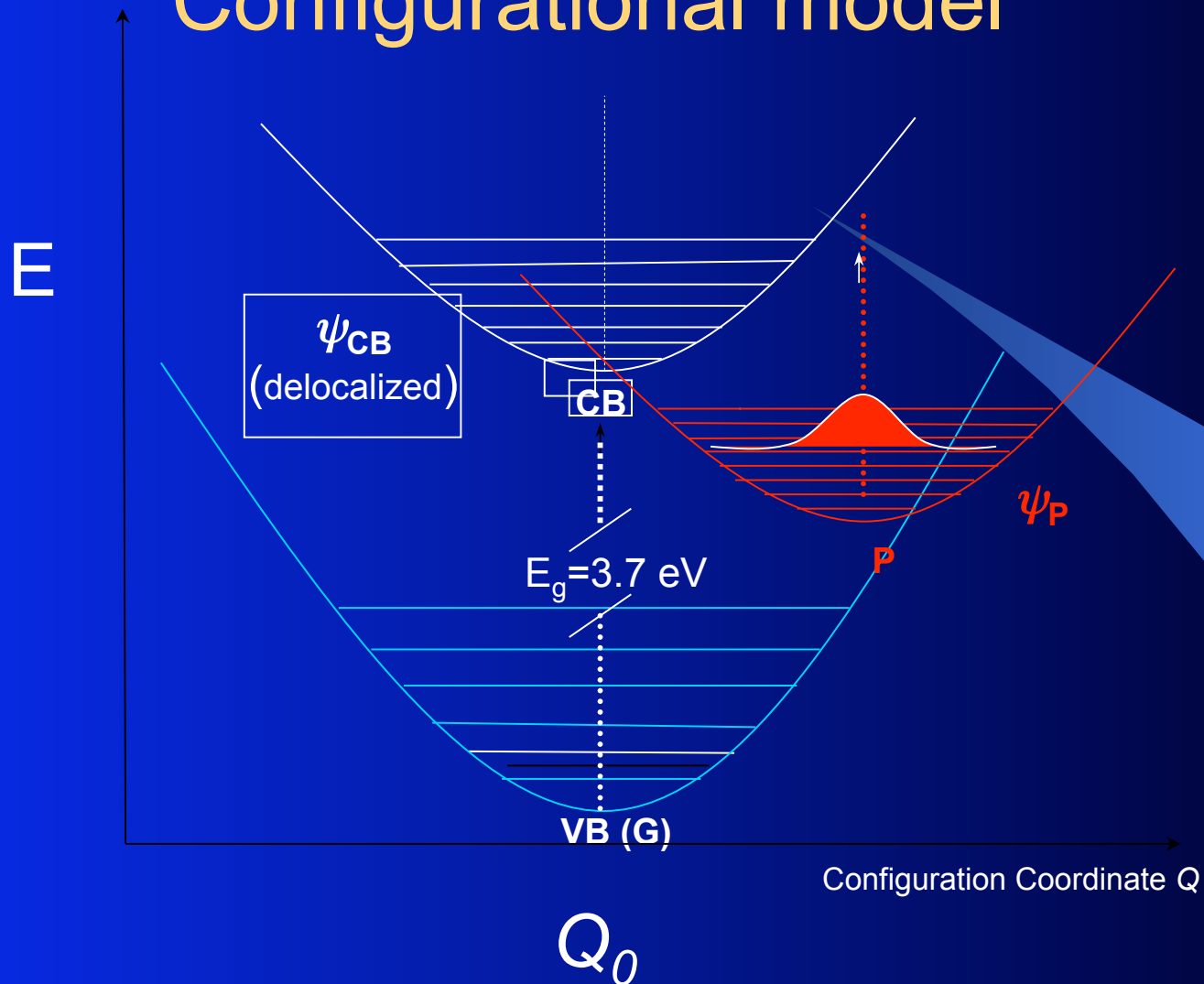
Configuration coordinate model for the local lattice with electron in **valence** and **conduction band** states and in **localized polaron** state.

Configurational model



Configuration coordinate model for the local lattice with electron in **valence** and **conduction band** states and in **localized polaron** state.

Configurational model



Configuration coordinate model for the local lattice with electron in **valence** and **conduction band** states and in **localized polaron** state.

Fundamental aspects of Scintillation

The 3 phases of the scintillation mechanism

1. Absorption : Creation of pair e-h

$$n_{e-h} = \frac{E_\gamma}{\beta' E_{gap}}$$

2. Transfer to the luminescence centre

Efficiency of energy transfer : **S**

3. Emission

Efficiency of emission : **q**



Efficiency of scintillation

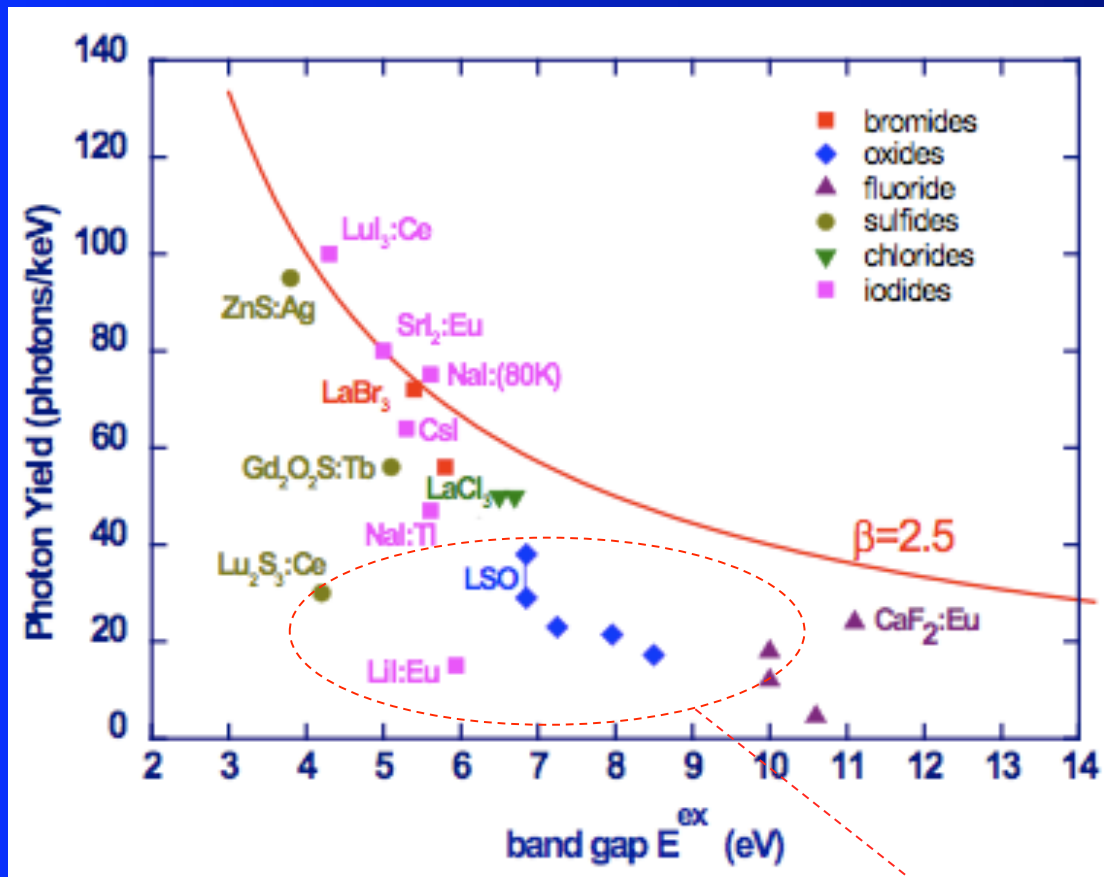
$$n_{photon} = n_{eh} S q = \frac{E_\gamma}{\beta' E_g} S q$$

Determination of the maximum of light

$$LY_{max} = \frac{n_{photon}}{E_\gamma} = \frac{1}{\beta E_g}$$

Usually $\beta = 2$ to 4

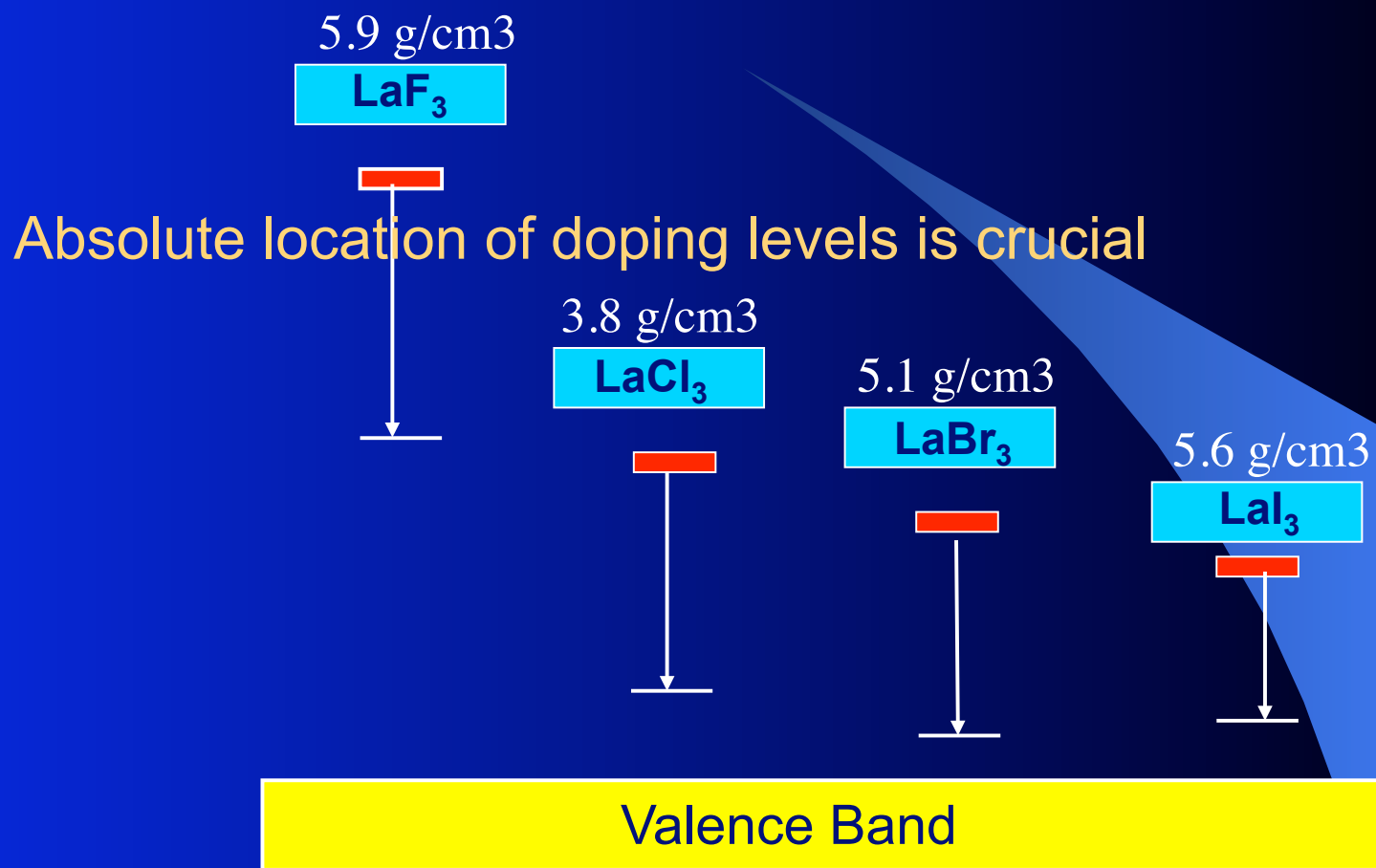
Fundamental limits to the LY



$$N_{ph} \leq N_{eh} = \frac{E_{\gamma}}{\beta E_{gap}}$$

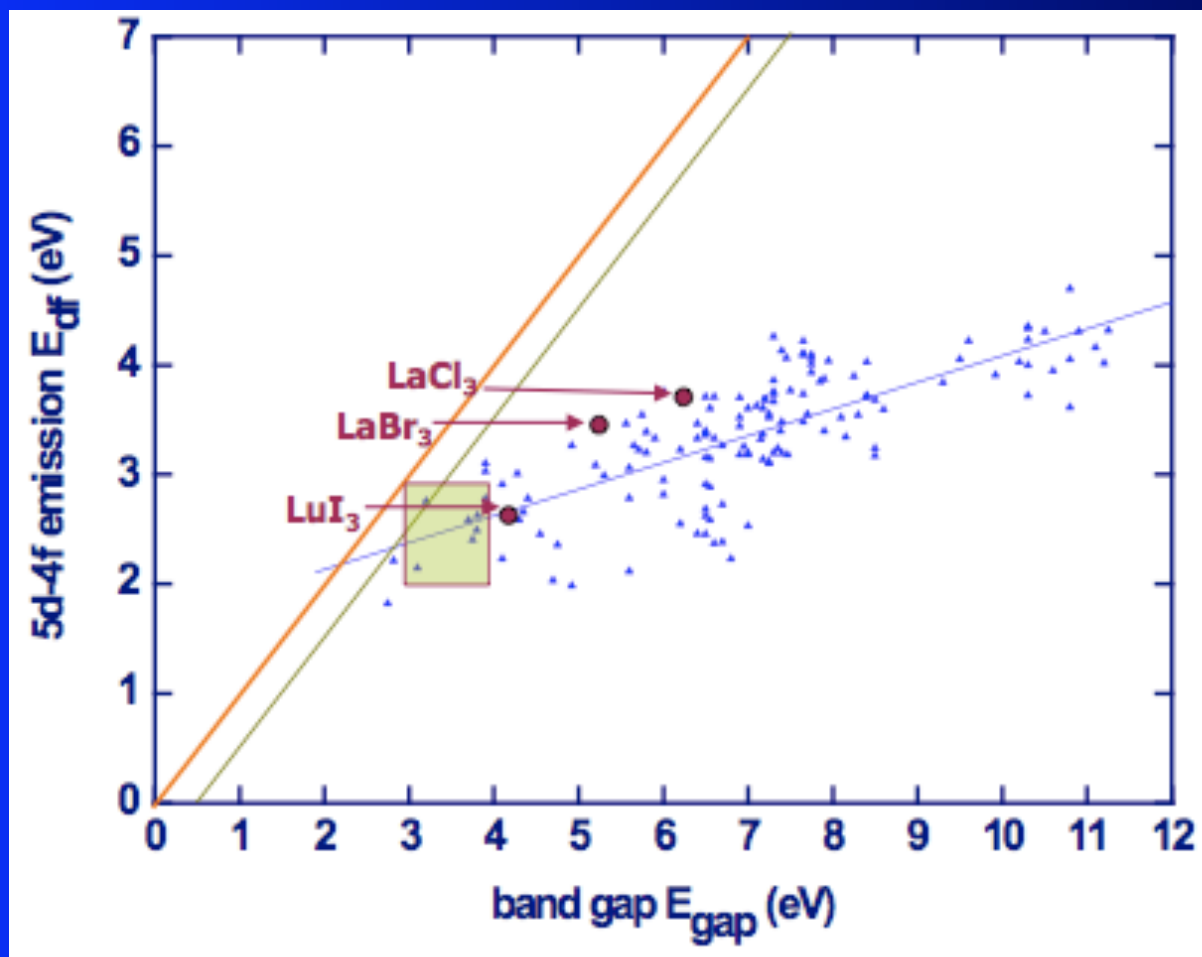
Why?

Towards smaller band gap compounds



Light yield (10 ³ ph/MeV)	2	53	73	0
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The host lattice must be transparent to the Ce^{3+} emission



Best result

$LuI_3: Ce^{3+}$

$E_{gap} = 4.2\text{eV}$

$Y_{ph} = 100000/\text{MeV}$

$\lambda_{em} \approx 470\text{ nm}$

Ultimate Ce^{3+}
Scintillator

$E_{gap} > 2.5\text{-}3\text{ eV}$

$Y_{ph} \leq 140000/\text{MeV}$

$\lambda_{em} \approx 600\text{ nm}$



Why is Ce^{3+} so popular? Limits on the scintillation speed

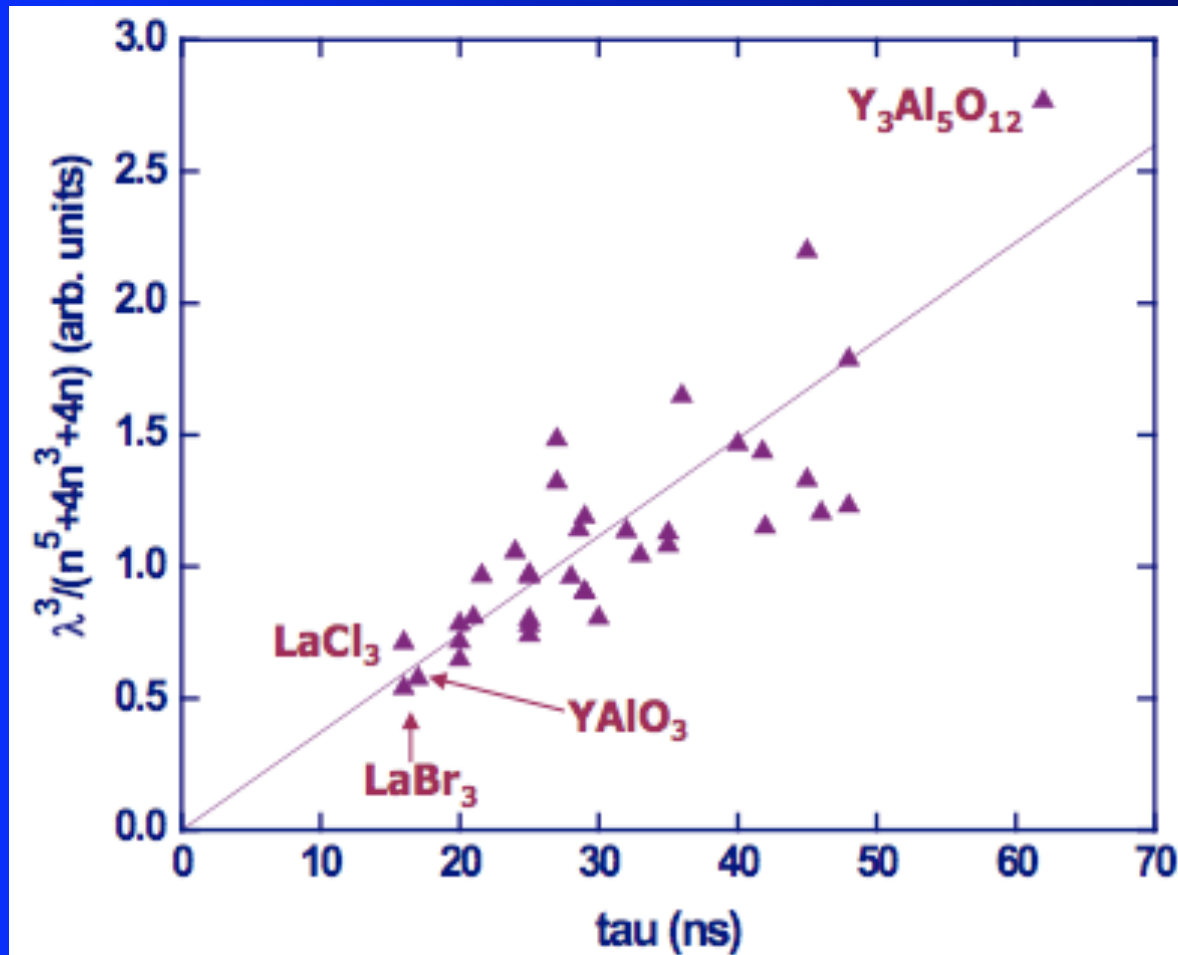


$$\Gamma_v = \frac{1}{\tau_v} \propto \frac{n}{\lambda^3} \left(\frac{n^2 + 2}{3} \right)^2 \sum_f |\langle f | \mu | i \rangle|^2$$

Three important aspects

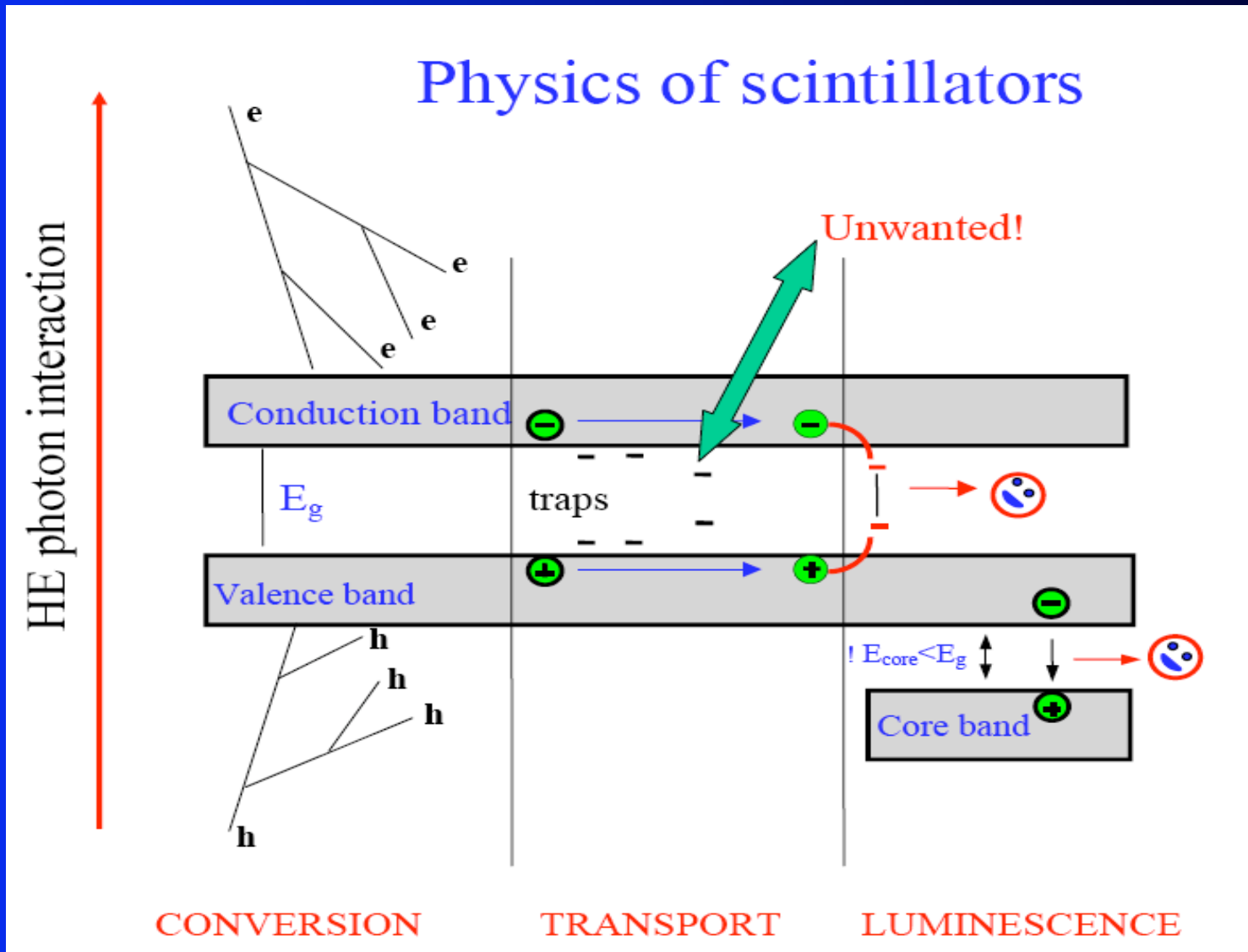
- Dipole and spin allowed transitions
- Short wavelength of emission
- High refractive index

Decay time of Ce^{3+} 5d-4f emission



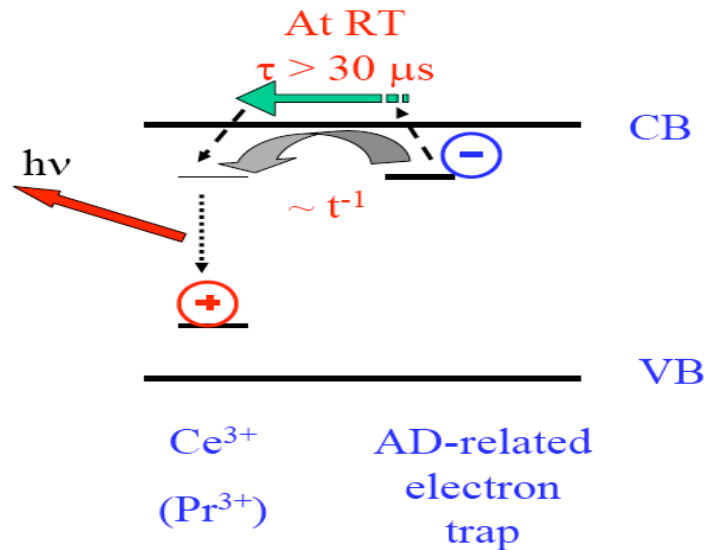
$$\tau \approx A \times \frac{\lambda^3}{n^5 + 4n^3 + 4n}$$

Effect of traps



Effect of traps

Ce^{3+} and Pr^{3+} -doped $Lu_3Al_5O_{12}$



Light yield (1 μs time gate)

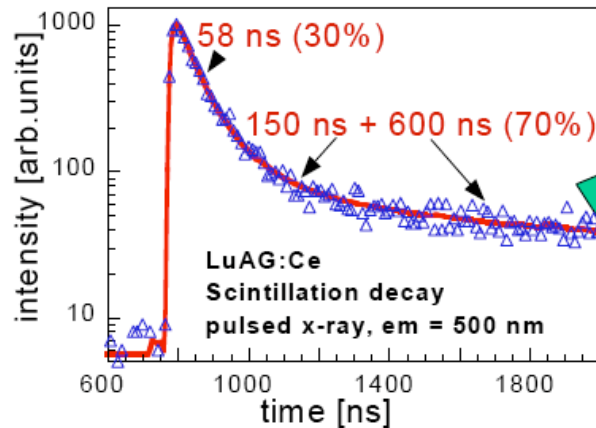
Best YAG:Ce $\sim 3x$ BGO

Best LuAG:Ce $\sim 60\%$ of YAG:Ce

A lot of "slow light" in these materials

Retrapping of electrons at shallow traps before their radiative recombination at Ce^{3+} ions

Nikl et al, pss (a) **201**, R41 (2004)



- Luminescence (laser)

- Characterized by direct interband excitation and relaxation
- Performance determined by Kramers-Kronig relations for the optical transitions

- Scintillation

- Characterized by higher excitation energy than interband transitions
- Performance characterized by:
 - Cooling and relaxation of hot e-h pairs
 - Density and possible interaction of excitations
 - Luminescence mechanism

Thank you