

THE SHiP EXPERIMENT AT CERN

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SHiP

On behalf of the SHiP Collaboration



CERN-SPSC-2015-017
SPSC-P-350-ADD-1
9 April 2015



CERN-SPSC-2015-016
SPSC-P-350
8 April 2015

arXiv:1504.04956v1.

SHiP 240 physicists, 15 Countries

Search for Hidden Particles

Steered west-southwest, and encountered a heavier sea than they had met with before in the whole voyage. Saw particles and a green nuch near the vessel. The crew of the Patria saw a cone and a log; they also picked up a stick which appeared to have been carved with an iron tool, a piece of cone, a plant which grows on land, and a board. The crew of the Nova saw other signs of land, and a stalle loaded with rose berries. These signs encouraged them, and they all grew cheerful. Sailed this day till sunset, twenty-seven leagues.

After sunset steered their original course west and sailed twelve miles an hour till two hours after midnight, going ninety miles, which are twenty-two leagues and a half, and as the Patria was the swiftest sailer, and kept ahead of the Achind,

the discovered land



Technical Proposal



SHiP

arXiv:1504.04855v1.

85 theorists

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Physics Proposal

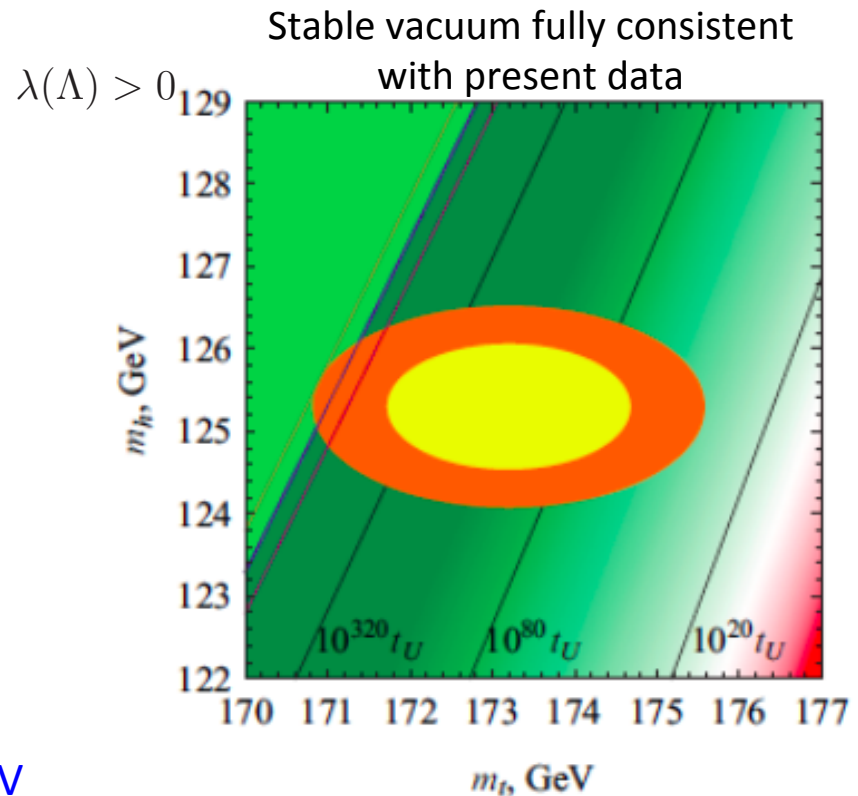
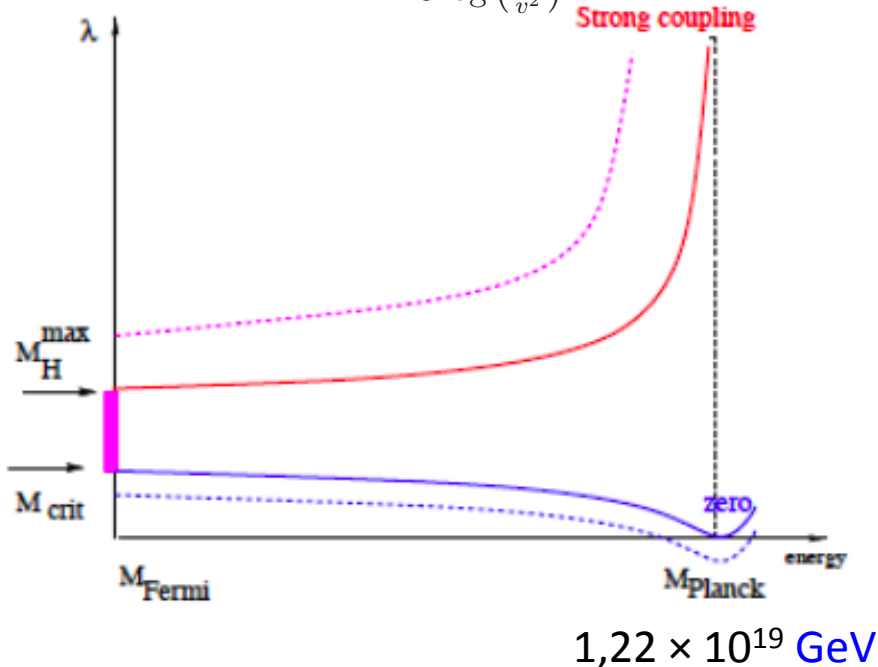
Instrumentation Seminar, DESY

SM may well be a consistent effective theory all the way up to the Plank scale

- ✓ $M_H < 175 \text{ GeV} \rightarrow \text{SM is a weakly coupled theory up to the Plank energies!}$
- ✓ $M_H > 111 \text{ GeV} \rightarrow \text{EW vacuum is stable or metastable with a lifetime greatly exceeding the age of our Universe (Espinosa et al)}$

S. Heinemeyer, Higgs Physics, arXiv:1405.3781

$$\lambda(\Lambda) < \infty \Rightarrow M_H^2 \leq \frac{8\pi^2 v^2}{3 \log\left(\frac{\Lambda^2}{v^2}\right)}$$



G. Degraasi et al., Higgs mass and vacuum stability

- ✓ *No sign of New Physics seen in the SM at NNLO, JHEP 1208 (2012) 098*

Nevertheless, many open questions in particle physics!

Among the most relevant ones:

Why is the Higgs boson so light (so-called “naturalness” or “hierarchy” problem) ?

What is the origin of the matter-antimatter asymmetry in the Universe ?

Why 3 fermion families ? Why do neutral leptons, charged leptons and quarks behave differently ?

What is the origin of neutrino masses and oscillations ?

What is the composition of dark matter (~25% of the Universe) ?

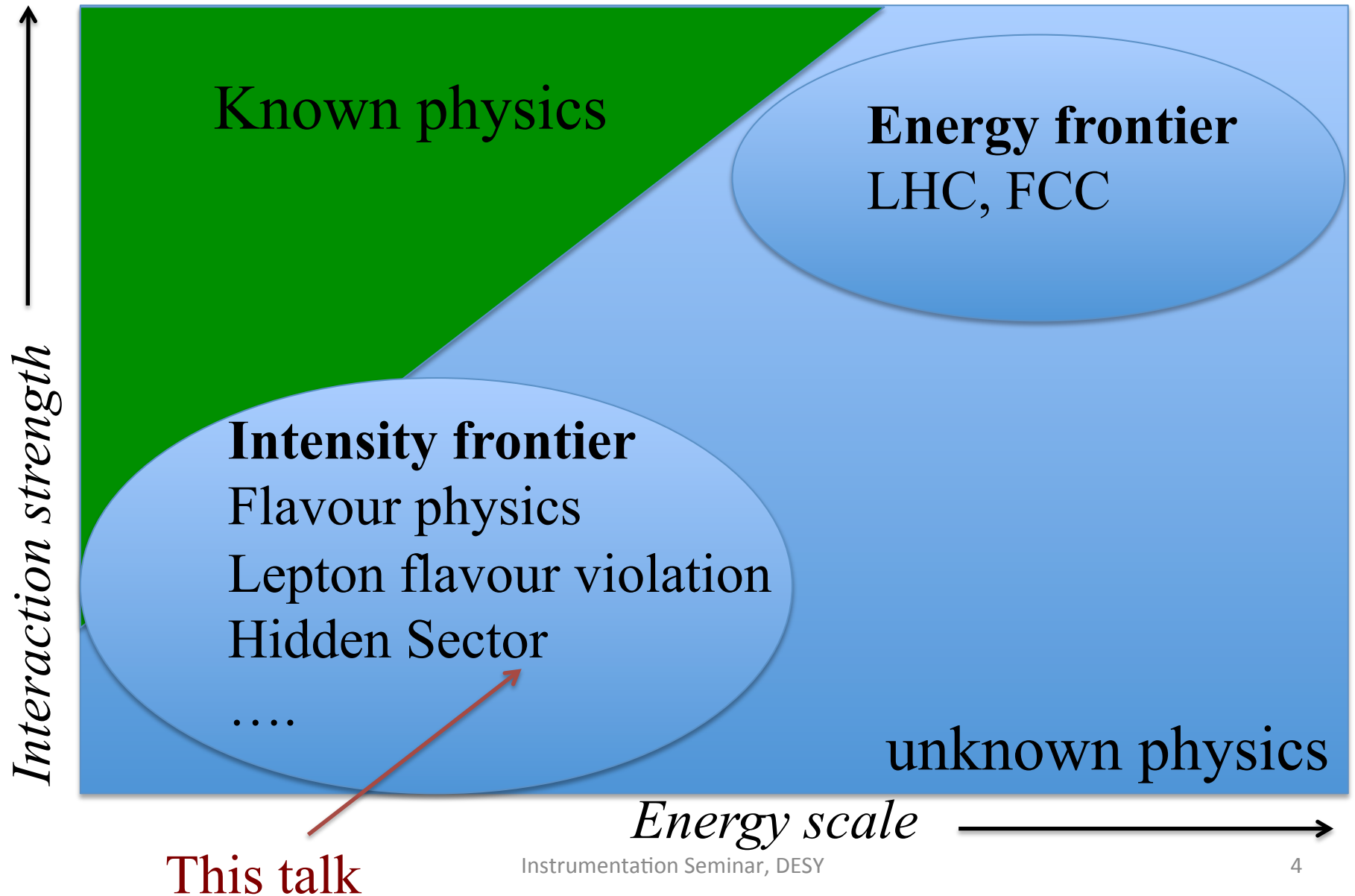


However: there is NO direct evidence for new particles (yet...)
from the LHC or other facilities

Where is the New Physics ?

i.e. at what E scale(s) will we find the answers to these questions ?

High Intensity Frontier



Search for Hidden Sector (HS)

or very weakly interacting NP

$$L = L_{SM} + L_{mediator} + L_{HS}$$

Visible Sector



Mediators or portals to the HS:
vector, scalar, axial, neutrino

Hidden Sector

Naturally accommodates Dark Matter
(may have very complicated structure)

- ✓ HS production and decay rates are strongly suppressed relative to SM
 - Production branching ratios $O(10^{-10})$
 - Long-lived objects
 - Travel unperturbed through ordinary matter

Models	Final states
HNL, SUSY neutralino	$l^+\pi^-, l^+K^-, l^+\rho^- \rightarrow \pi^+\pi^0$
Vector, scalar, axion portals, SUSY sgoldstino	l^+l^-
HNL, SUSY neutralino, axino	$l^+l^-\nu$
Axion portal, SUSY sgoldstino	$\gamma\gamma$
SUSY sgoldstino	$\pi^0\pi^0$

Full reconstruction and PID are essential to minimize model dependence

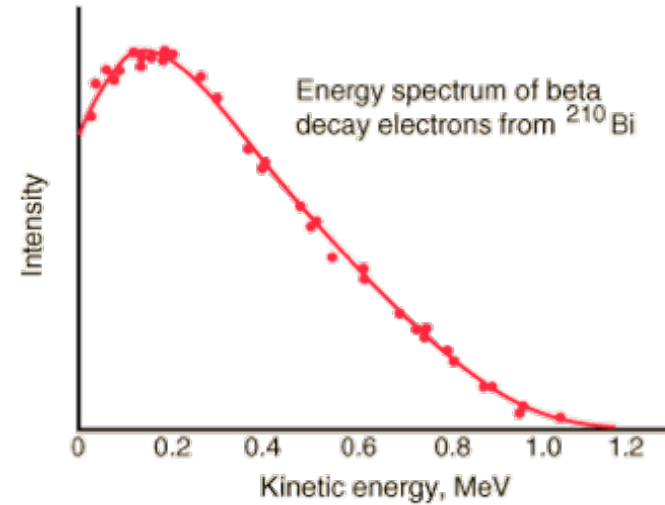
Experimental challenge is background suppression
→ requires $O(0.01)$ carefully estimated

History lesson - 1930s:

- Back then, the “Standard Model” was photon, electron, nucleons

- Beta decay: $n \rightarrow p + e^-$

Continuous spectrum!



- Pauli proposes a radical solution - the neutrino!

$$n \rightarrow p + e^- + \bar{\nu}$$

- Great example of a hidden sector!

- neutrino is electrically neutral (QED gauge singlet)
- very weakly interacting and light
- interacts with “Standard Model” through “portal” -

$$(\bar{p}\gamma^\mu n)(\bar{e}\gamma_\mu \nu)$$

Search for dark photons

- Assuming no lighter hidden particles, γ' decay into SM particles through a virtual photon:

$$\gamma' \rightarrow e^+e^-, \quad \mu^+\mu^-, \quad q\bar{q}, \dots$$

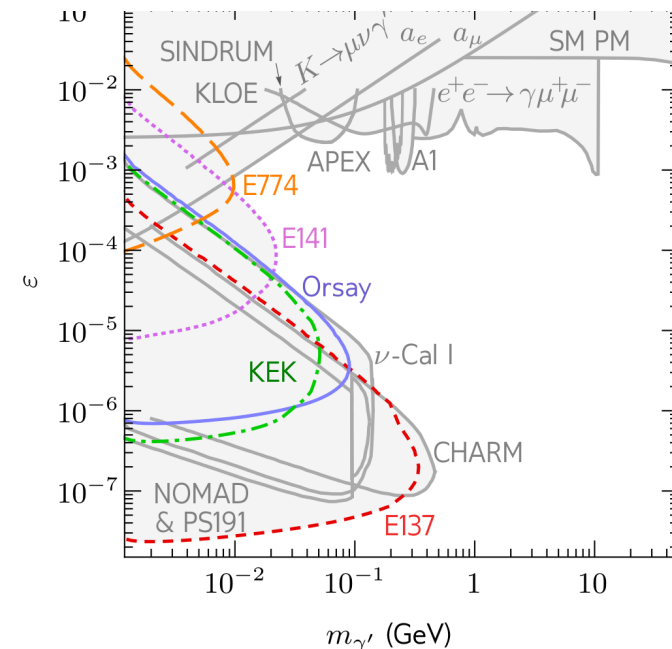
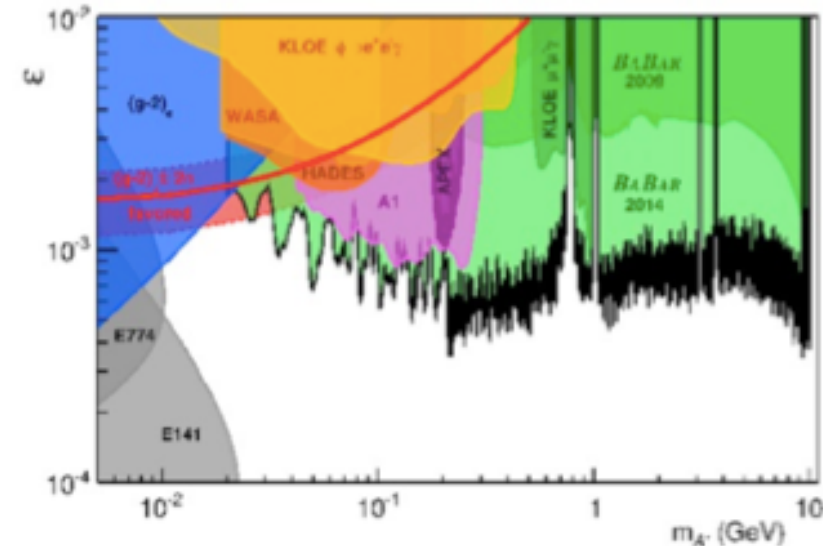
- decay length $c\tau \sim \varepsilon^{-2} m_{\gamma'}^{-1}$
- cosmological constraints (nucleo-synthesis):
 $\tau < 0.1 \text{ s} \Rightarrow \varepsilon^2 m_{\gamma'} > 10^{-21} \text{ GeV}$

γ' production

- proton bremsstrahlung:
 - initial-state radiation from the incoming proton, followed by a hard proton-nucleus interaction
- secondary particles decay:

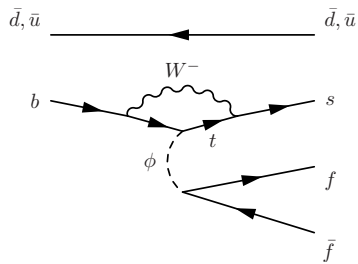
Mass interval (GeV)	Process	$n_{\gamma'}/p.o.t$
$m_{\gamma'} < 0.135$	$\pi^0 \rightarrow \gamma\gamma'$	$\varepsilon^2 \times 5.41$
$0.135 < m_{\gamma'} < 0.548$	$\eta \rightarrow \gamma\gamma'$	$\varepsilon^2 \times 0.23$
$0.548 < m_{\gamma'} < 0.648$	$\omega \rightarrow \pi^0\gamma'$	$\varepsilon^2 \times 0.07$
$0.648 < m_{\gamma'} < 0.958$	$\eta' \rightarrow \gamma\gamma'$	$\varepsilon^2 \times 10^{-3}$

Instrumentation Seminar, DESY



Higgs (scalar) portal: production and decay modes

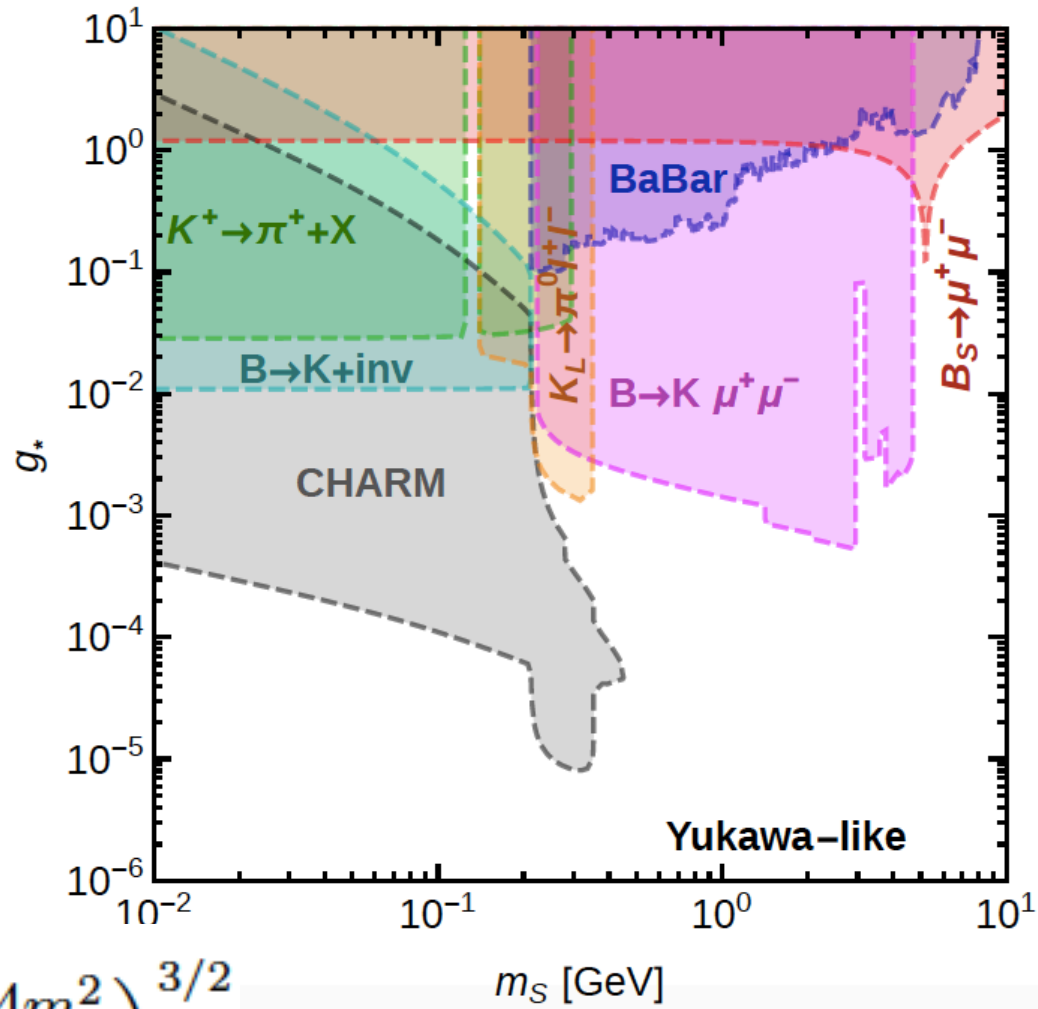
Rare B meson decays mediated by a light scalar ϕ



$$\Gamma(D \rightarrow \pi \phi) \sim (m_b^2 |V_{cb}^* V_{ub}|)^2 \propto m_b^4 \lambda^5$$

$$\Gamma(B \rightarrow K \phi) \sim (m_t^2 |V_{ts}^* V_{tb}|)^2 \propto m_t^4 \lambda^2$$

B decays favoured compared to D



$$\Gamma(S \rightarrow \ell \bar{\ell}) = \frac{g_*^2 m_\ell^2 m_S}{8\pi v^2} \left(1 - \frac{4m_\ell^2}{m_S^2}\right)^{3/2}$$

Motivation for Heavy Neutral Leptons

See-saw mechanism for neutrino masses

Most general renormalisable Lagrangian of SM particles (+3 singlets wrt SM gauge group):

$$L_{\text{singlet}} = i\bar{N}_I \partial_\mu \gamma^\mu N_I - Y_{I\alpha} \bar{N}_I^c \tilde{H} L_\alpha - M_I \bar{N}_I^c N_I + h.c.$$

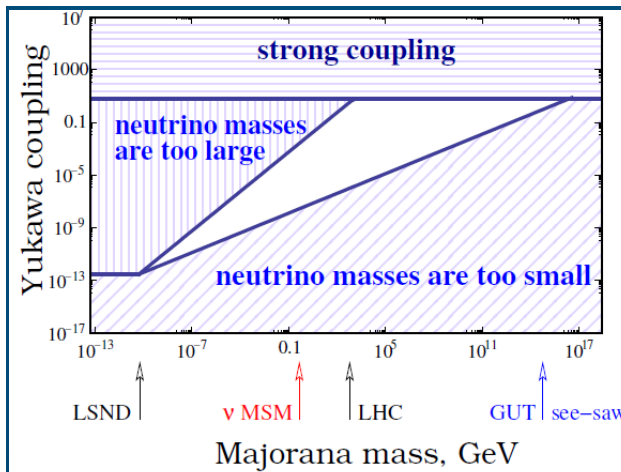
Yukawa term: mixing of N_I with active neutrinos to explain oscillations

Majorana term which carries no gauge charge

$$v \sim 246 \text{ GeV}$$

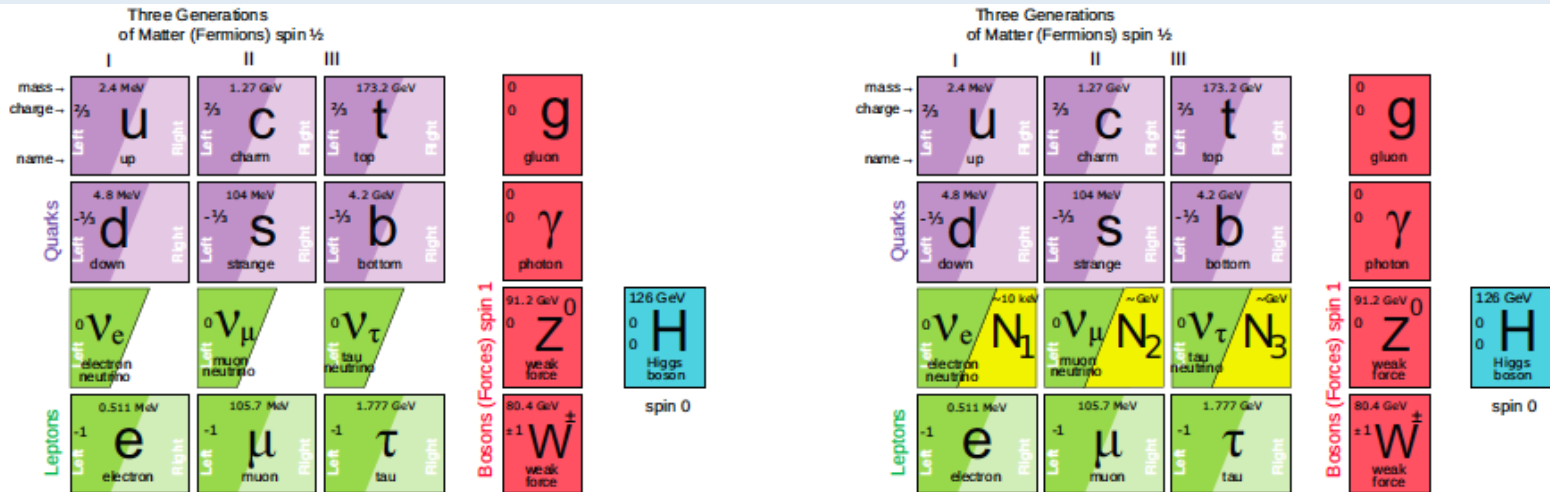
The scale of the active neutrino mass is given by the see-saw formula: $m_\nu \sim \frac{m_D^2}{M}$ where $m_D \sim Y_{I\alpha} v$ - typical value of the Dirac mass term

Four “popular” N mass ranges



	N mass	v masses	eV v anomalies	BAU	DM	M _H stability	direct search	experiment
GUT see-saw	10 ⁻¹⁶ - 10 ¹⁶ GeV	YES	NO	YES	NO	NO	NO	-
EWSB	10 ²⁻³ GeV	YES	NO	YES	NO	YES	YES	LHC
v MSM	keV - GeV	YES	NO	YES	YES	YES	YES	a'la CHARM
v scale	eV	YES	YES	NO	NO	YES	YES	a'la LSND

The ν MSM model: leptogenesis and dark matter



N = Heavy Neutral Lepton - HNL

Role of N_1 with mass in keV region: dark matter

Role of N_2 , N_3 with mass in 100 MeV – GeV region: “give” masses to neutrinos and produce baryon asymmetry of the Universe

Role of the Higgs: give masses to quarks, leptons, Z and W and inflate the Universe.

ν MSM: T.Asaka, M.Shaposhnikov PL B620 (2005) 17
M.Shaposhnikov Nucl. Phys. B763 (2007) 49

global lepton-number symmetry broken at the level of $O(10^{-4})$ leads to the required pattern of sterile neutrino masses consistent with neutrino oscillations data

Masses and couplings of HNLs

- $M(N_2) \approx M(N_3) \sim \text{a few GeV} \rightarrow$ CPV can be increased dramatically to explain **Baryon Asymmetry of the Universe (BAU)**

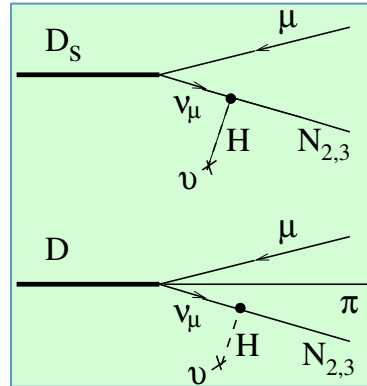
Very weak $N_{2,3}$ -to- ν mixing ($\sim U^2$) $\rightarrow N_{2,3}$ are much longer-lived than SM particles

- Produced in semi-leptonic decays,

$$K \rightarrow \mu\nu, D \rightarrow \mu\pi\nu, B \rightarrow D\mu\nu$$

Example:

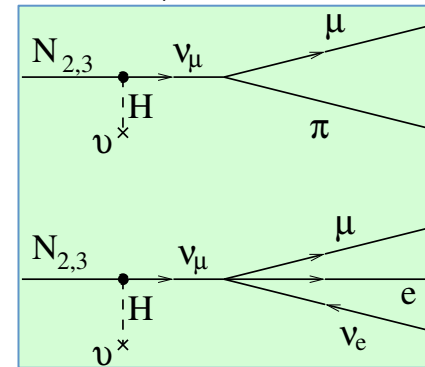
$N_{2,3}$ production in charm



- $\propto \sigma_D \times U^2$

- $U_2^2 = U_{2,\nu_e}^2 + U_{2,\nu_\mu}^2 + U_{2,\nu_\tau}^2$

and subsequent decays



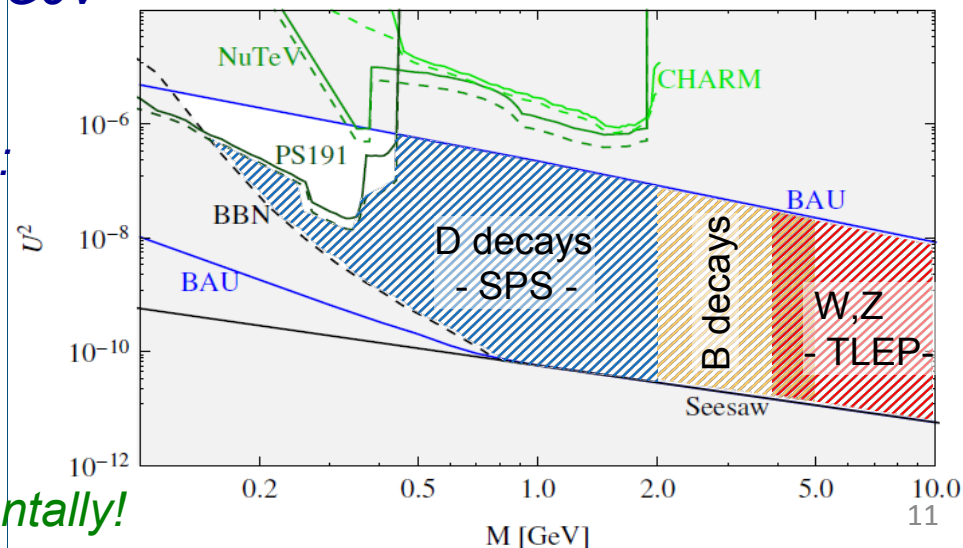
- Typical lifetimes $> 10 \mu\text{s}$ for $M(N_{2,3}) \sim 1 \text{ GeV}$
Decay distance $\mathcal{O}(\text{km})$

- Typical BRs (depend on flavour mixing):

$$\text{Br}(N \rightarrow \mu/e \pi) \sim 0.1 - 50\%$$

$$\text{Br}(N \rightarrow \mu/e^- \rho^+) \sim 0.5 - 20\%$$

$$\text{Br}(N \rightarrow \nu\mu e) \sim 1 - 10\%$$



Domain only marginally explored, experimentally!



Common experimental features of Hidden Sector (HS)

✓ Production through hadron decays (π , K , D , B , proton bremsstrahlung, ...)

✓ Decays:

Models	Final states
Neutrino portal, SUSY neutralino	$\ell^\pm \pi^\mp, \ell^\pm K^\mp, \ell^\pm \rho^\mp, \rho^\pm \rightarrow \pi^\pm \pi^0$
Vector, scalar, axion portals, SUSY sgoldstino	$\ell^+ \ell^-$
Vector, scalar, axion portals, SUSY sgoldstino	$\pi^+ \pi^-, K^+ K^-$
Neutrino portal, SUSY neutralino, axino	$\ell^+ \ell^- \nu$
Axion portal, SUSY sgoldstino	$\gamma \gamma$
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✓ Full reconstruction and PID are essential to minimize model dependence

✓ Production and decay rates are strongly suppressed when compared to SM

- Production branching ratios $O(10^{-10})$
- Long-lived objects
- Travel unperturbed through ordinary matter

✓ **Challenge is background suppression \rightarrow requires $O(0.01)$ carefully estimated**

✓ **Physics with ν_τ produced in D_s decays share many of these features**

ν_τ STUDIES

- Less known particle in the Standard Model
- **First observation** by DONUT at Fermilab in 2001 with 4 detected candidates, *Phys. Lett. B504 (2001) 218-224*
- 9 events (with an estimated background of 1.5) reported in 2008 with looser cuts

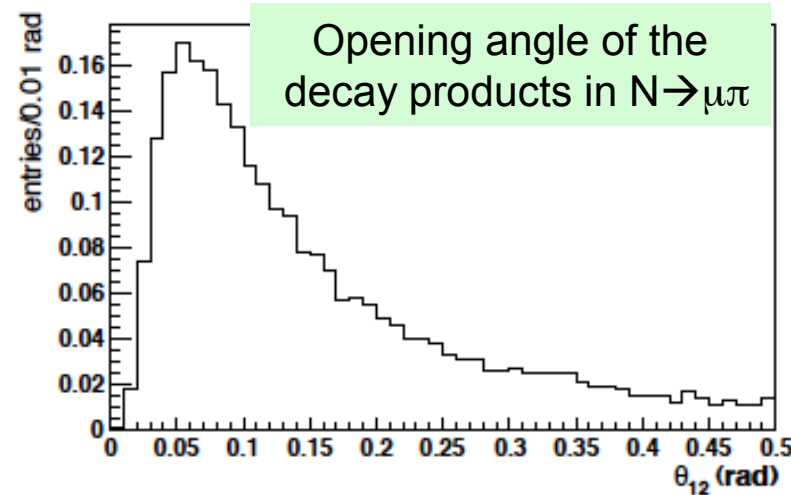
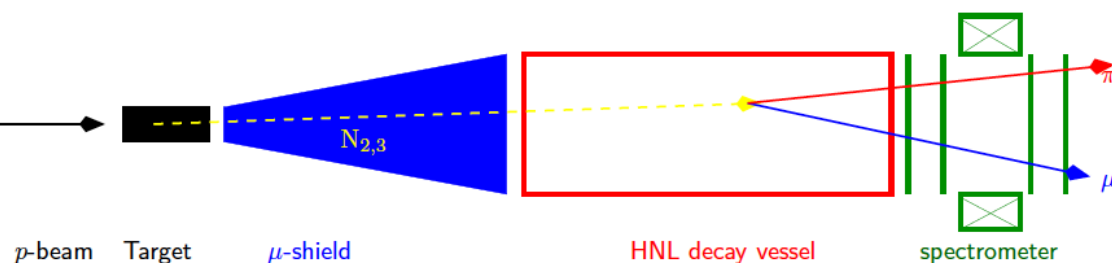
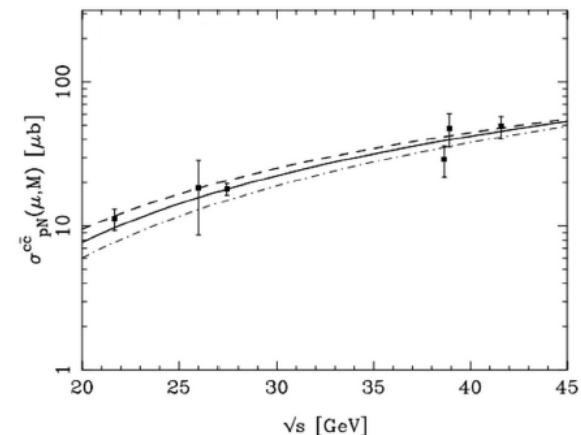
$$\sigma^{\text{const}}(\nu_\tau) = (0.39 \pm 0.13 \pm 0.13) \times 10^{-38} \text{ cm}^2 \text{ GeV}^{-1}$$

- 5 ν_τ candidates reported by OPERA for the discovery (5.1 σ result) of **ν_τ appearance** in the CNGS neutrino beam *PRL 115 (2015) 121802*
- Tau anti-neutrino never observed

$$N_{\nu_\tau + \bar{\nu}_\tau} = 4N_p \frac{\sigma_{c\bar{c}}}{\sigma_{pN}} f_{D_s} Br(D_s \rightarrow \tau) = 2.85 \times 10^{-5} N_p = 5.7 \times 10^{15}$$

General experimental requirements

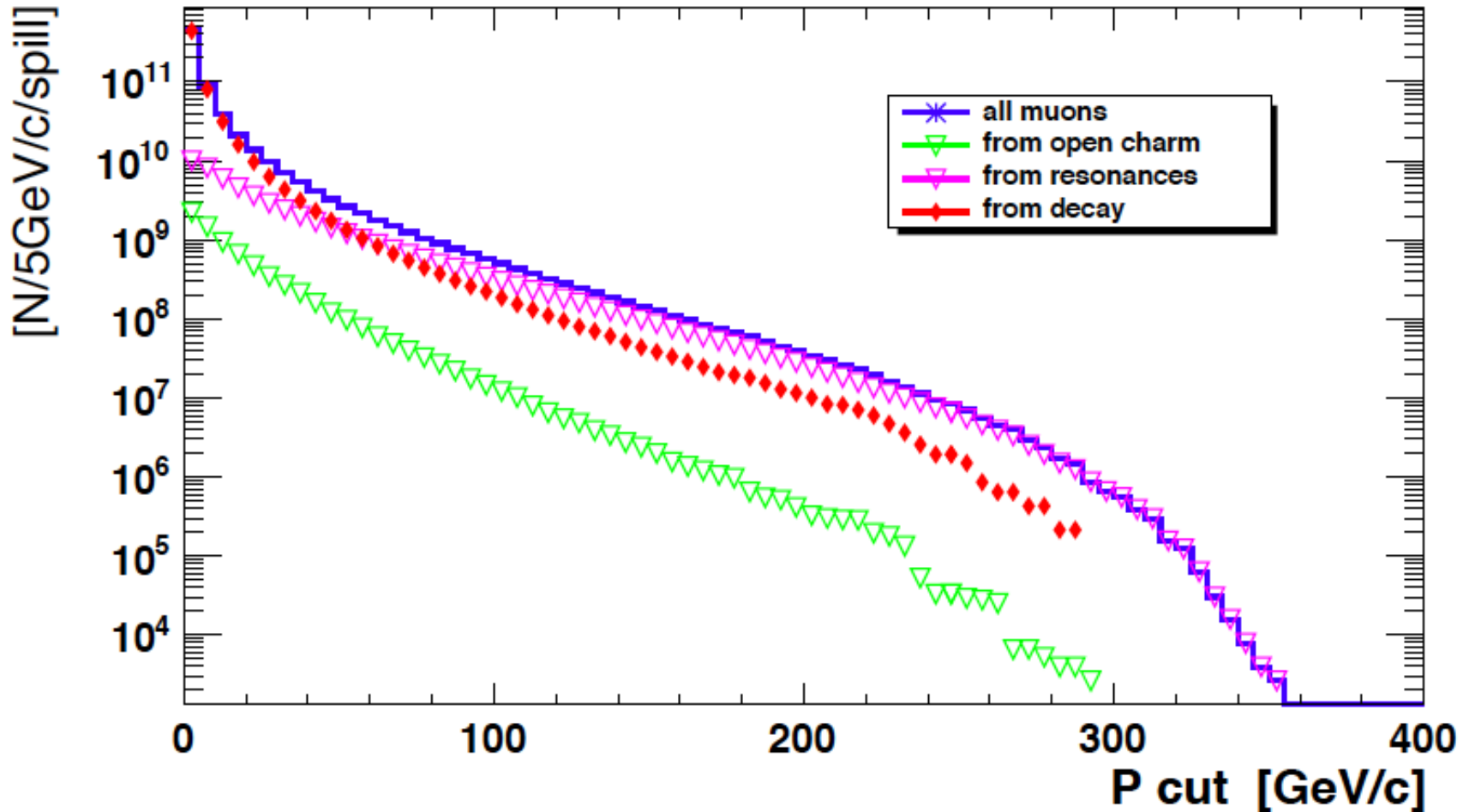
- ✓ Search for HS particles in Heavy Flavour decays
- ✓ HS produced in charm and beauty decays have significant P_T



- ✓ Detector must be placed close to the target to maximize geometrical acceptance
- ✓ Effective (and “short”) muon shield is essential to reduce muon-induced backgrounds
- ✓ With 2×10^{20} 400 GeV pot, $\sim 3 \times 10^{17}$ charm produced

Main sources of muons in beam dump

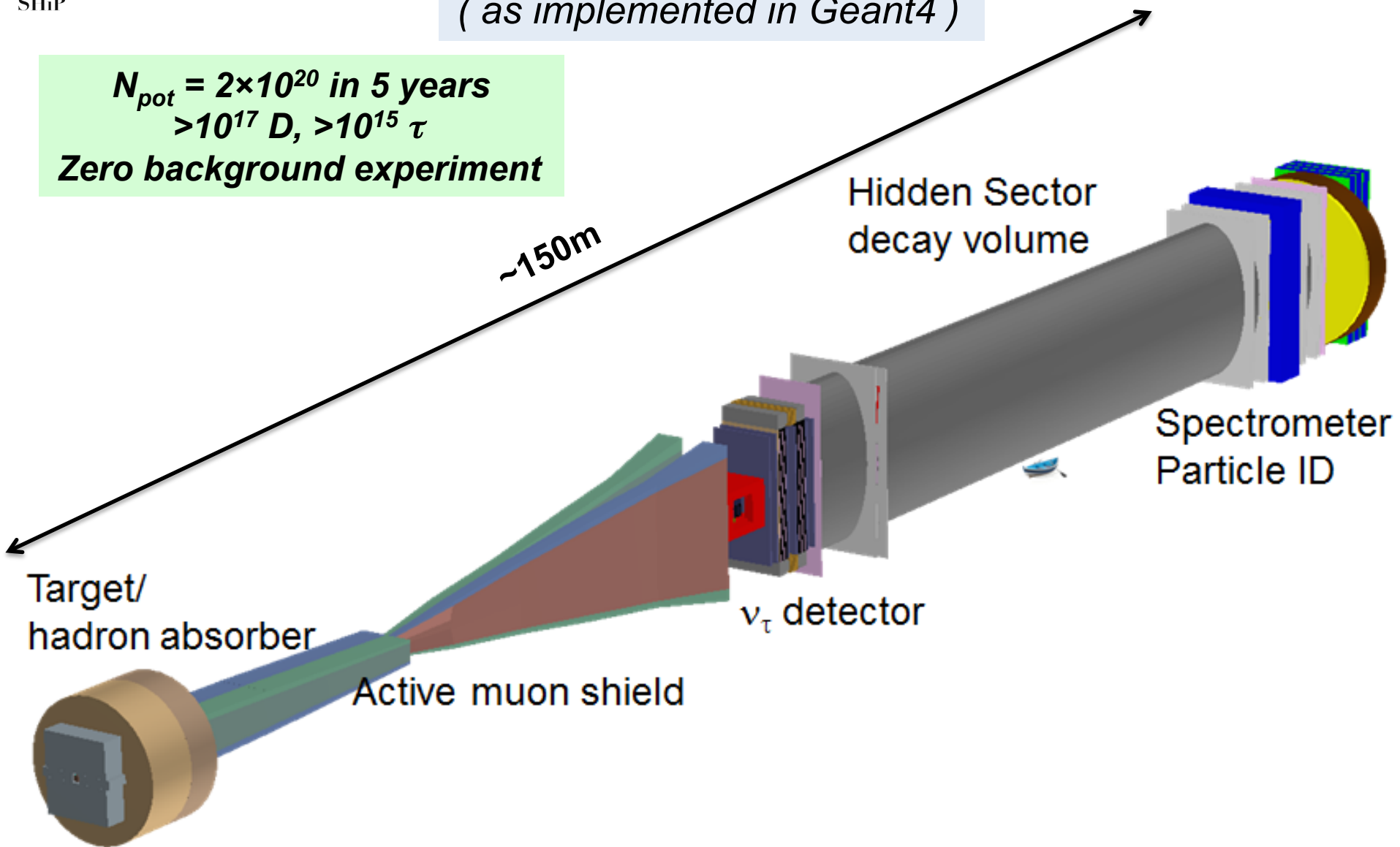
- Decays of pions populate mainly low momenta
- Electromagnetic decays of resonances (η , ρ , etc) populate mainly high momenta
- Negligible fraction of muons from charm decays





The SHiP experiment (as implemented in Geant4)

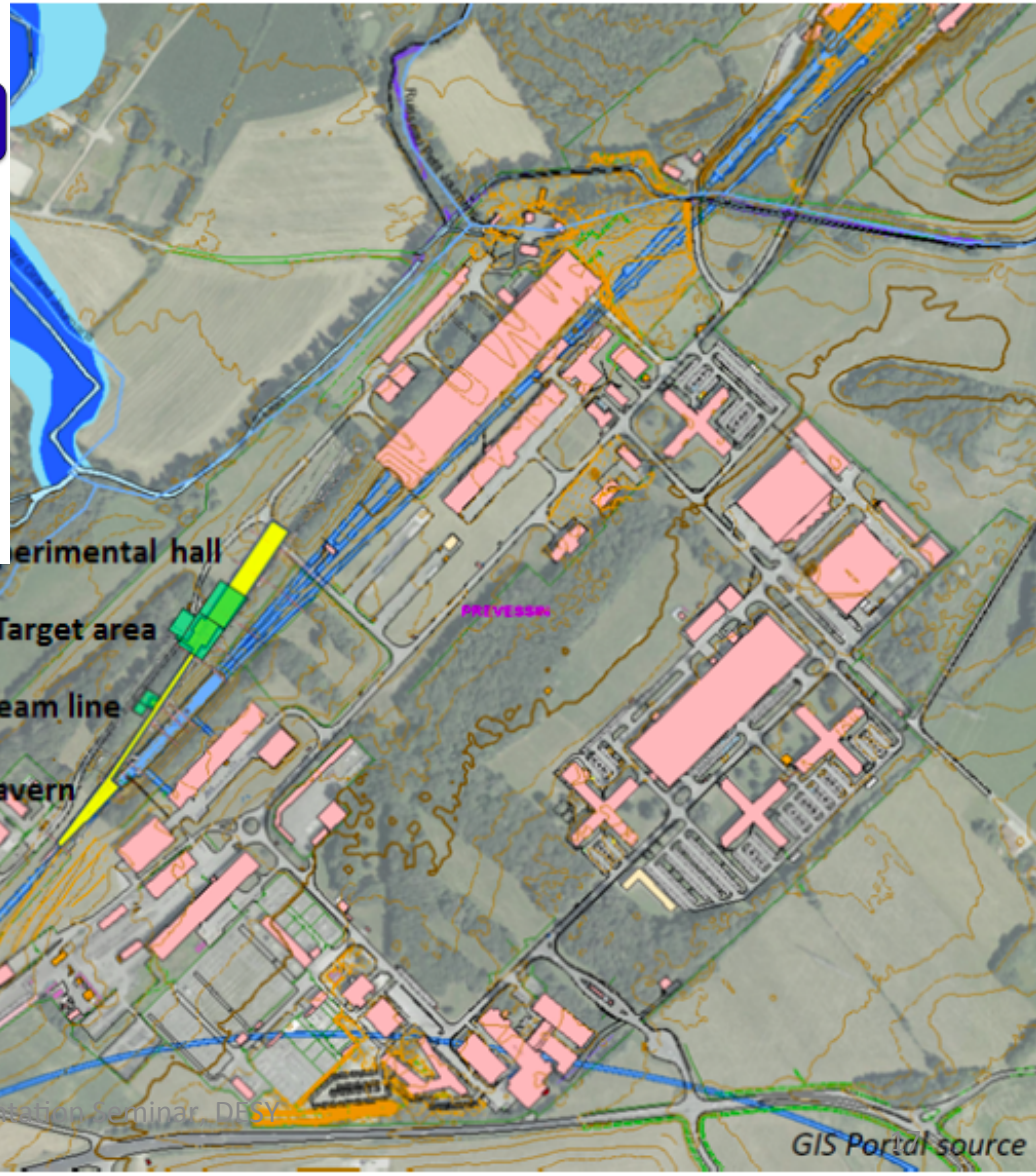
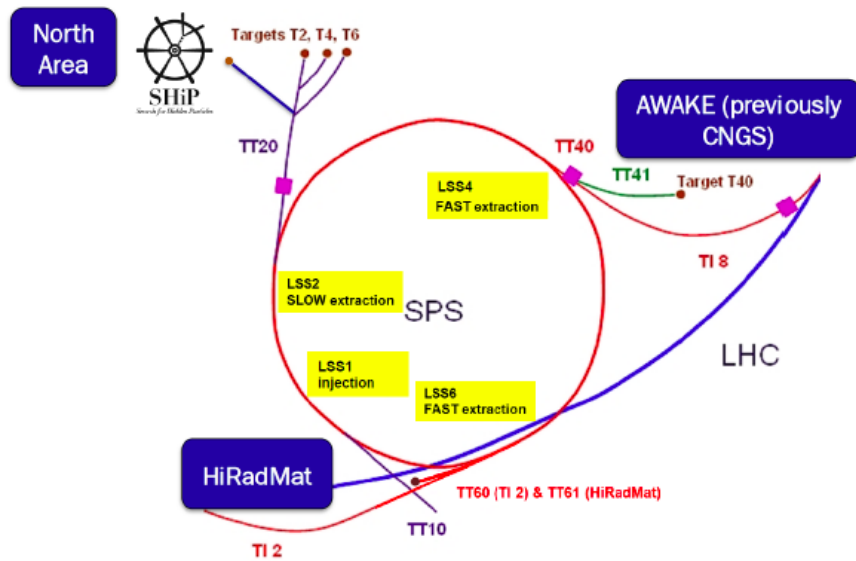
$N_{pot} = 2 \times 10^{20}$ in 5 years
 $> 10^{17} D, > 10^{15} \tau$
Zero background experiment





The Fixed-target facility at the SPS: Preveessin North Area site

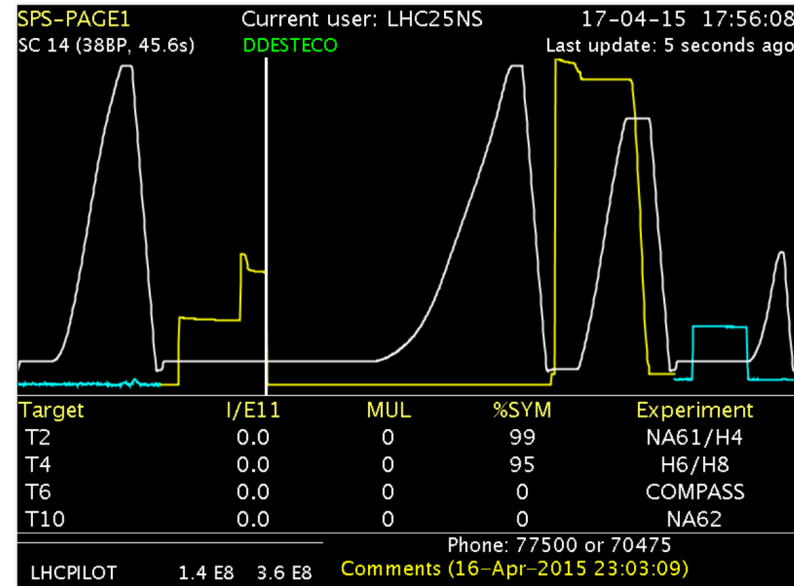
Proposed implementation based on minimal modification of the SPS complex
High-intensity proton beam: $4 \cdot 10^{13}$ ppp, $4 \cdot 10^{19}$ pot/yr, 5 years run (as for CNGS)



The SHiP facility is located on the North Area, and shares the TT20 transfer line and slow extraction mode with the fixed target programmes

R&D at CERN for extraction and beam lines

- Deployment of the new SHiP cycle
- Extraction loss characterisation and optimisation
 - Reduce p density on septum wires
 - Probe SPS aperture limits during slow extraction
- Development of new TT20 optics
 - Change beam at splitter on cycle-to cycle basis
- Characterisation of spill structure
- R&D and development of laminated splitter and dilution (sweep) magnets



Successful test in April 2015

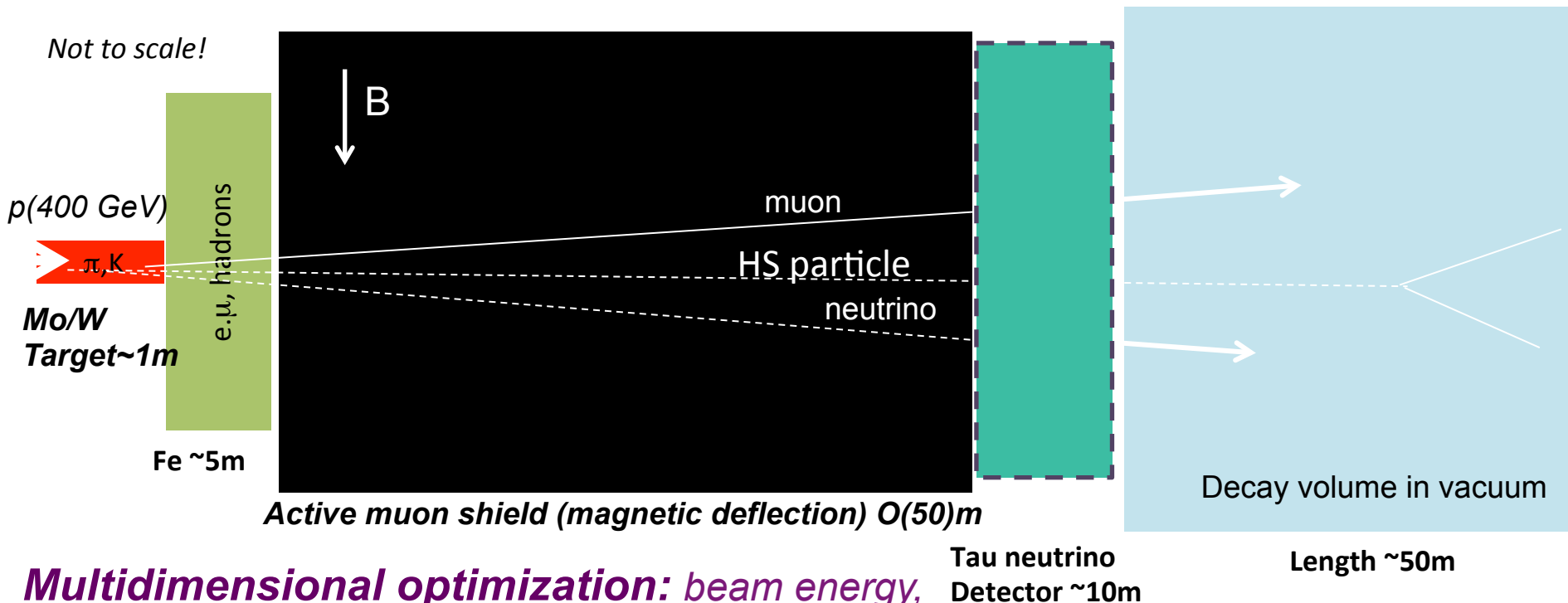
SHiP beam-line

(incompatible with conventional neutrino facility)

Initial reduction of beam induced backgrounds

- Heavy target to maximize Heavy Flavour production (large A) and minimize production of neutrinos in $\pi/K \rightarrow \mu\nu$ decays (short λ_{int})
- Hadron absorber
- Effective muon shield (without shield: muon rate $\sim 10^{10}$ per spill of 4×10^{13} pot)
- Slow (and uniform) beam extraction ~ 1 s to reduce occupancy in the detector

Not to scale!

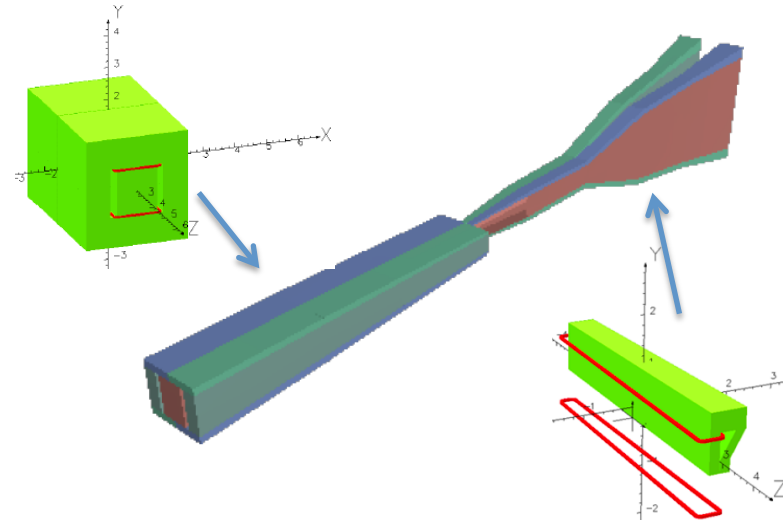


Multidimensional optimization: beam energy, beam intensity, background conditions and detector acceptance

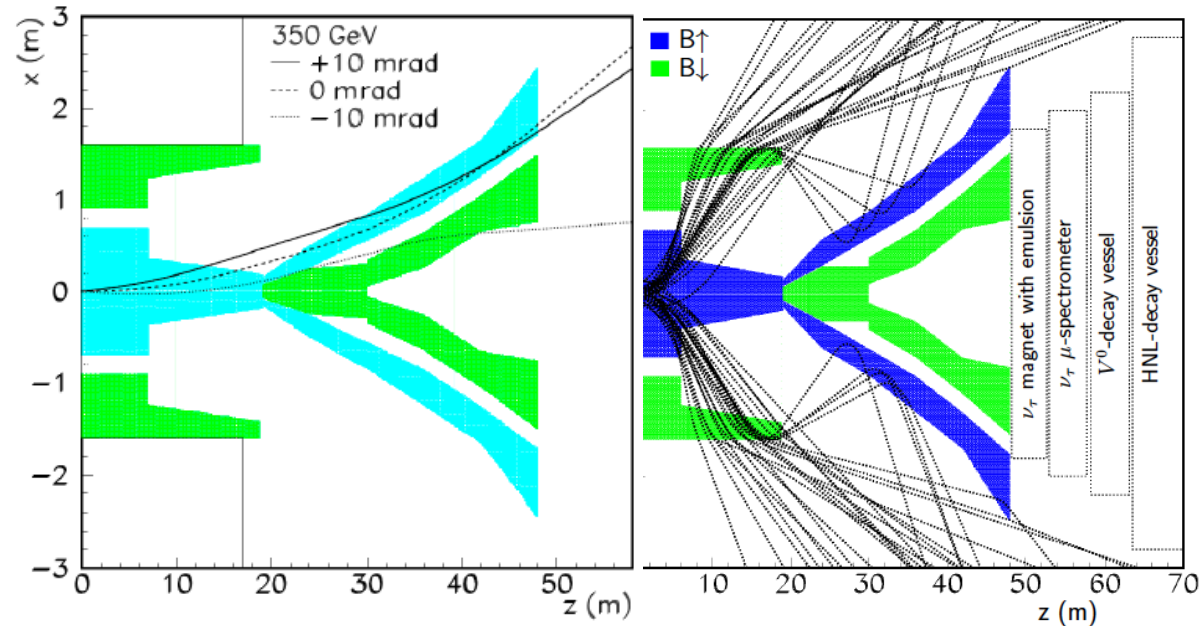


SHiP muon shield

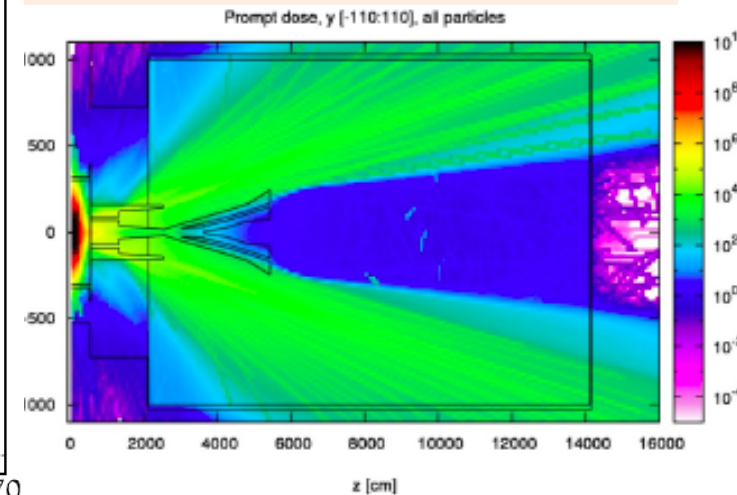
- ✓ Muon flux limit driven by HS background and emulsion-based neutrino detector
- ✓ Active muon shield based entirely on magnet sweeper with a total field integral $B_y = 86.4 \text{ Tm}$
- Realistic design of sweeper magnets in progress
- Challenges: flux leakage, constant field profile, modeling magnet shape
- ✓ $< 7k \text{ muons / spill } (E_\mu > 3 \text{ GeV}), \text{ from } 10^{10}$
- ✓ Negligible flux in terms of detector occupancy



Magnetic sweeper field



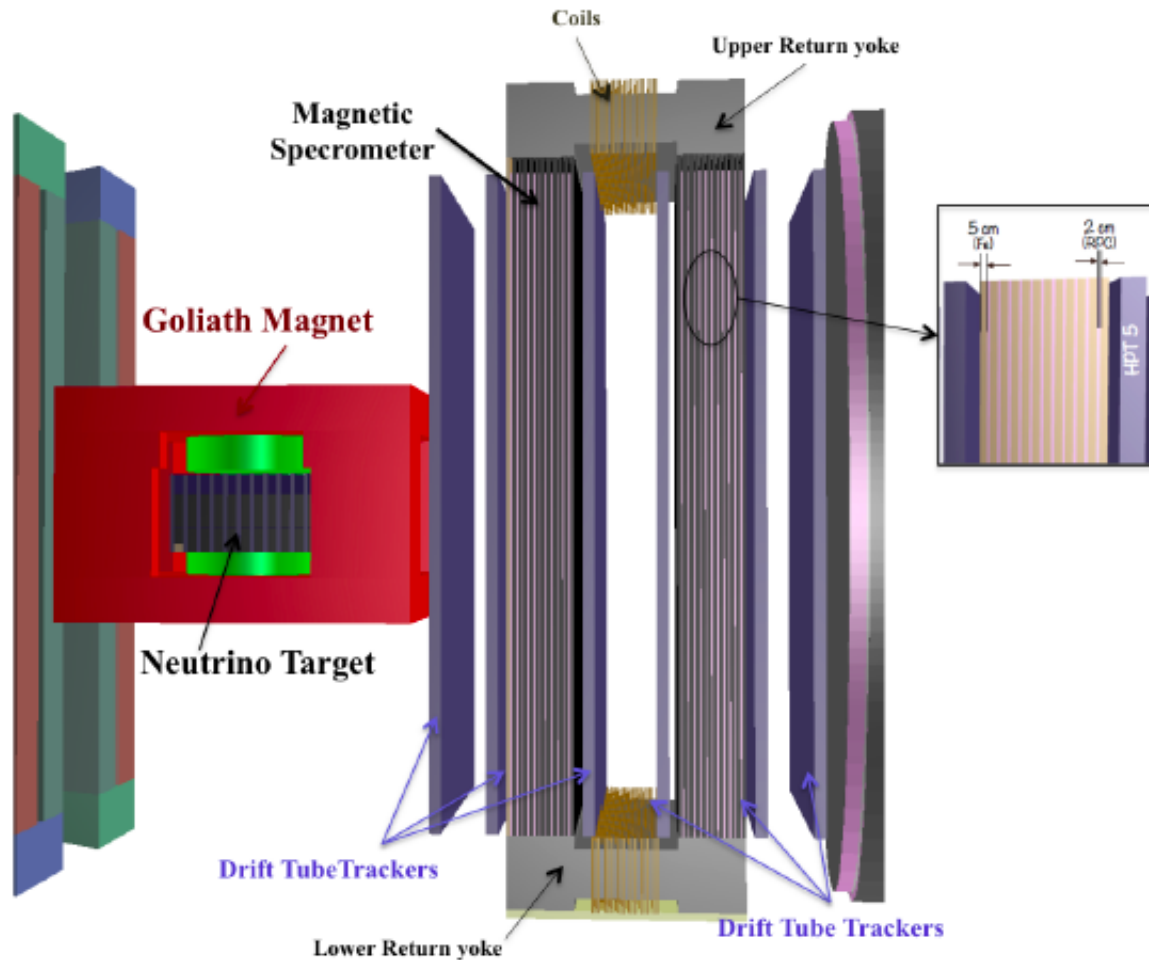
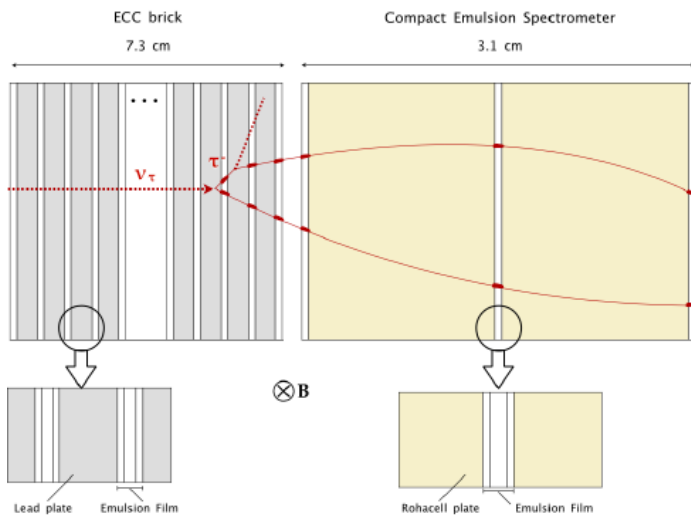
Dose rate ($\mu\text{Sv/h}$) in the SHiP hall



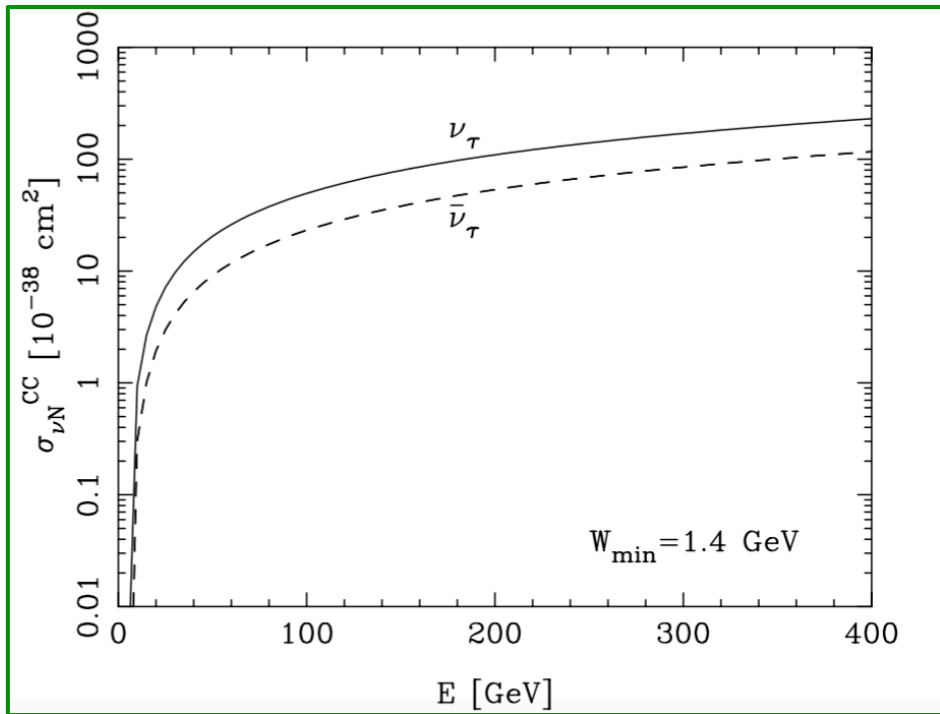
ν_τ detector follows the concept of OPERA



**Emulsion Cloud Chamber
is the key element of ν_τ detection**



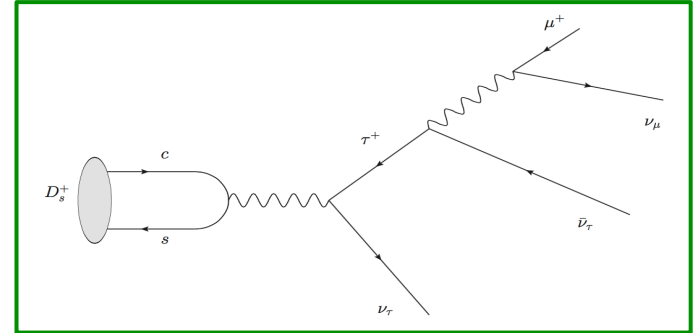
ν_τ INTERACTIONS IN THE TARGET



M. H. Reno, Phys. Rev. D74 (2006) 033001

Uncertainty ($\lesssim 10\%$) from:

- Scale choices
- Pdf
- Target mass correction



Expected number of interactions*

*in 5 years run (2×10^{20} pot)

target mass ~ 9.6 ton (Pb)

$$N_{\nu_\tau} \simeq 6.7 \times 10^3$$

$$N_{\bar{\nu}_\tau} \simeq 3.4 \times 10^3$$

20% uncertainty mainly
from scale variations in
c-cbar differential cross-section

ν_τ DETECTOR

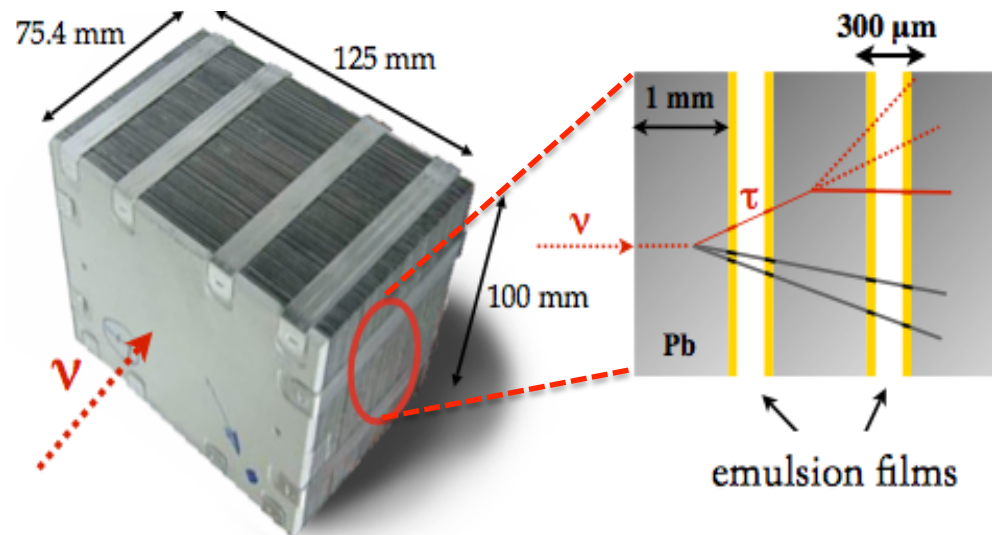
THE UNITARY CELL

Emulsion Cloud Chamber (ECC)

BRICK

- passive material \rightarrow lead
(*massive target*)
- tracking device \rightarrow nuclear emulsions
(*high resolution*)

$10 X_0$



sensitivity 30 grains/100 μm

mip \rightarrow

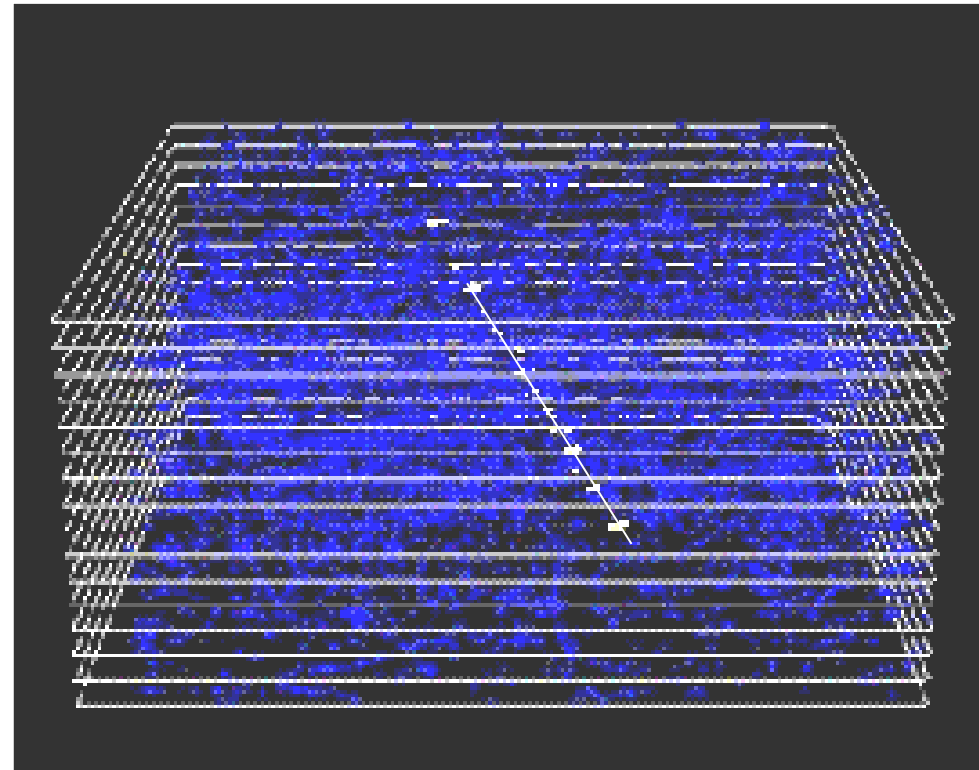
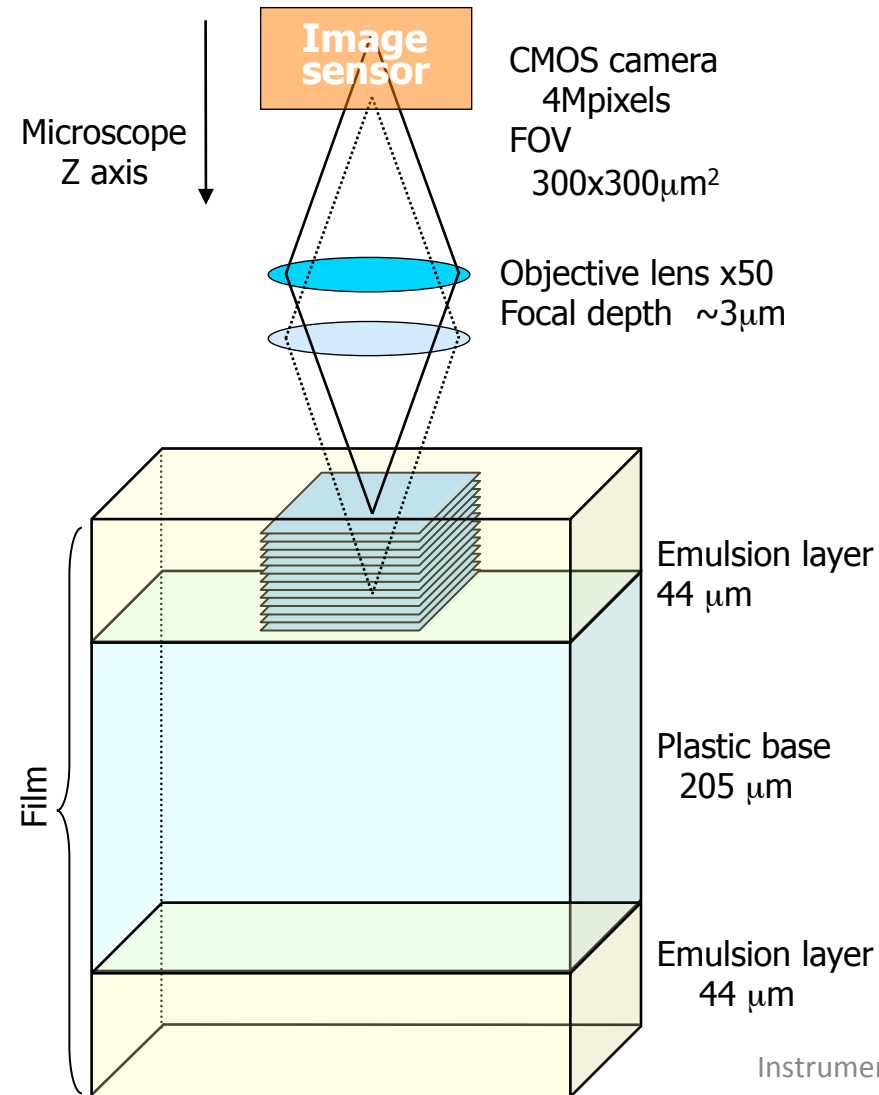
NIM A556 (2006) 80-86

PERFORMANCES

- Primary and secondary **vertex definition** with μm resolution
- **Momentum measurement** by Multiple Coulomb Scattering
 - largely exploited in the OPERA experiment
- **Electron identification**: shower ID through calorimetric technique

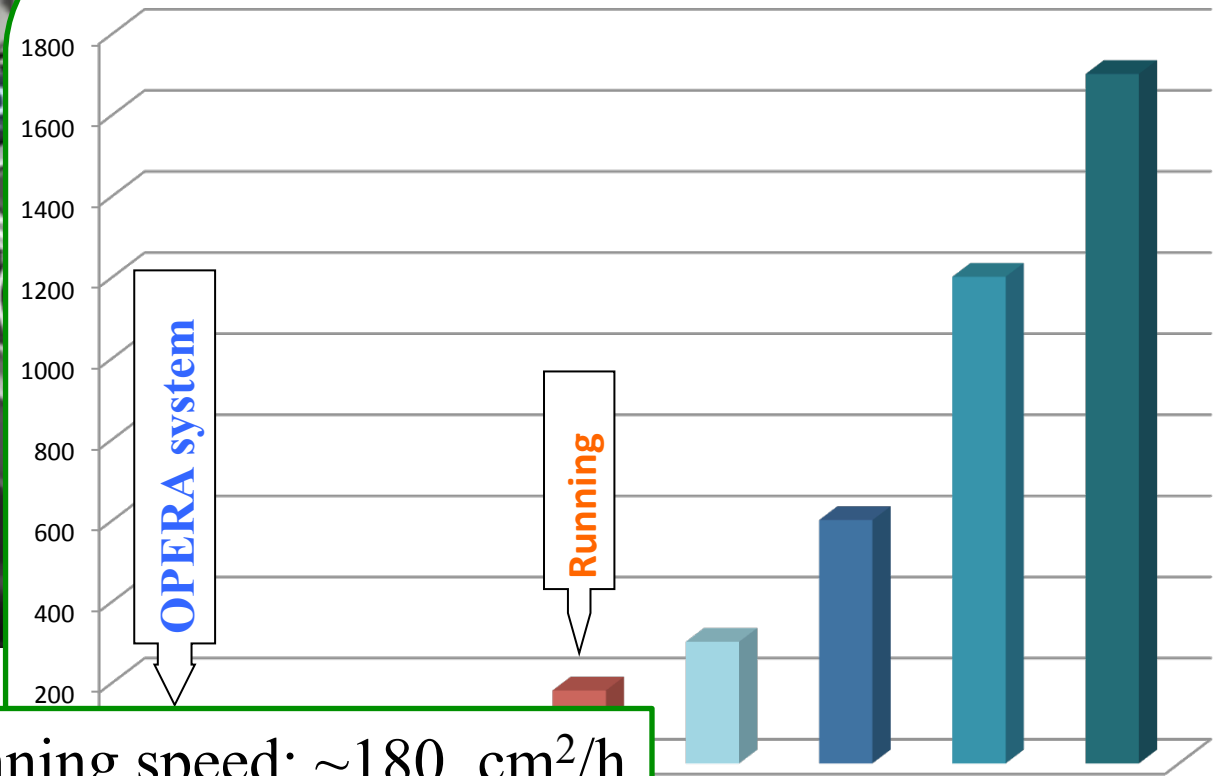
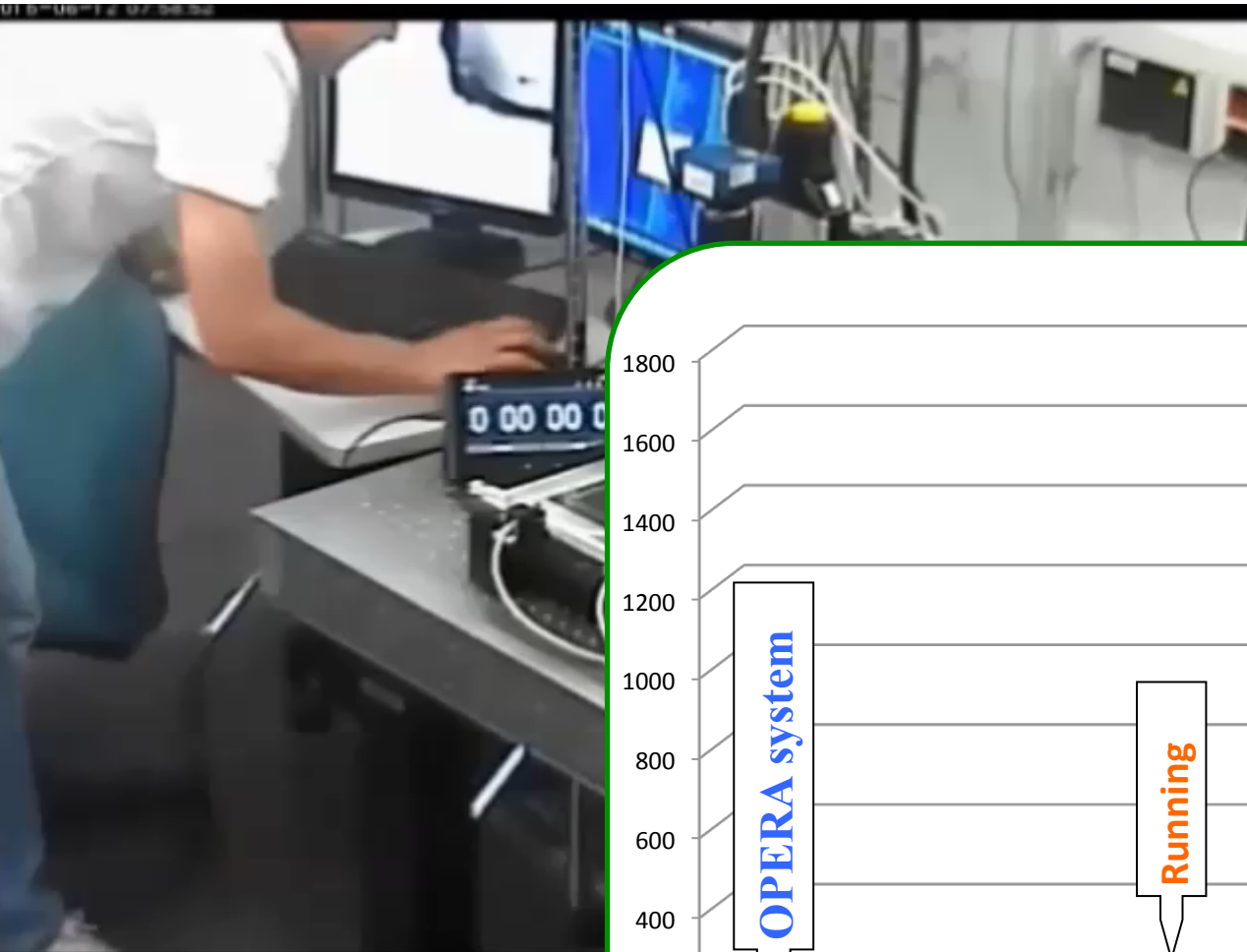
OPERA: 1 event in 1 brick
SHIP: ~ 230 events/brick

Digitizing Nuclear Emulsion Films



← 300 μm →
Grain Density ~ 15 (/45 μm)

IMPROVEMENTS IN THE SCANNING SYSTEM



Scanning speed: $\sim 180 \text{ cm}^2/\text{h}$
 ~ 10 times faster

2011-2014 R&D

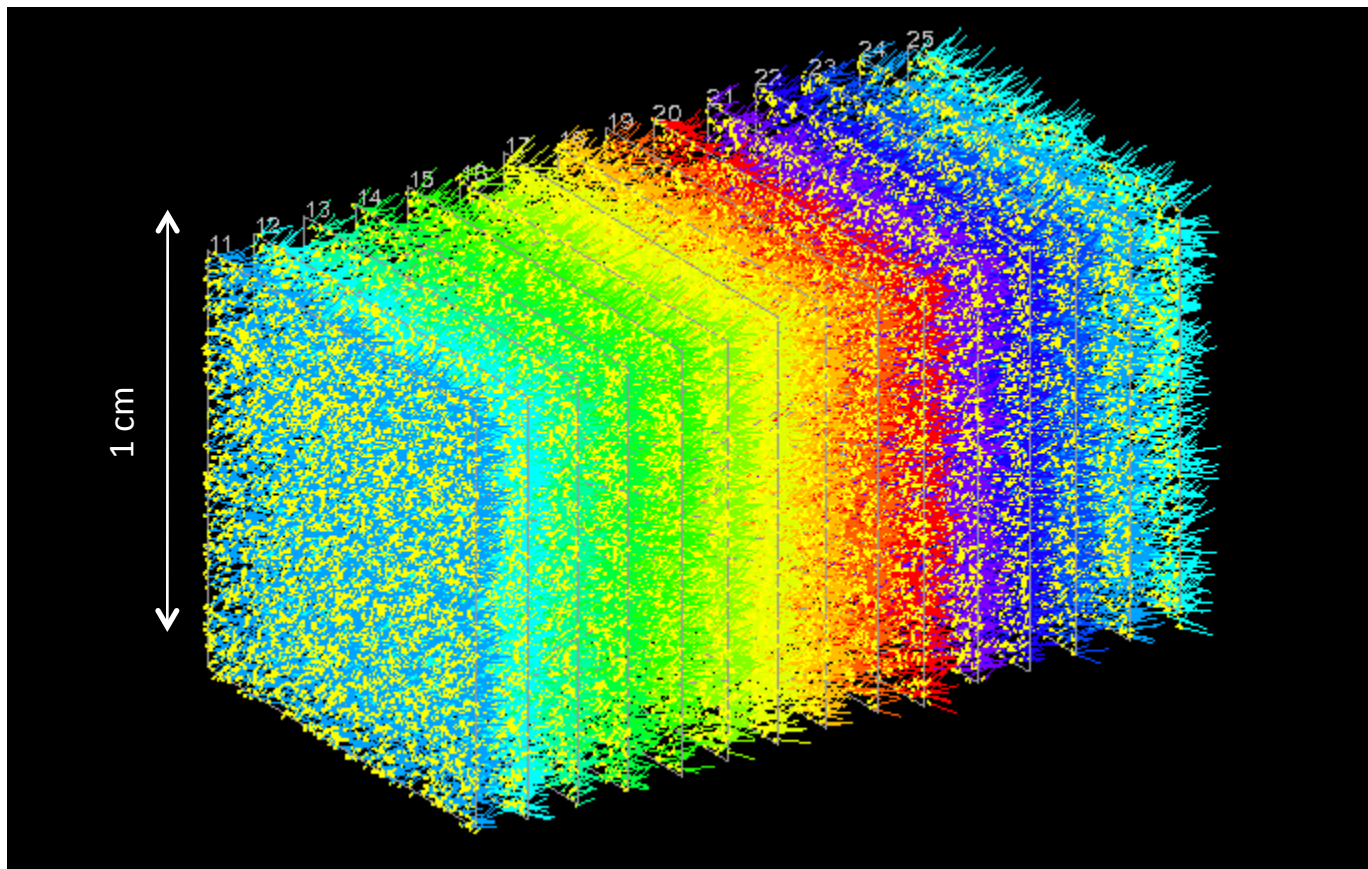
1-2 years

Volume ($\sim 2 \text{ cm}^3$) analysed

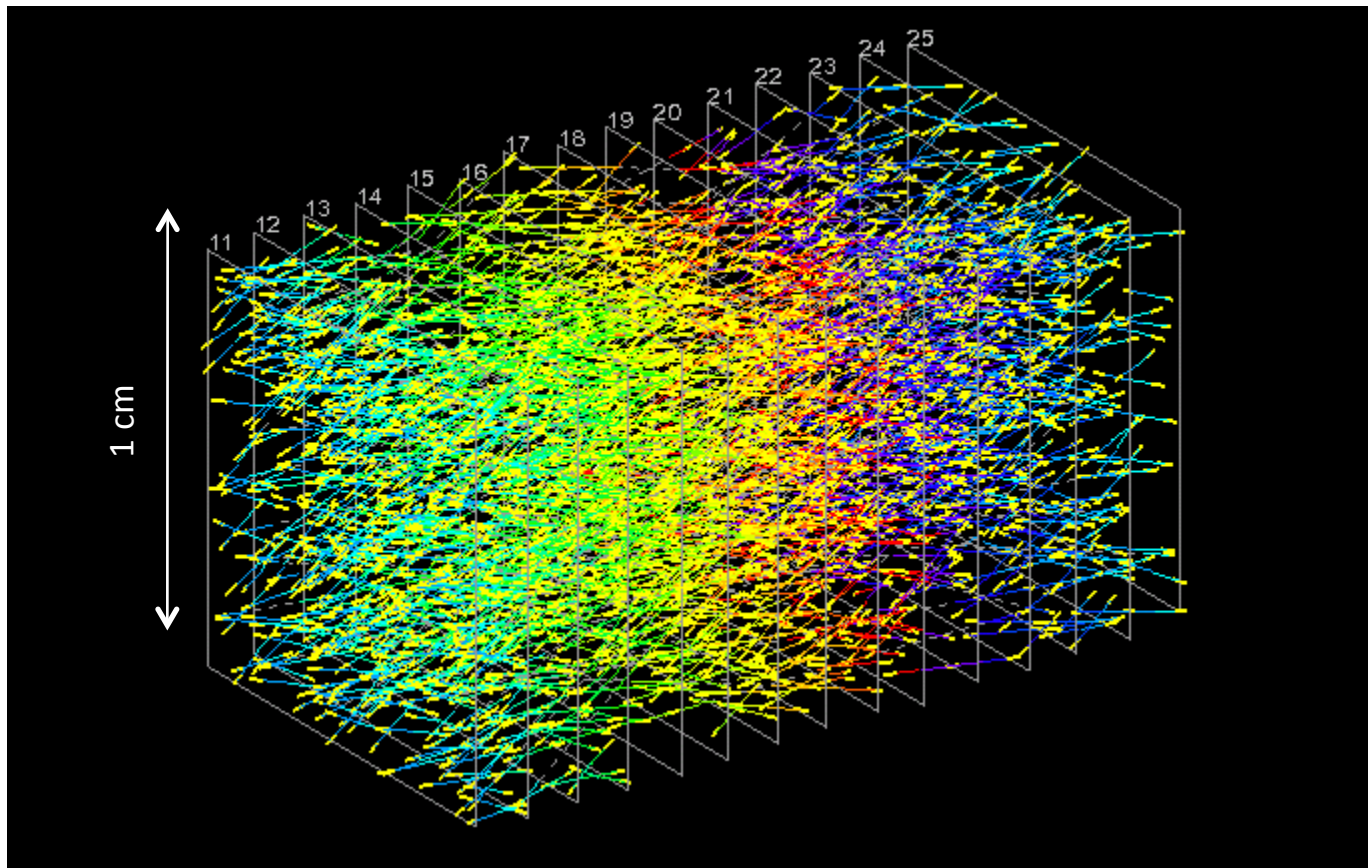
3D tracks with sub-micrometric accuracy

Short Yellow lines \rightarrow measured tracks

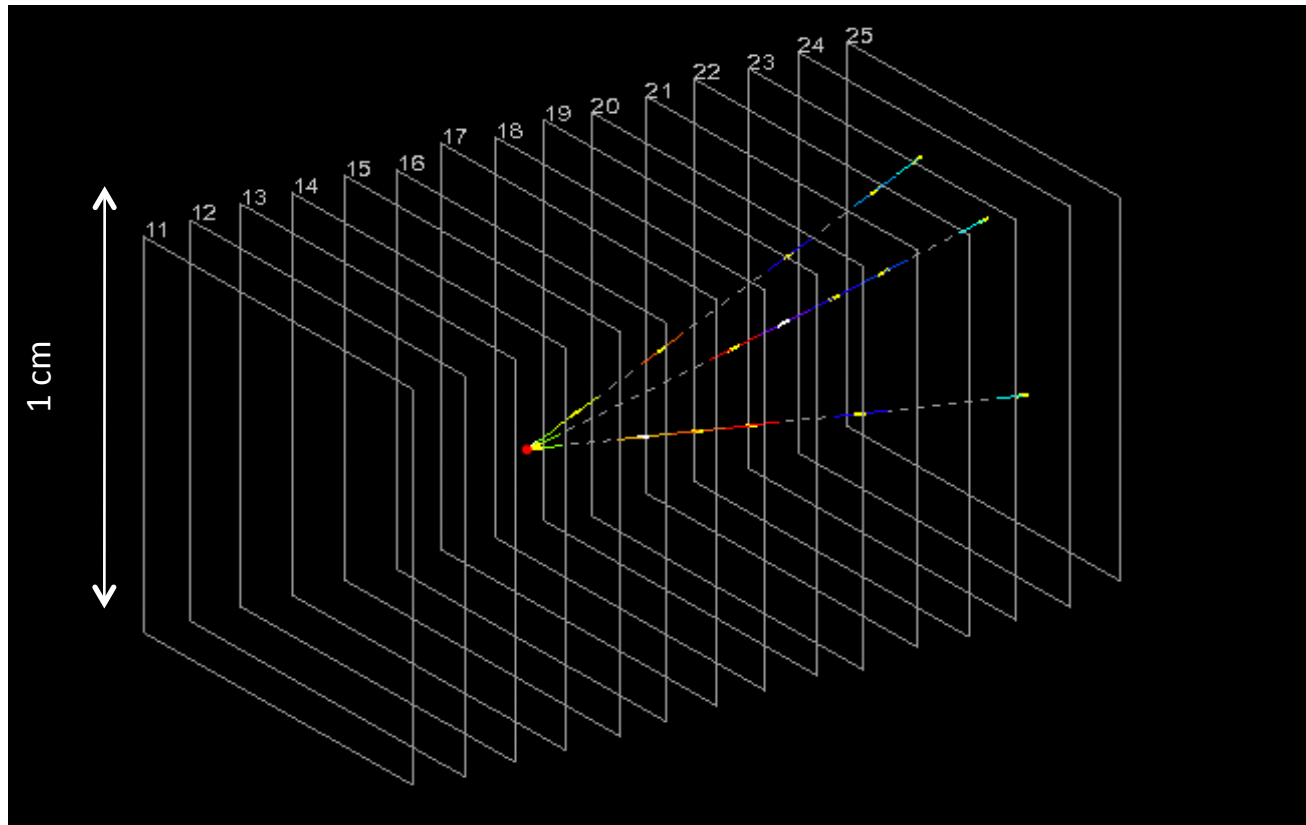
Other colours \rightarrow extrapolated segments



Film to film connection

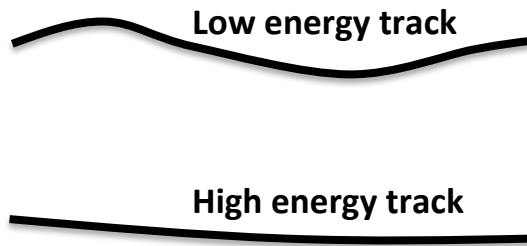
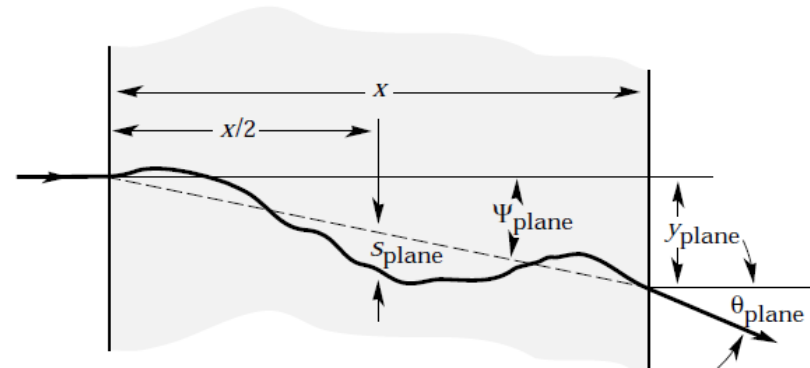


LOCATED NEUTRINO INTERACTION



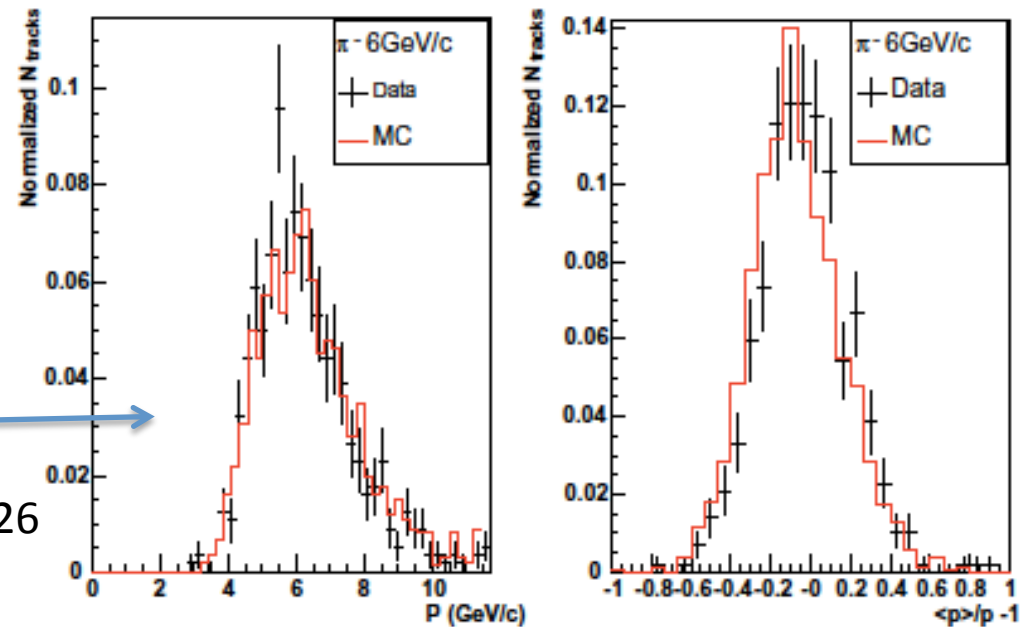
Momentum measurement by multiple Coulomb scattering

$$\theta_{plane}^{rms} = \frac{14.64 MeV}{\beta c p} \sqrt{\frac{x}{X_0}} \left(1 + 0.038 \ln \frac{x}{X_0} \right)$$



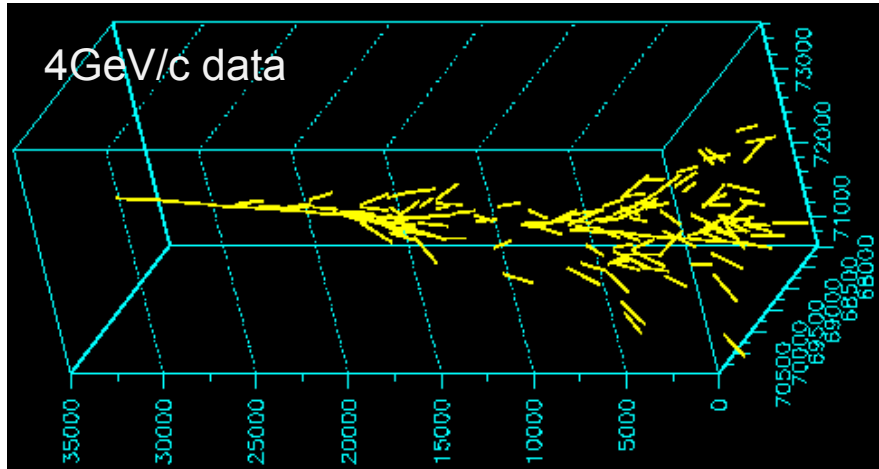
6 GeV π data/MC comparison

New Journal of Physics 14 (2012) 013026

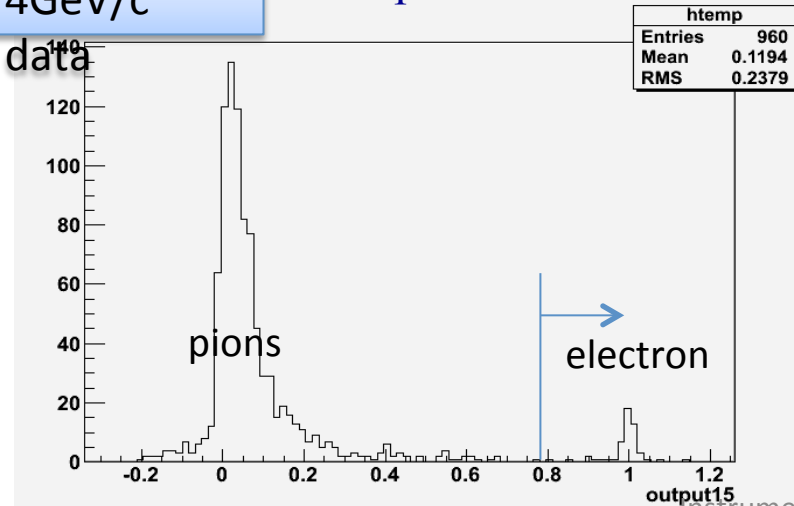


Electromagnetic shower analysis

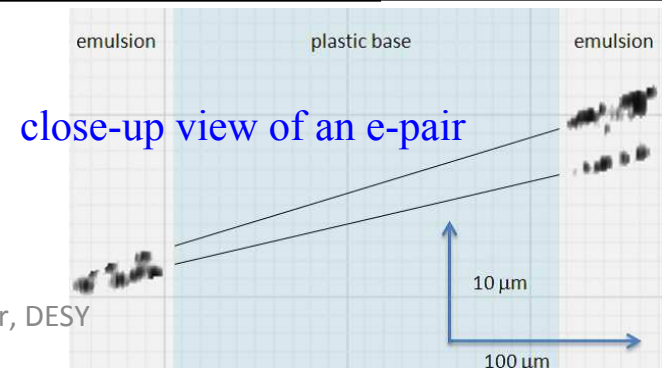
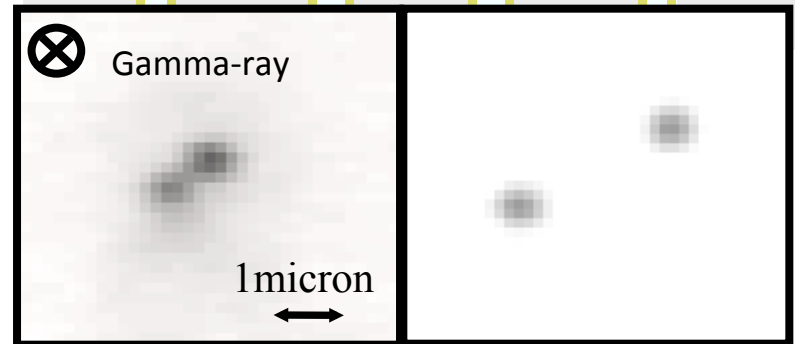
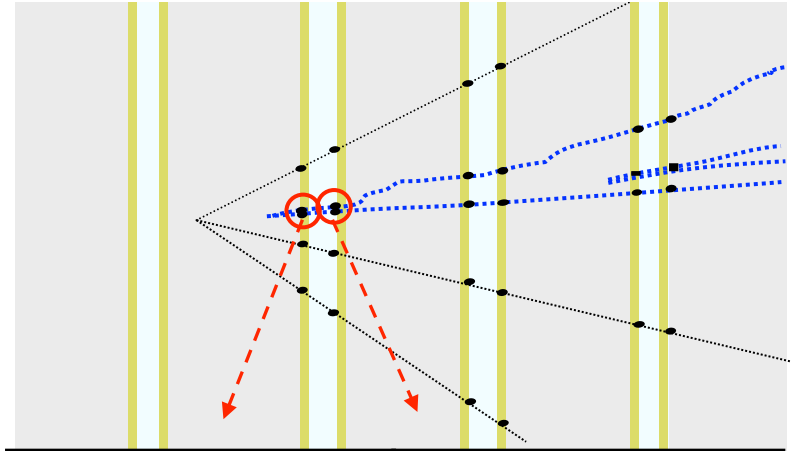
Shower reconstruction



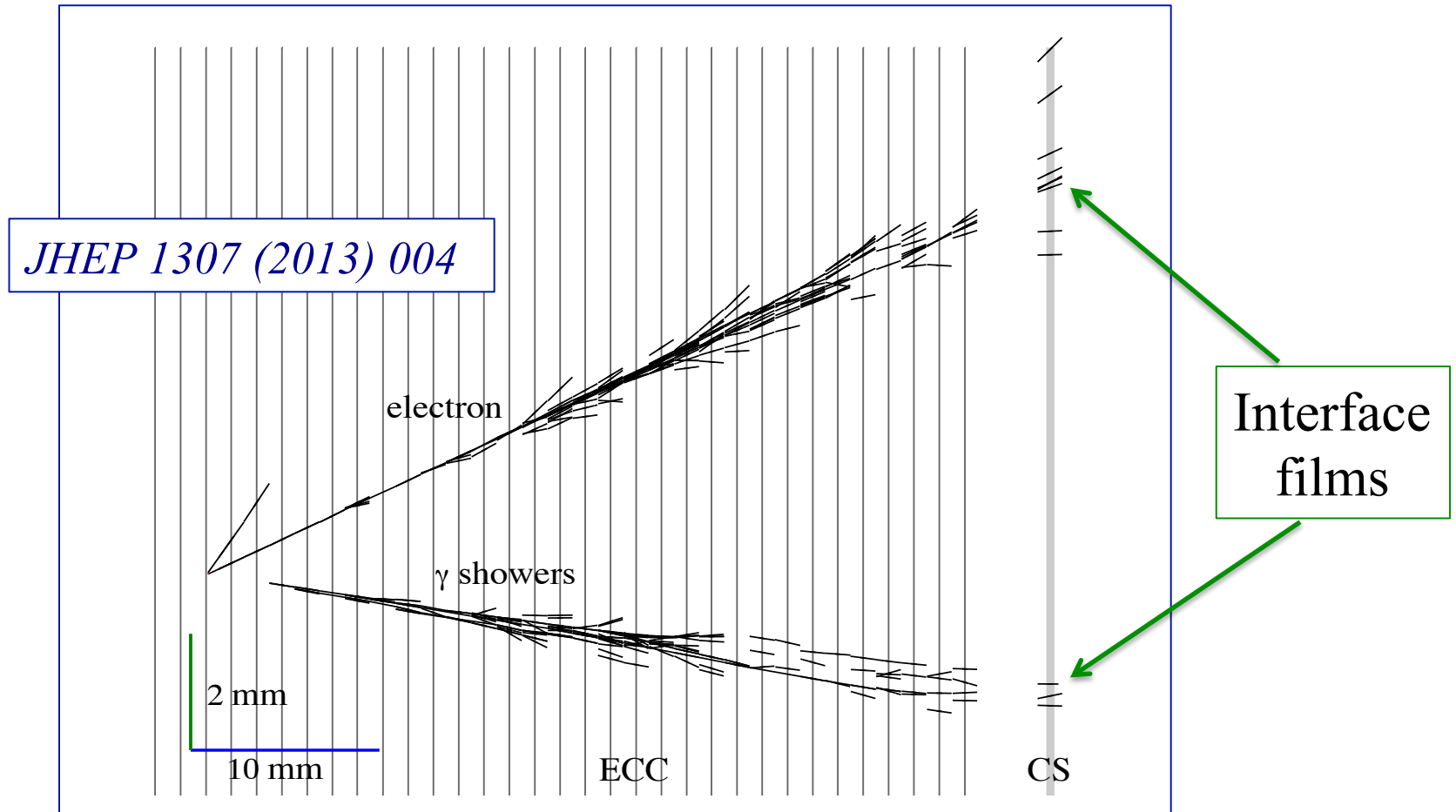
4GeV/c e/π separation



Electron/ γ separation by e-pair detection



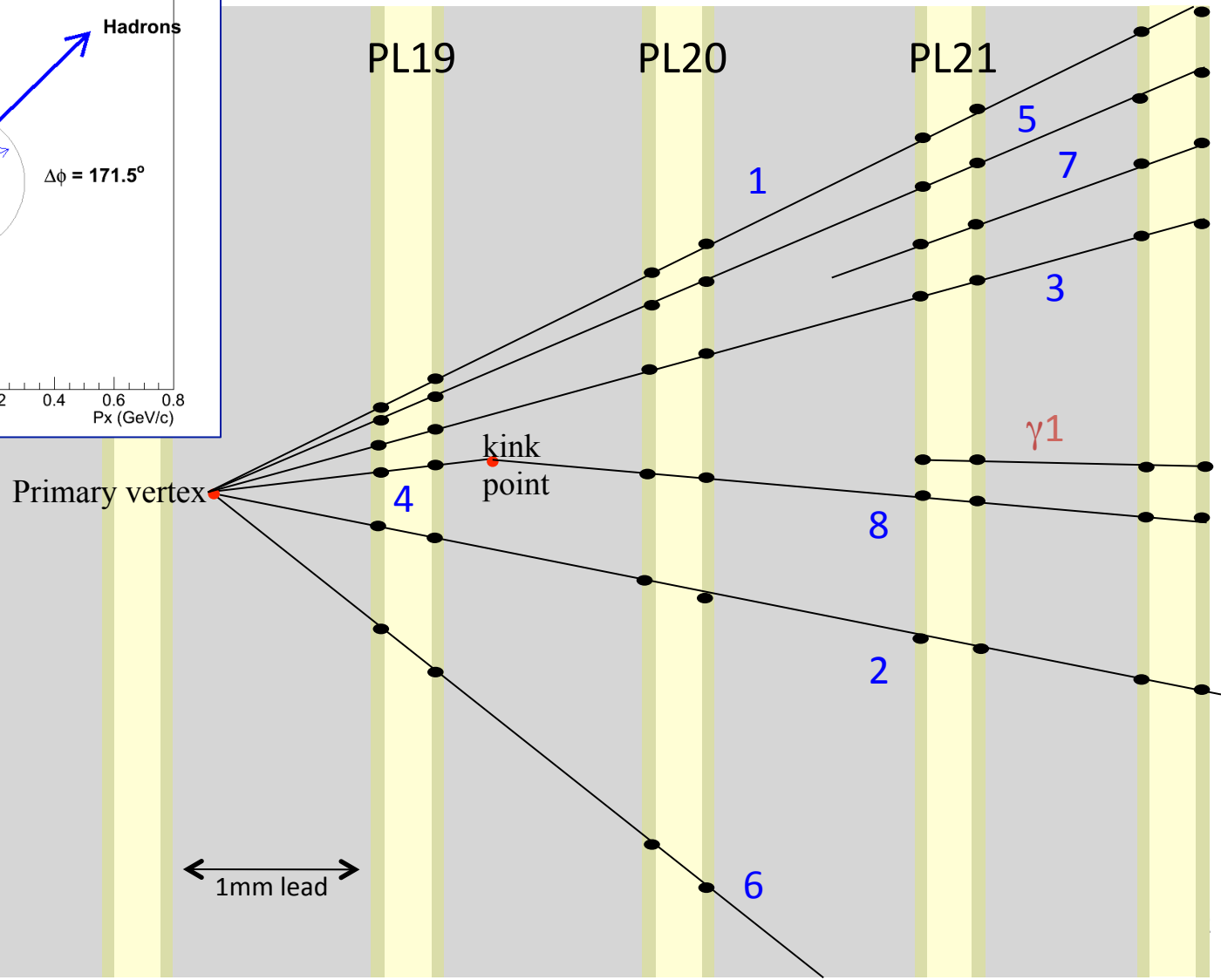
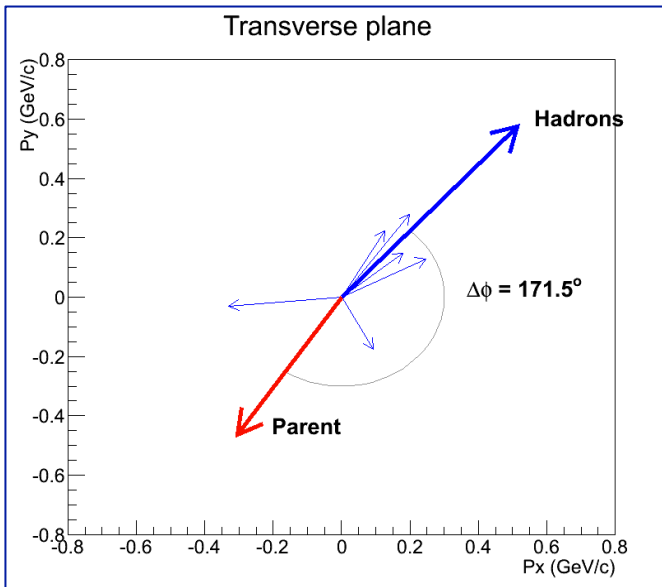
ν_e INTERACTION DETECTED IN AN OPERA BRICK



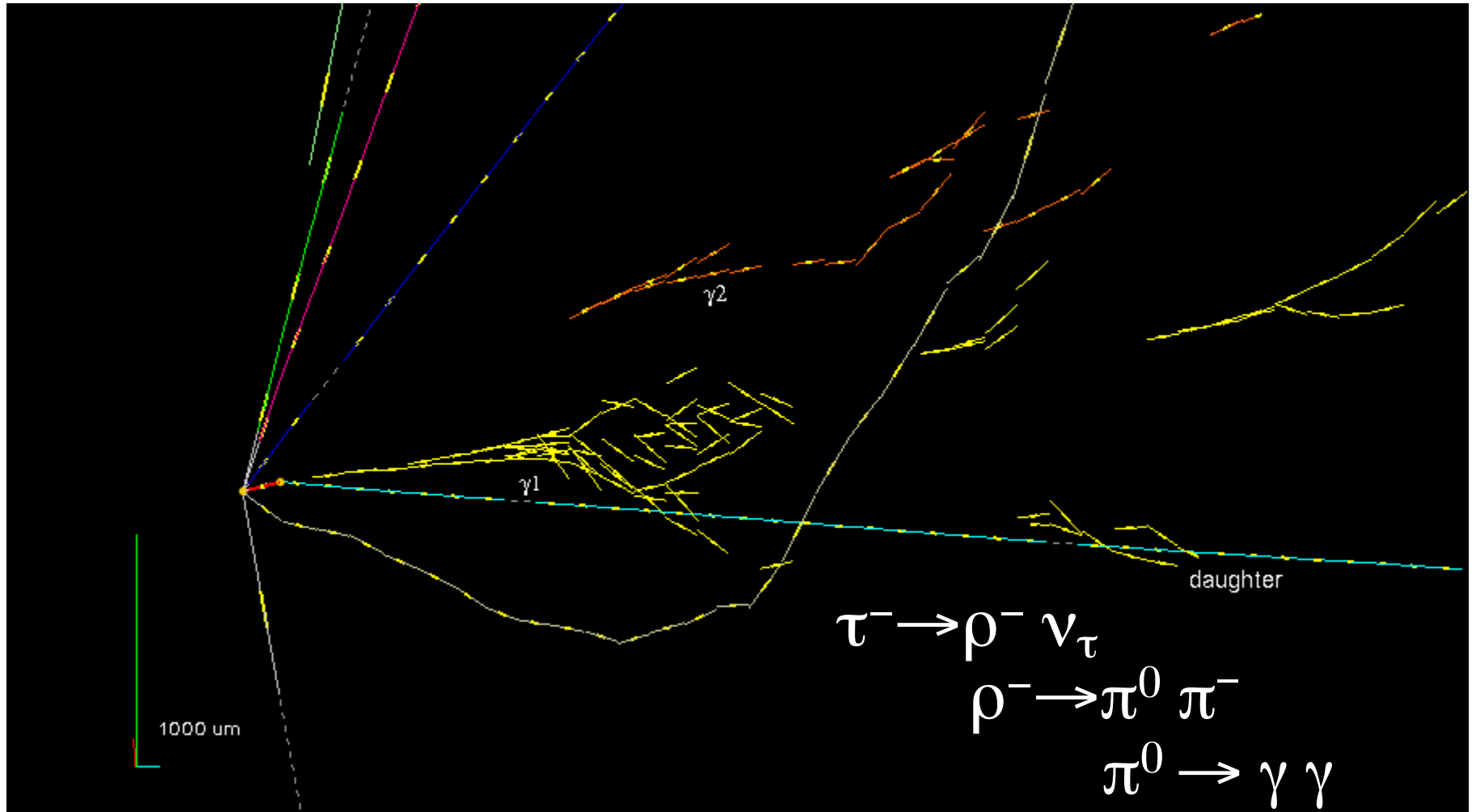
a π^0 is produced at the primary interaction vertex and a γ is detected

ONE OF THE OPERA ν_τ CANDIDATES

Physics Letters B691 (2010) 138



THE FIRST OPERA ν_τ CANDIDATE



ν_τ /ANTI- ν_τ SEPARATION

THE COMPACT EMULSION SPECTROMETER

Magnetised target \rightarrow charge and momentum measurement for hadrons

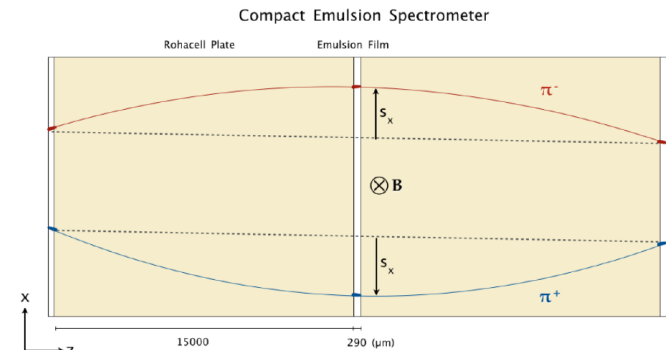
$\text{BR}(\tau \rightarrow \text{hadrons}) \sim 65\%$

Use Compact Emulsion Spectrometer (CES) \rightarrow R&D going on

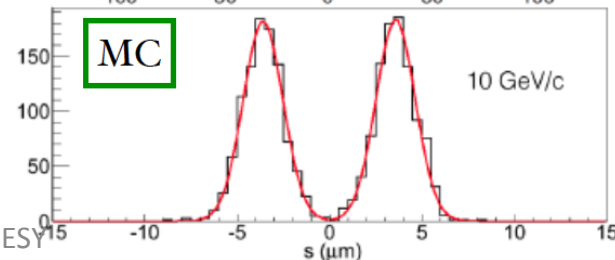
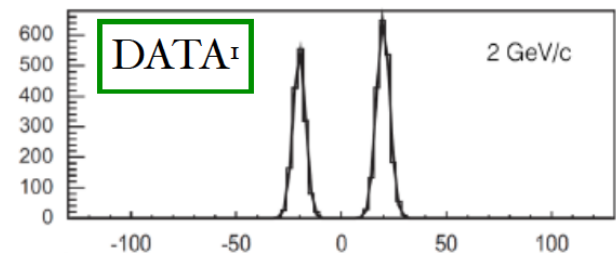
- 1T field
- 3 films interleaved with 2 Rohacell layers (15 mm)
- Thin chamber: 3cm in total
- 90% efficiency for hadronic τ daughters reaching the CES
- Sagitta to discriminate between positive and negative charge

Performances to be achieved

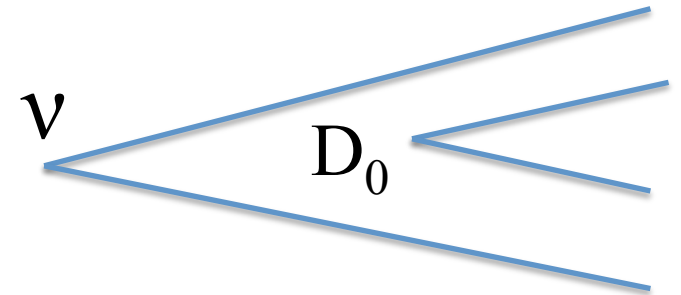
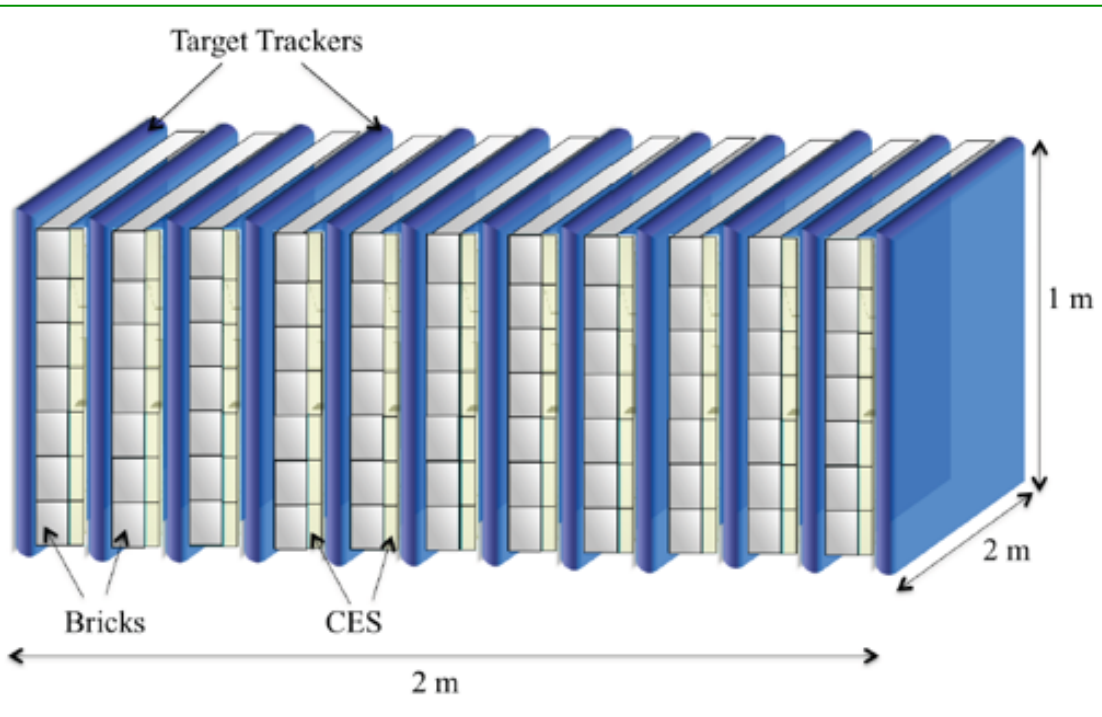
- charge measured up to 10 GeV/c (3 sigma level)
- $\Delta p/p < 20\%$ up to 12 GeV/c



NIM A 592 (2008) 56–62



THE TARGET TRACKER



- 12 target tracker (TT) planes interleaving the 11 brick walls
- first TT plane used as veto
- Transverse size $\sim 2 \times 1 \text{ m}^2$

FEATURES

- Provide time stamp
- Link muon track information from the target to the magnetic spectrometer

REQUIREMENTS

- Operate in 1T field
- X-Y position resolution $< 100 \mu\text{m}$
- high efficiency ($>99\%$) for angles up to 1 rad

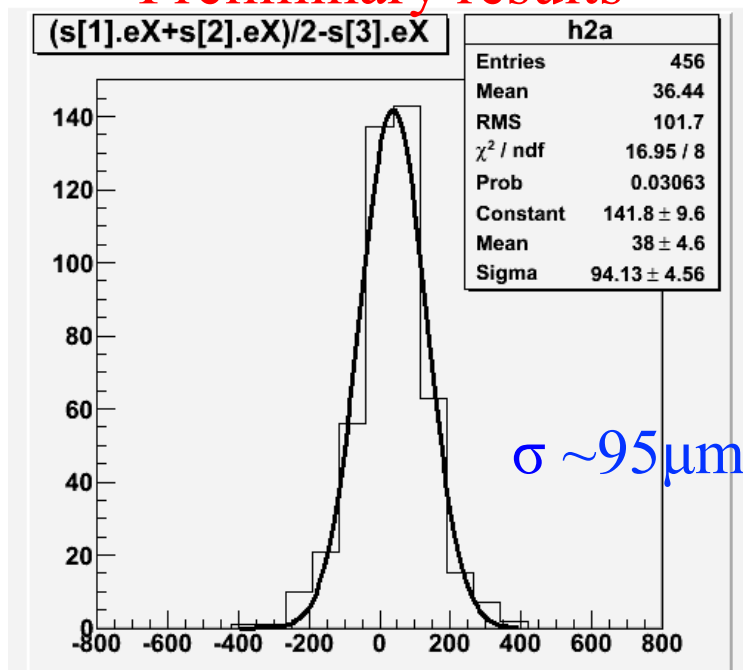
TARGET TRACKER PLANES

POSSIBLE OPTIONS

- Scintillating fibre trackers
- Micro-pattern gas detectors (GEM, Micromegas)

Testing some options

Preliminary results

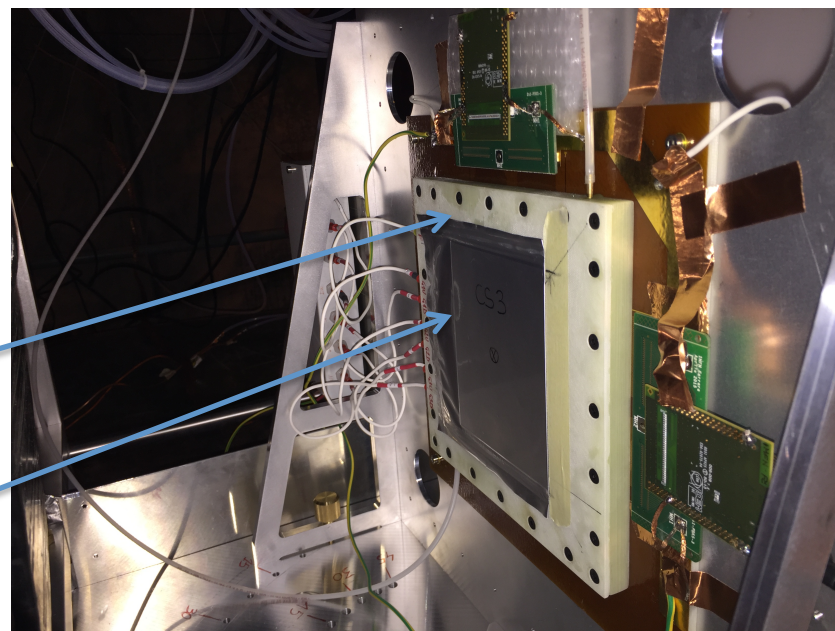
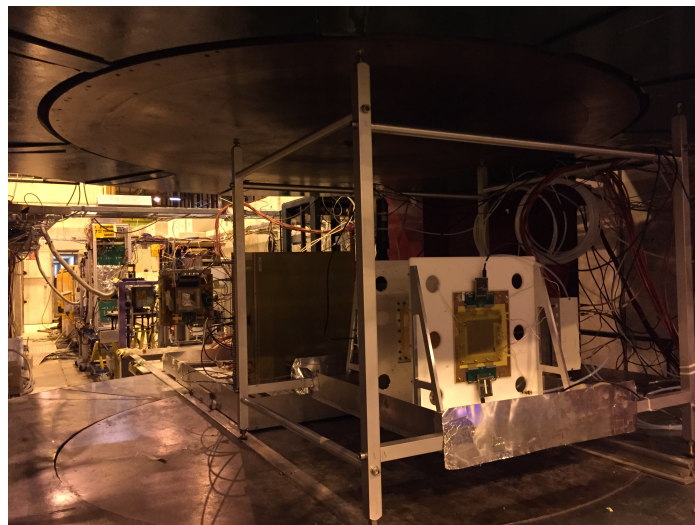


Perpendicular tracks with magnetic field off

Magnetic field and angular
effect being studied

GEM

Emulsion films



SPECTROMETER MAGNET



TOP (IRON) YOKE

TOP (COPPER) COIL

(IRON) SLABS

BOTTOM (COPPER) COIL

BOTTOM (IRON) YOKE

Isometric view
Scale: 1:50

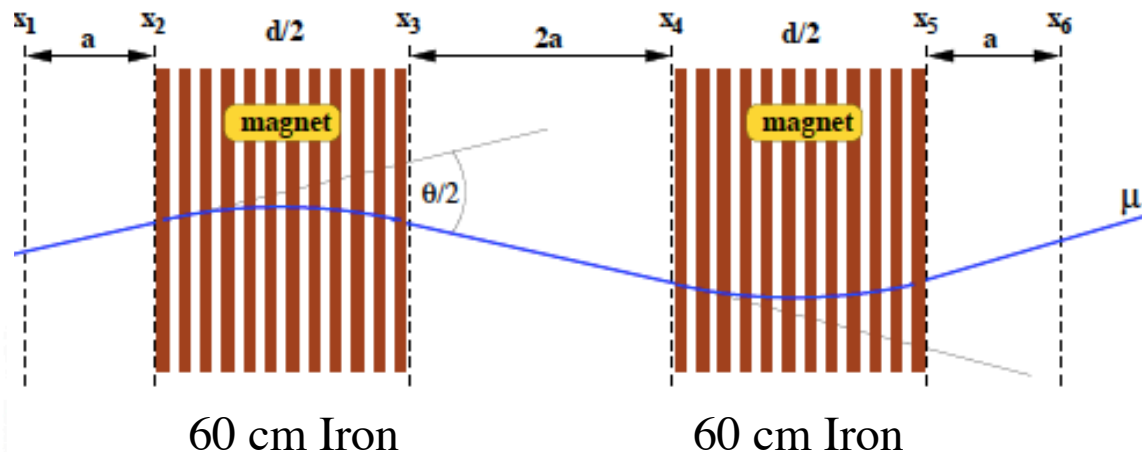
8200

4000

2840

Total WEIGHT 454 Ton

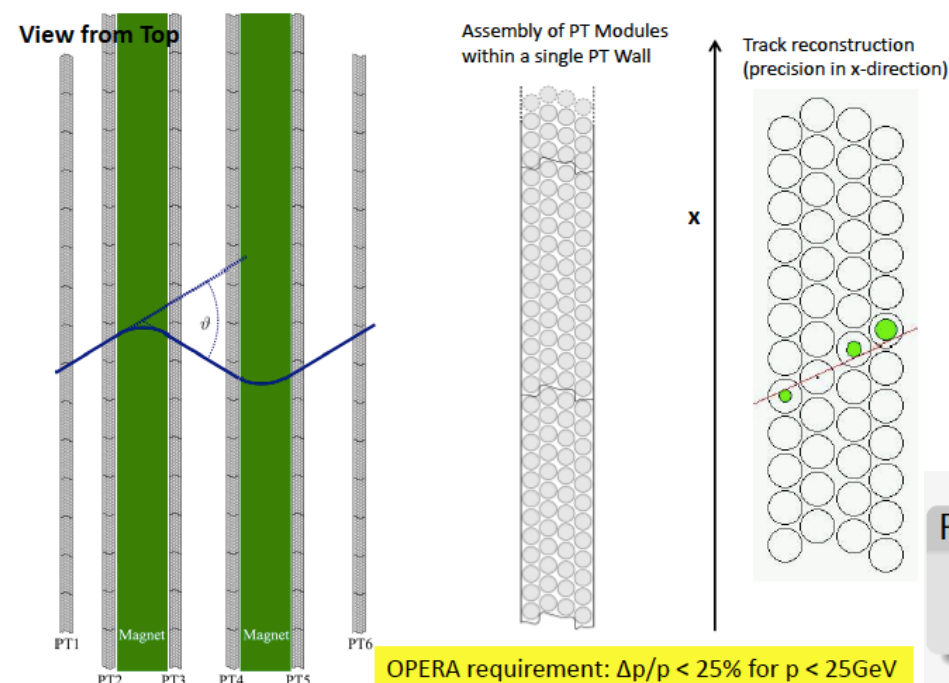
Tracking stations inside the magnet



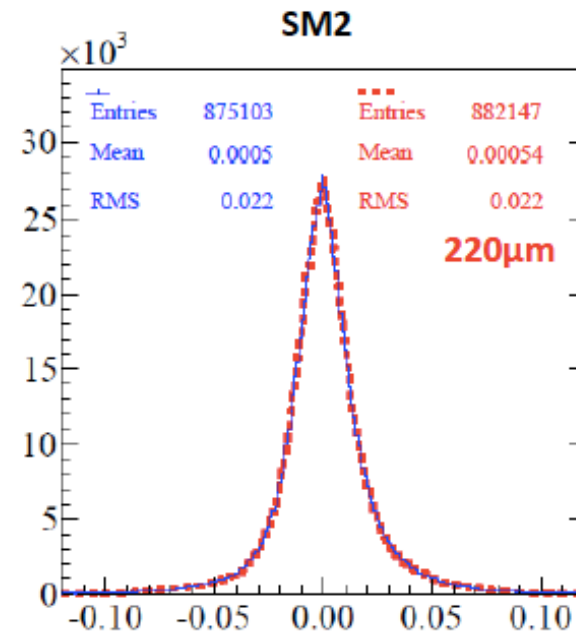
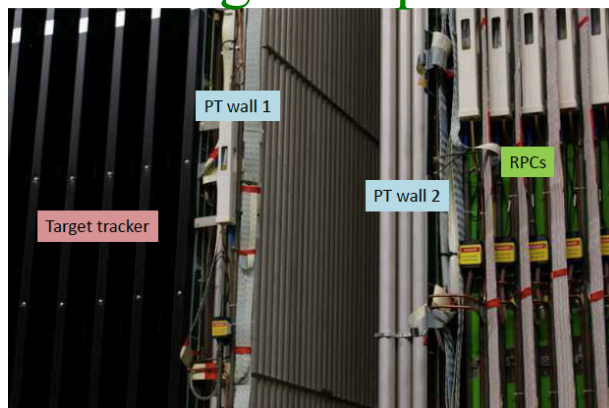
- Requirements:
 - coarse (1 cm) tracking inside the magnetised volume
 - 1ns time resolution
 - Muon rate $\sim 5\text{kHz/m}^2$ rate
 - Electron rate ~ 1 order higher
- RPC's technology is one option
- Streamer versus avalanche to be studied



Muon momentum measurement and identification

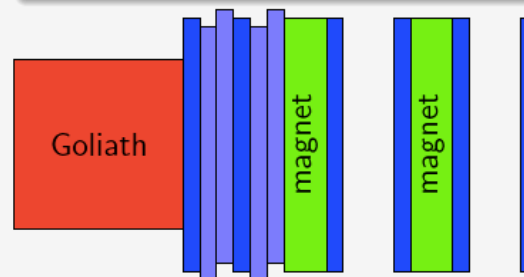


OPERA drift tubes are
a good option

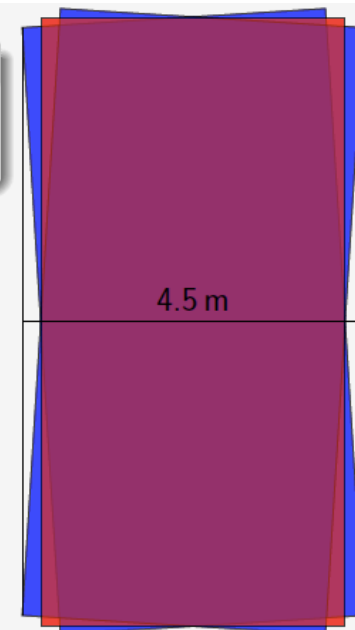


Requirements for 3d reconstruction

- Stereo angle between planes
- 3 projections to avoid ambiguities



- 2×3 projections
- 2 projections rotated by $\pm 3.6^\circ$
- maximal width: 4.5 m





Neutrino detector performances

✓ *Unique capability of detecting all three neutrino flavours*

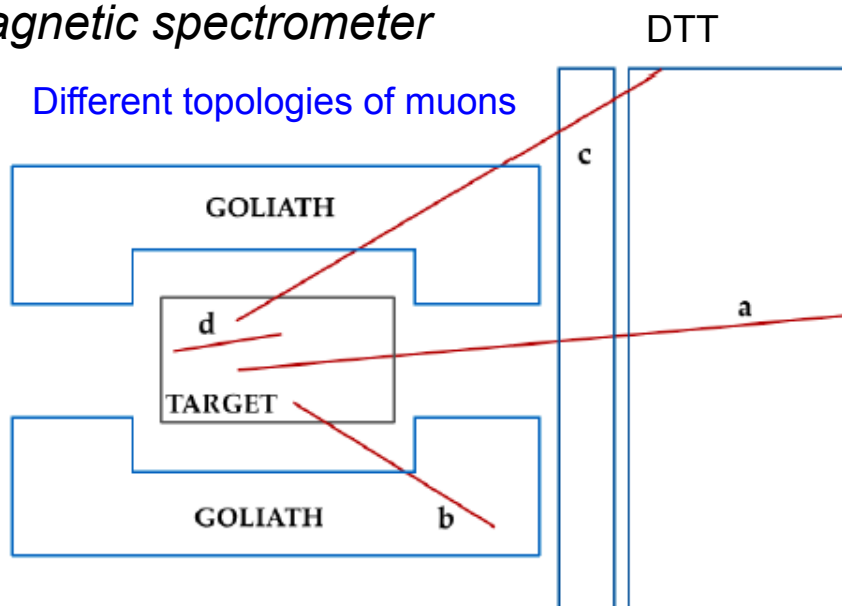
- $\nu_\tau / \bar{\nu}_\tau \rightarrow \nu$ interaction and τ decay vertices in emulsion target
- $\nu_e \rightarrow$ electrons producing em shower in emulsion target
- $\nu_\mu \rightarrow$ muons identified by TT, DTT and the muon spectrometer of the tau neutrino detector

	ϵ_{tot} (%)
$\tau \rightarrow \mu X$	60
$\tau \rightarrow hX$	62
$\tau \rightarrow 3hX$	63
$\tau \rightarrow eX$	56

✓ *Separation between tau and anti tau-neutrinos by the charge measurement*

- charge of hadrons is measured by CES
- charge of muons is measured by CES and magnetic spectrometer

	$\tau \rightarrow hX$	$\tau \rightarrow 3hX$	$\tau \rightarrow \mu X$
Correct charge	70%	49%	94%
Wrong charge	0.5%	1.0%	1.5%

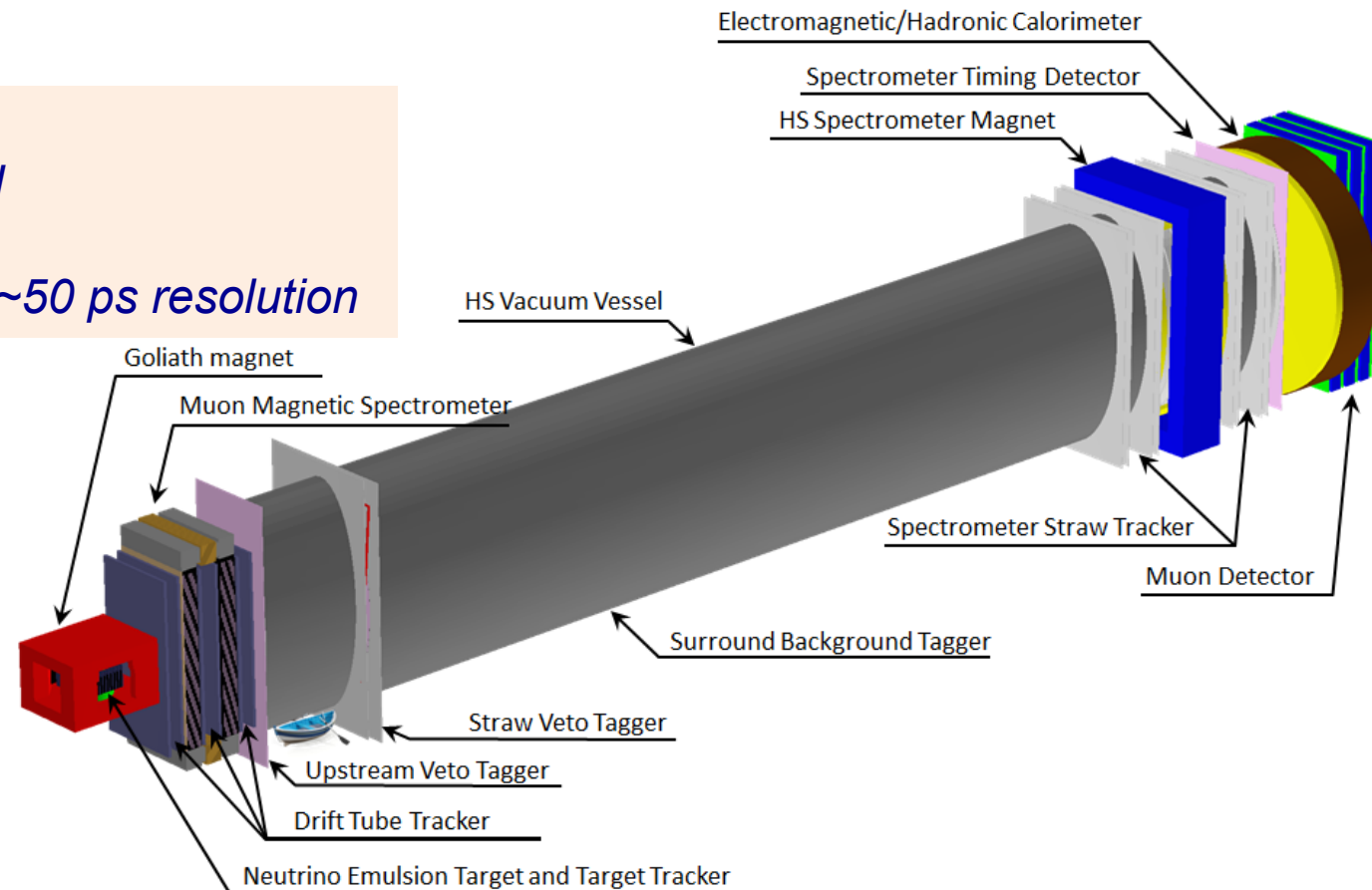


Hidden Sector detector concept

- ✓ *Reconstruction of HS decays in all possible final states*
Long decay volume protected by various Veto Taggers, Magnetic Spectrometer followed by the Timing Detector, and Calorimeters and Muon systems.
All heavy infrastructure is at distance to reduce neutrino / muon interactions in proximity of the detector

Challenges:

- Large vacuum vessel
- 5 m long straw tubes
- Timing detector with ~ 50 ps resolution

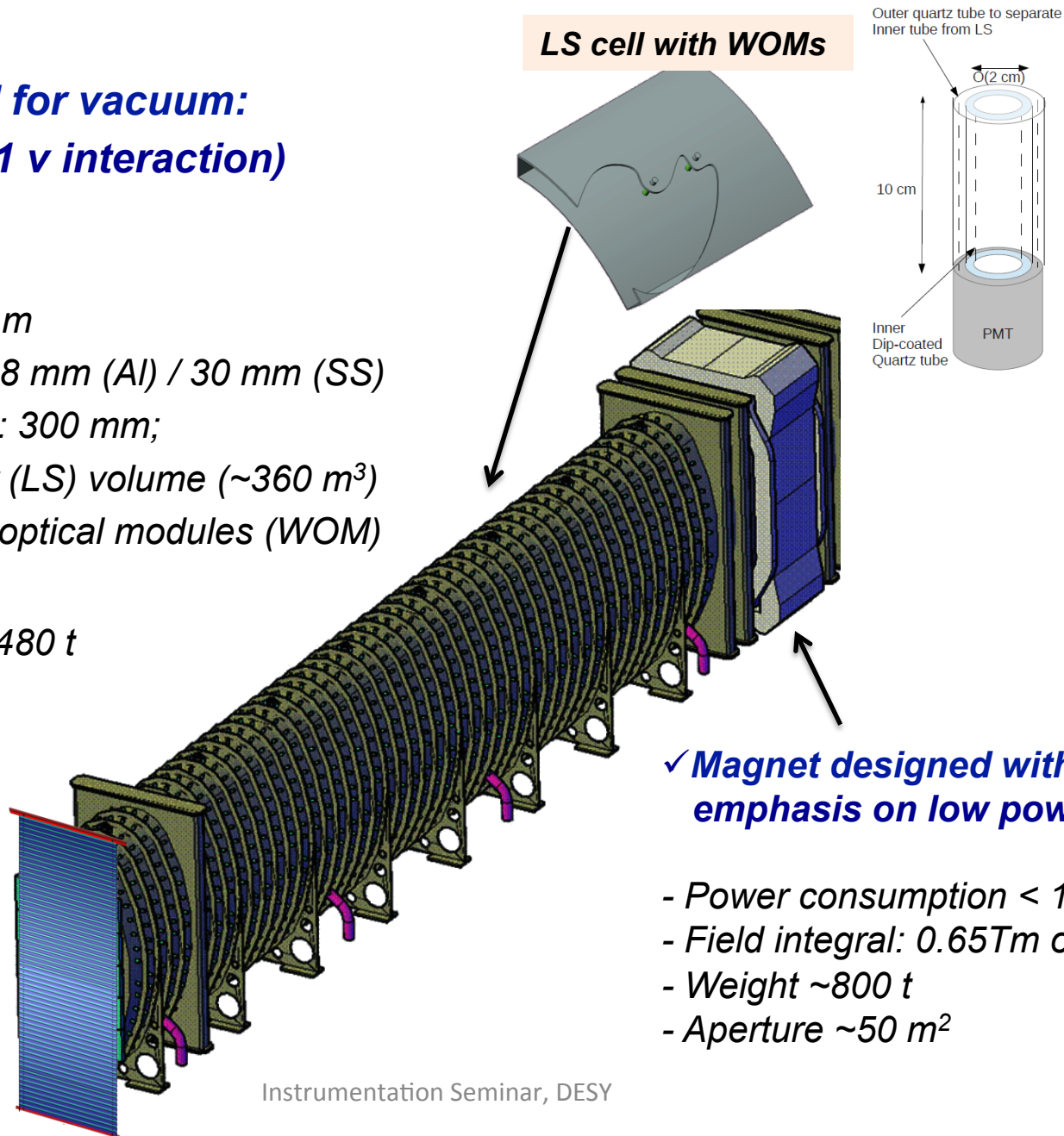


HS decay volume and spectrometer magnet

✓ **Estimated need for vacuum:**
 $\sim 10^{-3}$ mbar (<1 v interaction)

✓ **Vacuum vessel**

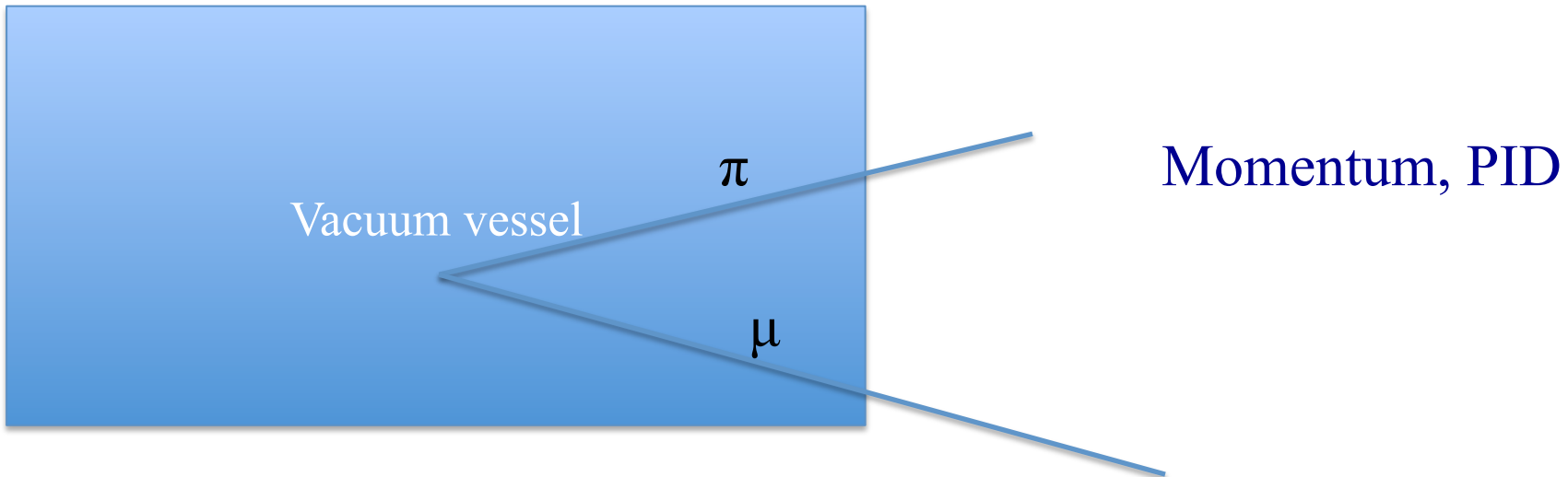
- 10 m x 5 m x 60 m
- Walls thickness: 8 mm (Al) / 30 mm (SS)
- Walls separation: 300 mm;
- Liquid scintillator (LS) volume (~ 360 m³) readout by WLS optical modules (WOM) and PMTs
- Vessel weight ~ 480 t



✓ **Magnet designed with an emphasis on low power**

- Power consumption < 1 MW
- Field integral: 0.65Tm over 5m
- Weight ~ 800 t
- Aperture ~ 50 m²

Signal features



- Main background: neutrino interactions
- Reduce this background by:
 - IP cut
 - Invariant mass
- Important to
 - Measure precisely the momentum
 - Identification the particle
- Reduce combinatorial background by precise timing



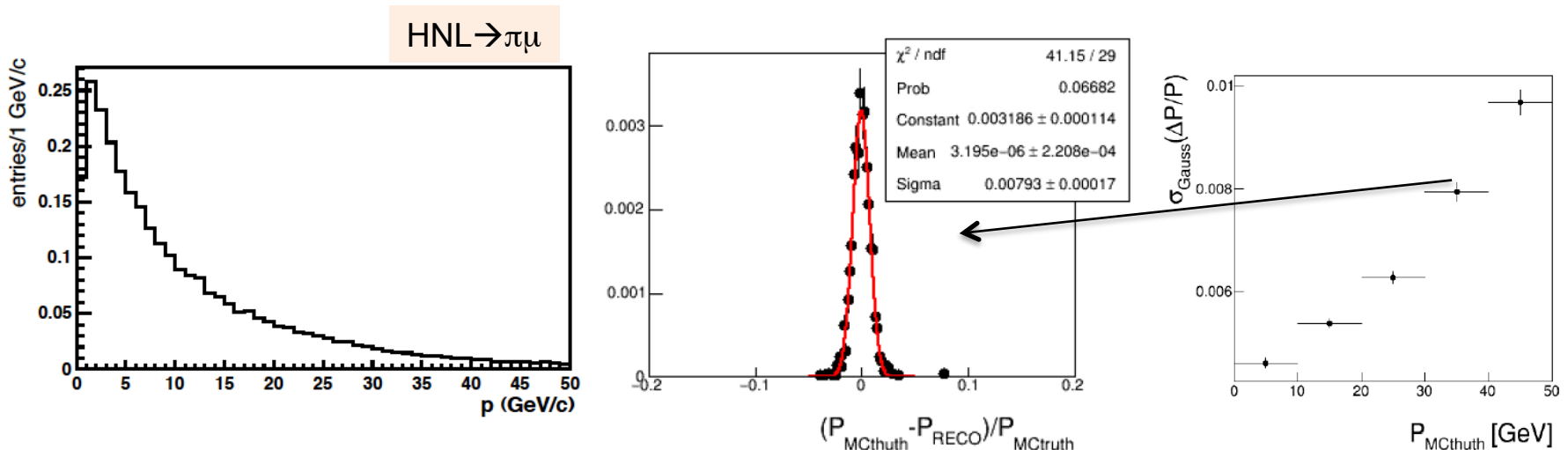
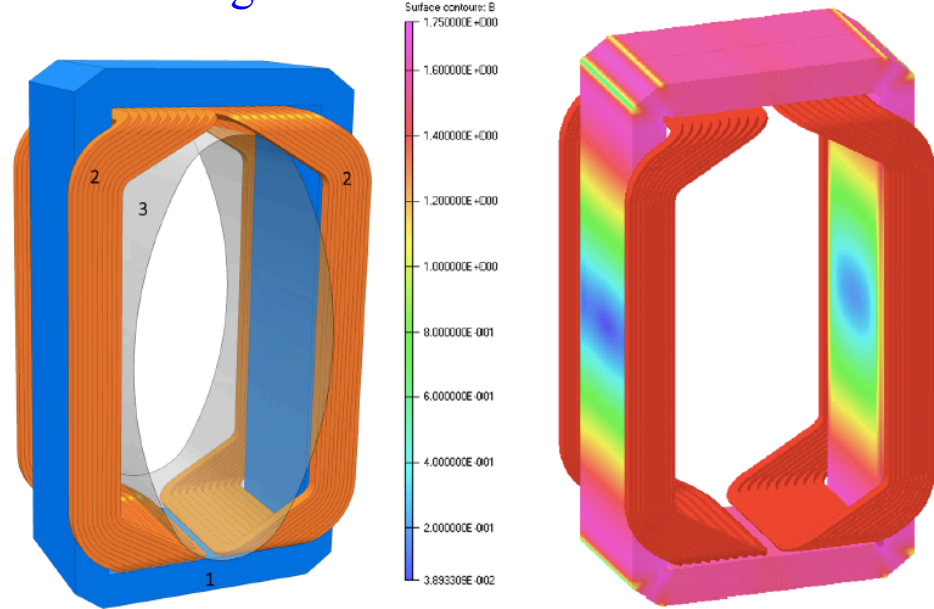
Momentum resolution of the HS (straw tubes) tracker

Magnet with vacuum vessel

- material budget per station $0.5\% X_0$
- position resolution $120 \mu\text{m}$ per straw, 8 hits per station on average

$$\left(\frac{\sigma_p}{p}\right)^2 \approx [0.49\%]^2 + [0.022\% / (\text{GeV}/c)]^2 \cdot p^2$$

Momentum resolution is dominated by multiple scattering below 22 GeV/c
(For $\text{HNL} \rightarrow \pi\mu$, 75% of both decay products have $P < 20 \text{ GeV}/c$)

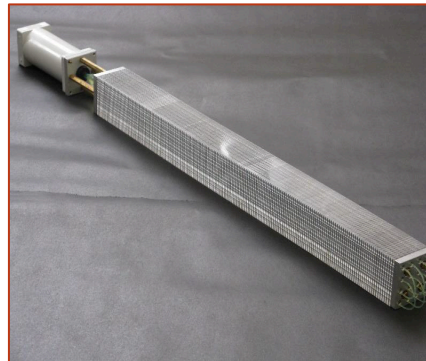
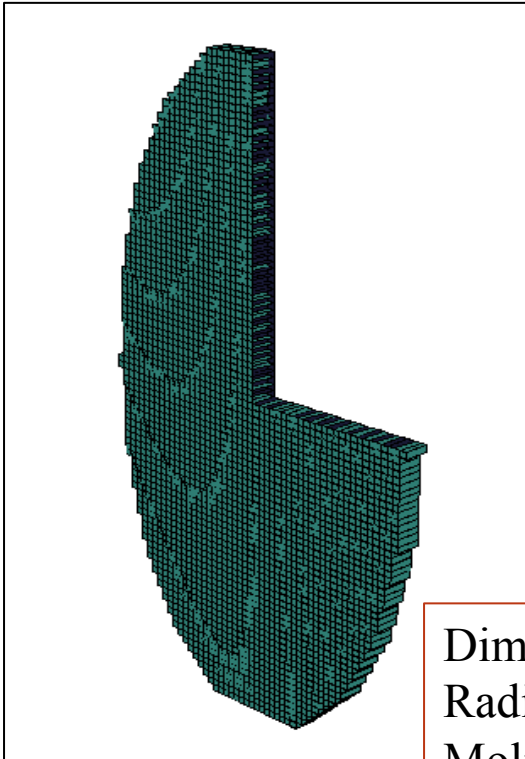


Vertex resolution (also driven by multiple scattering and $\Delta P/P$):

$$\sigma_{xy} \sim \mathcal{O}(\text{mm}), \sigma_z \sim \mathcal{O}(\text{cm})$$

ECAL

- Almost elliptical shape (5 m x 10 m)
- 2876 Shashlik modules
- 2x2 cells/modules, width=6 cm
- 11504 independent readout channels

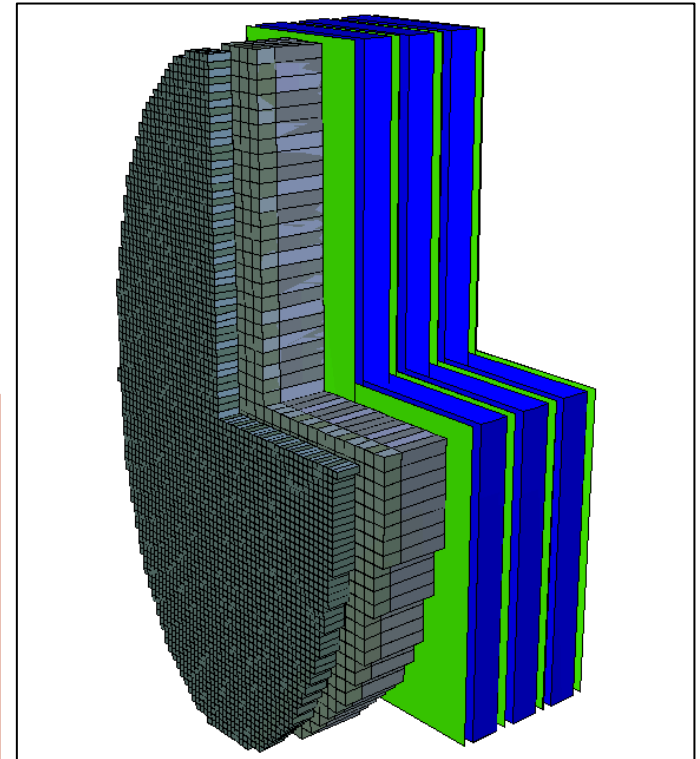


Dimensions	60x60 mm ²
Radiation length	17 mm
Moliere radius	36 mm
Radiation thickness	25 X ₀
Scintillator thickness	1.5 mm
Lead thickness	0.8 mm
Energy resolution	1%

Calorimeters

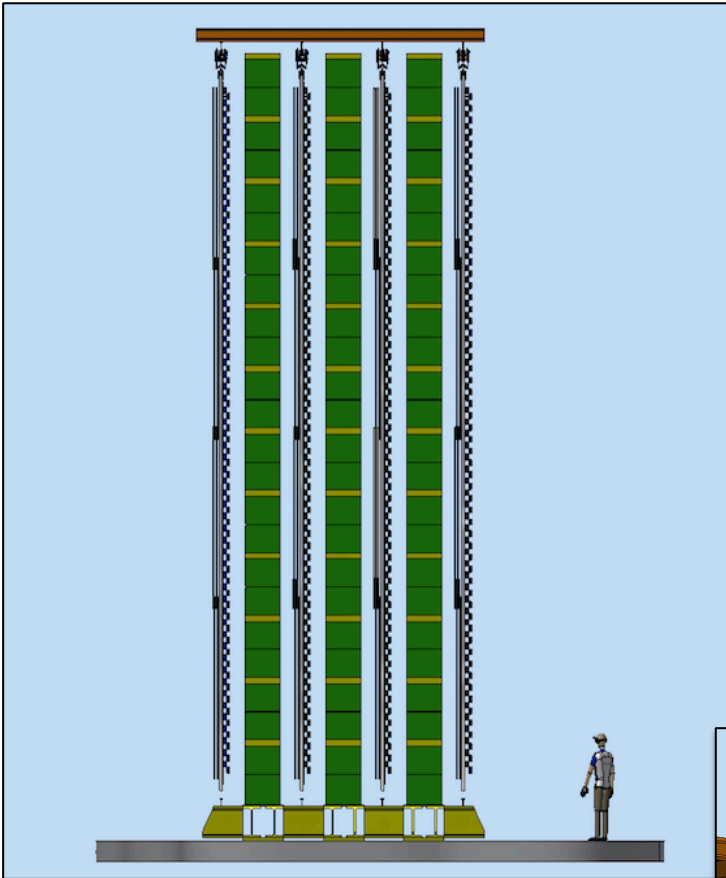
HCAL

- Matched with ECAL acceptance
- 2 stations
- 5 m x 10 m
- 1512 modules
- 24x24 cm² dimensions
- Stratigraphy: N x (1.5 cm steel+0.5 cm scint)
- 1512 independent readout channels



Muon System

Based on scintillating bars, with WLS fibers and SiPM readout

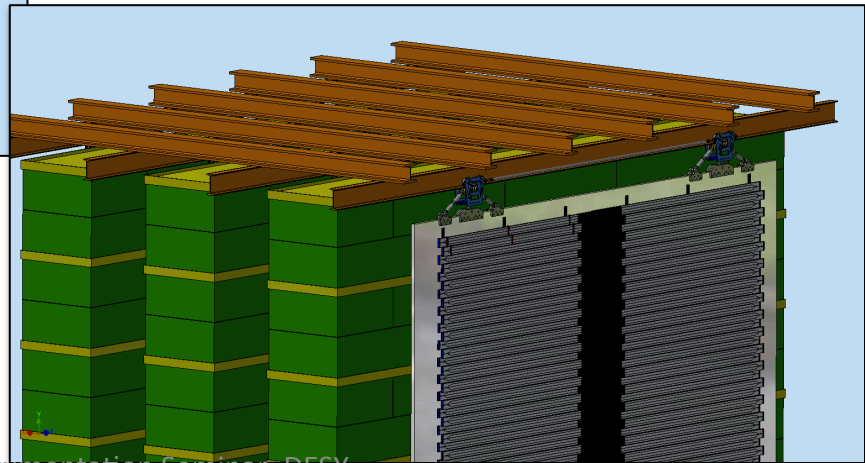


Requirements:

- High-efficiency identification of muons in the final state
- Separation between muons and hadrons/electrons
- Complement timing detector to reject combinatorial muon background

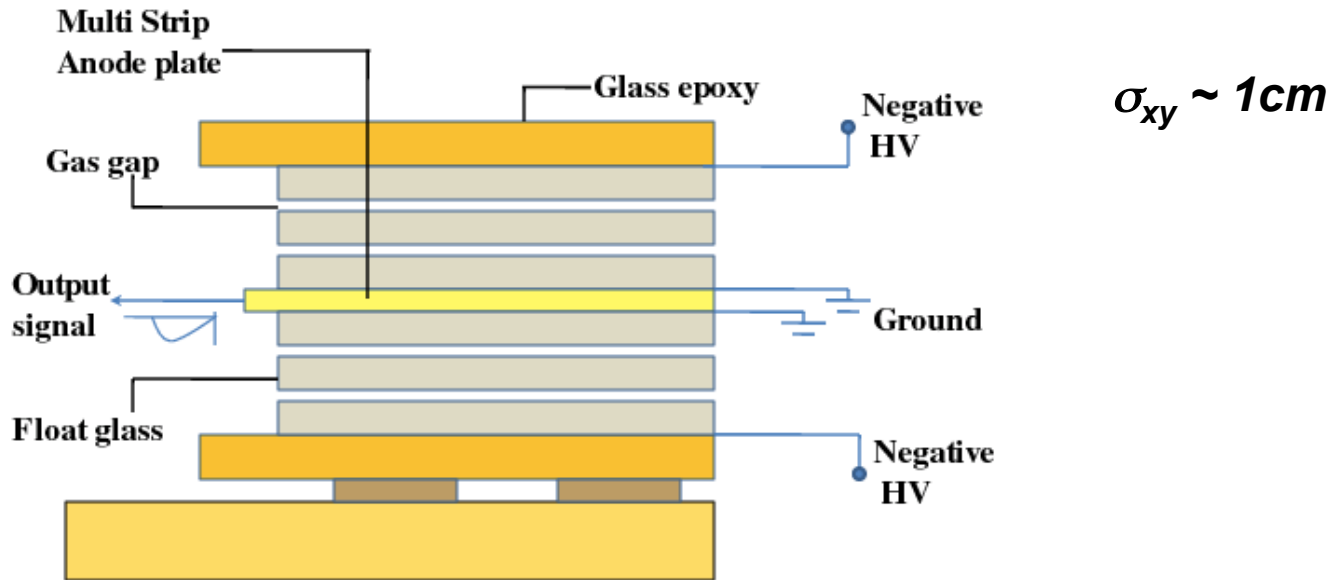
Technical Proposal (preliminary design)

- 4 active stations
- transverse dimensions: $1200 \times 600 \text{ cm}^2$
- x,y view
- 3380 bars, $5 \times 300 \times 2 \text{ cm}^3$ /each
- 7760 FEE channels
- 1000 tons of iron filters



Timing detector ($< 100\text{ps}$)

Multi-gap RPC is one option





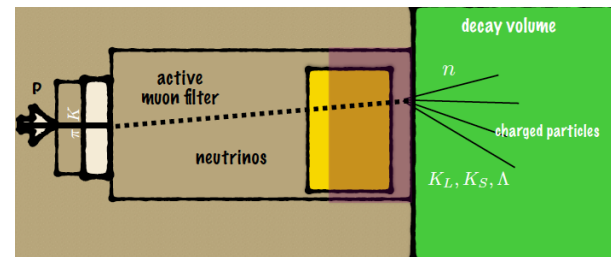
HS Backgrounds (1)

Main sources of background

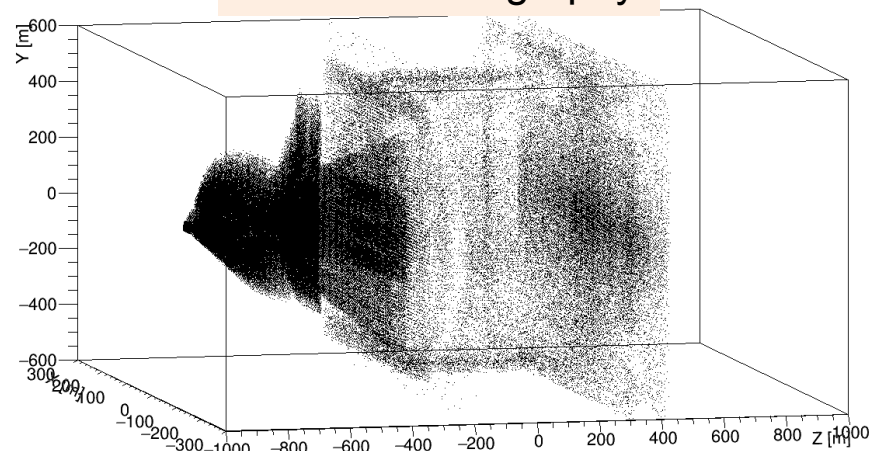
- ✓ **Neutrino DIS interactions with material in the vicinity of the HS decay volume**
(interactions of ν with air in the decay volume are negligible at 10^{-3} mbar)

Origin of neutrino interactions

- Walls of the decay volume (>80%)
- Tau neutrino detector
- HS tracking system

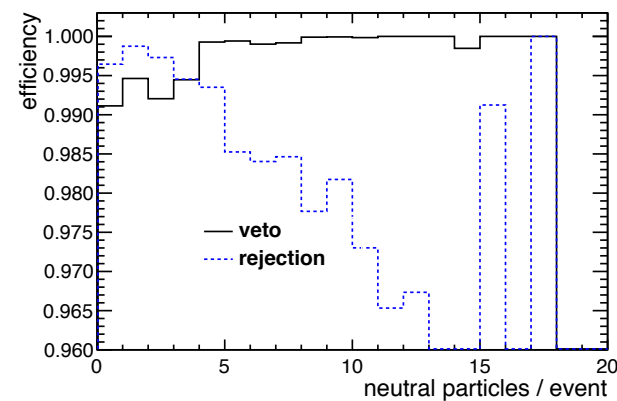
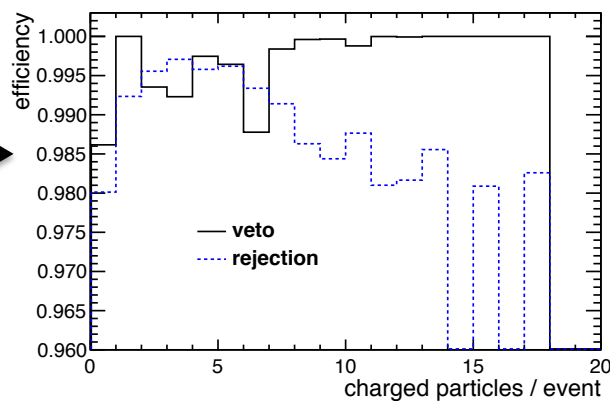


Neutrino tomography



Combination of veto and selection cuts reduces the ν -induced background to zero

Veto efficiency increases with event multiplicity →



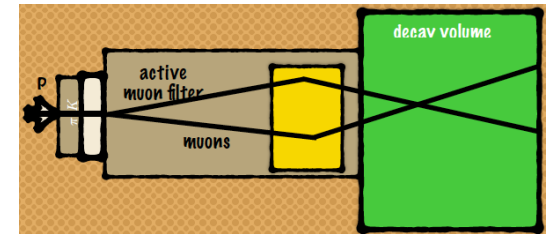
HS Backgrounds (2)

✓ Muon combinatorial background

Simulation predicts $O(10^{12})$ muon pairs in the decay volume in 5 years of data taking

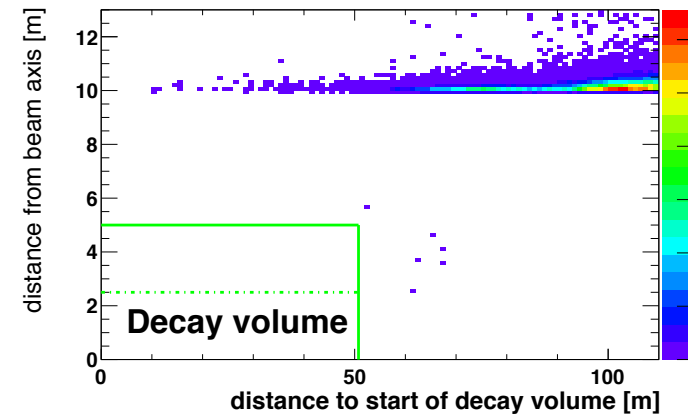
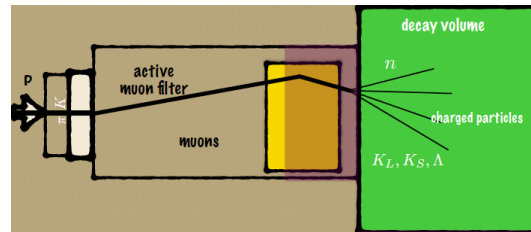
Suppressed by:

- Basic kinematic and topological cuts $\sim 10^4$
- Timing veto detectors $\sim 10^7$
- Upstream veto and surrounding veto taggers $\sim 10^4$



✓ Muon DIS interactions

- V^0 s produced in the walls of the cavern
- DIS close to the entry of the decay volume
→ smaller than neutrino induced background

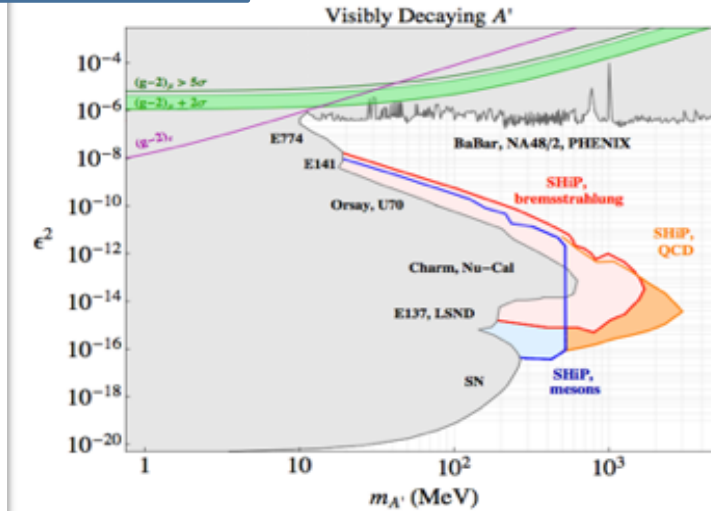


Background summary: no evidence for any irreducible background

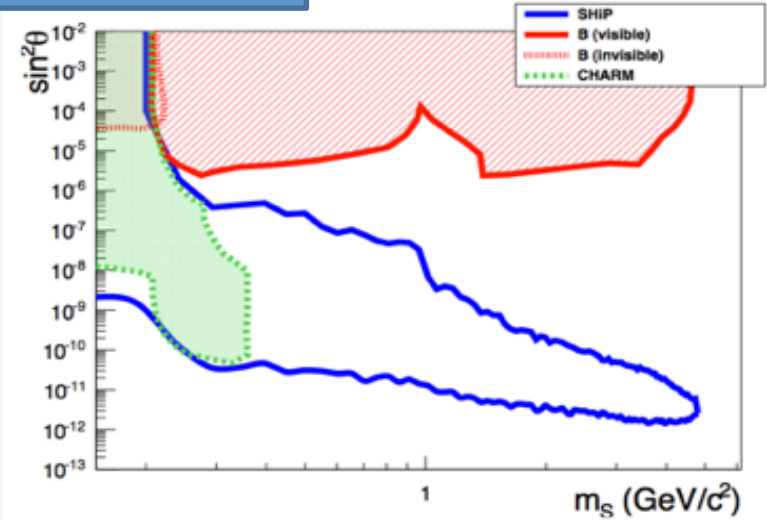
SHiP sensitivity to Hidden Sector

Based on 2×10^{20} pot
@400 GeV in 5 years

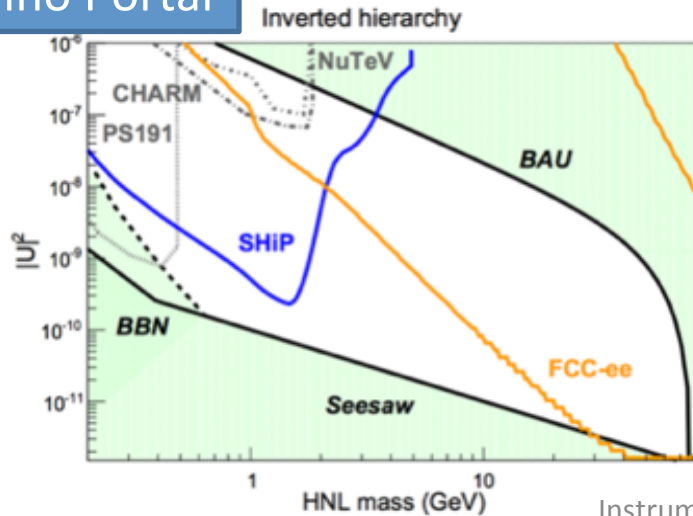
Vector Portal



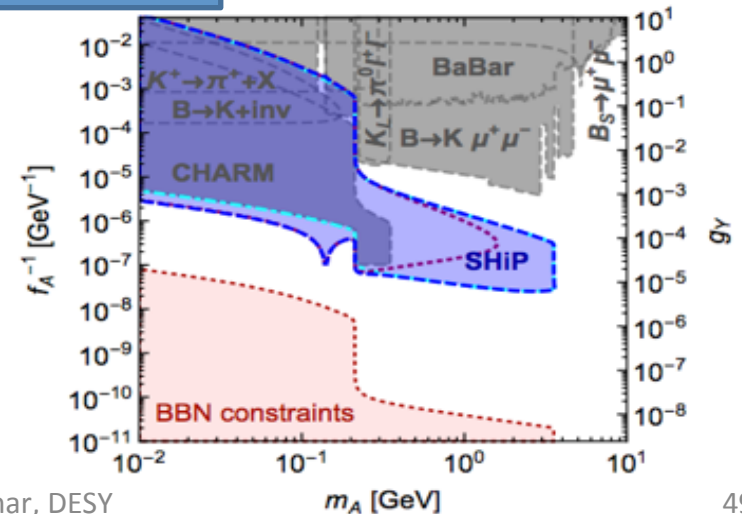
Scalar Portal



Neutrino Portal



Axion Portal

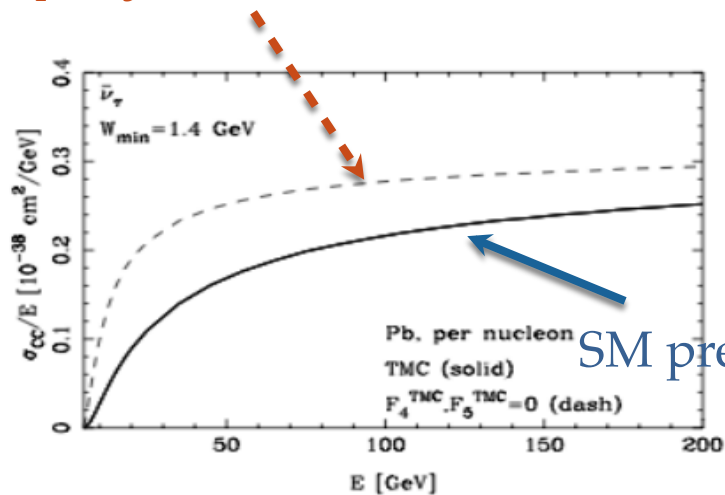


F₄ AND F₅ STRUCTURE FUNCTIONS

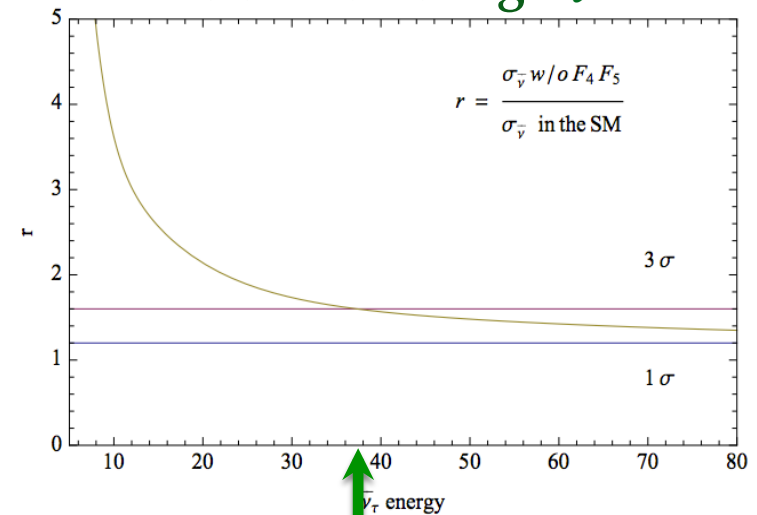
First evaluation of F₄ and F₅, not accessible with other neutrinos

$$\frac{d^2\sigma^{\nu(\bar{\nu})}}{dxdy} = \frac{G_F^2 M E_\nu}{\pi(1 + Q^2/M_W^2)^2} \left((y^2x + \frac{m_\tau^2 y}{2E_\nu M}) F_1 + \left[(1 - \frac{m_\tau^2}{4E_\nu^2}) - (1 + \frac{Mx}{2E_\nu}) \right] F_2 \right. \\ \left. \pm \left[xy(1 - \frac{y}{2}) - \frac{m_\tau^2 y}{4E_\nu M} \right] F_3 + \frac{m_\tau^2(m_\tau^2 + Q^2)}{4E_\nu^2 M^2 x} F_4 - \frac{m_\tau^2}{E_\nu M} F_5 \right),$$

F₄ = F₅ = 0



CC interacting $\bar{\nu}_\tau$

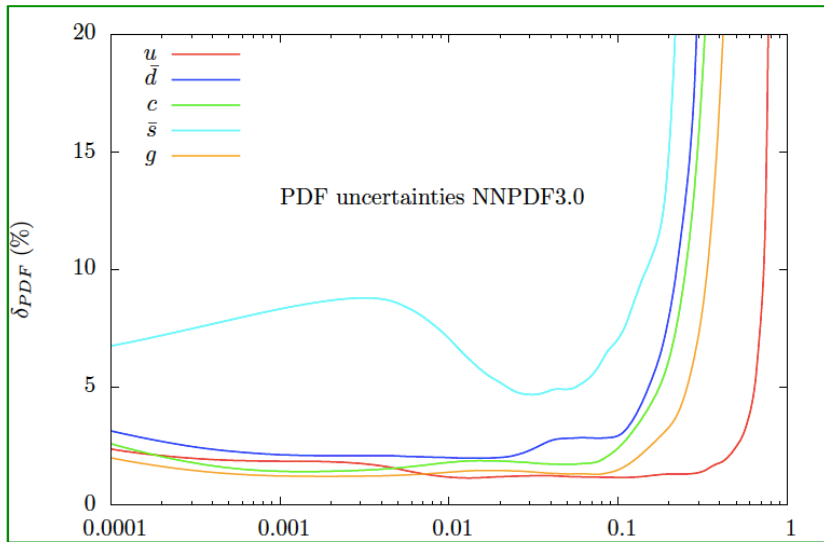


E($\bar{\nu}_\tau$) < 38 GeV

- At LO $F_4 = 0$, $2xF_5 = F_2$
- At NLO $F_4 \sim 1\%$ at 10 GeV

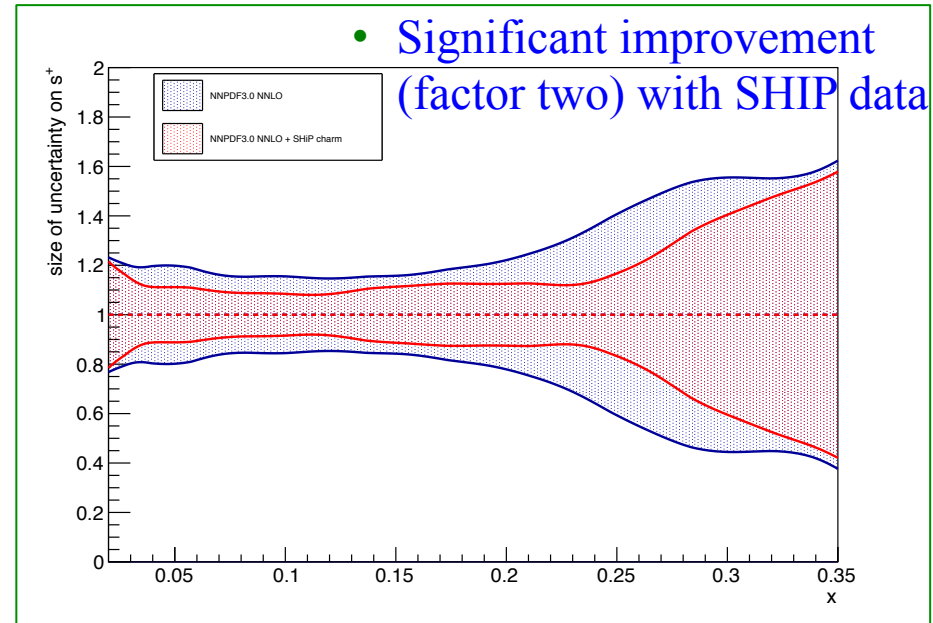
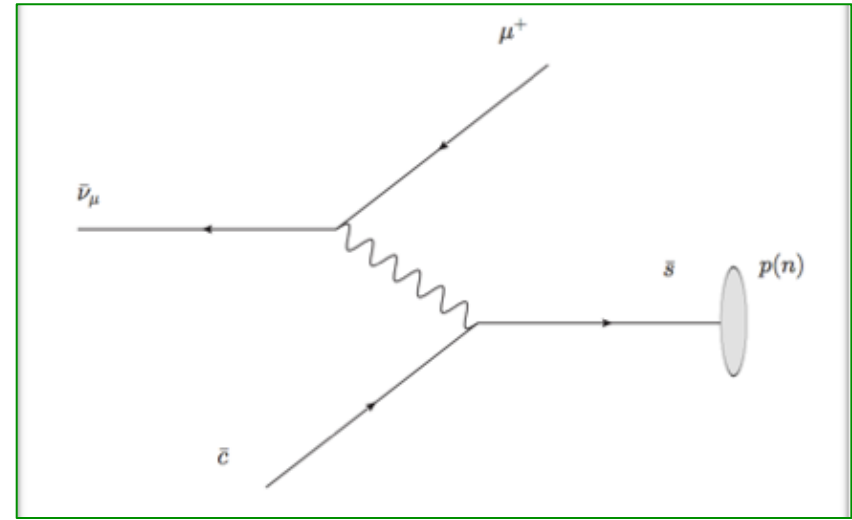
STRANGE QUARK NUCLEON CONTENT

- Charmed hadron production in anti-neutrino interactions selects anti-strange quark in the nucleon
- Strangeness important for precision SM tests and for BSM searches
- W boson production at 14 TeV: 80% via $u\bar{d}$ and 20% via $c\bar{s}$



Phys. Rev. D91 (2015) 113005

Fractional uncertainty of the individual parton densities $f(x; m_W^2)$ of NNPDF3.0



$$s^+ = s(x) + \bar{s}(x)$$

Added to NNPDF3.0 NNLO fit, Nucl. Phys. B849 (2011) 112–143, at $Q^2 = 2 \text{ GeV}^2$

DARK MATTER SEARCH

WITH THE NEUTRINO DETECTOR

χ produced by a dark photon decay

$$\chi e^- \rightarrow \chi e^-$$

P. deNiverville, D. McKeen, and A. Ritz,

Phys.Rev. D86 (2012) 035022

$\alpha' =$ dark photon coupling with χ

SIGNAL SELECTION

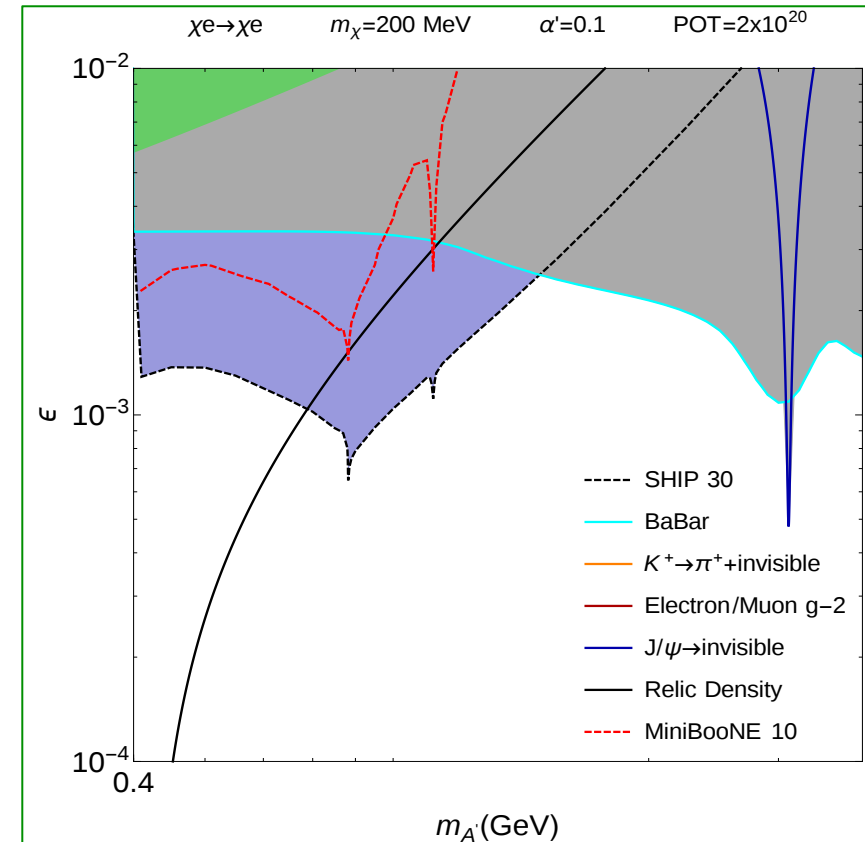
$$\left\{ \begin{array}{l} 0.01 < \theta < 0.02 \\ E < 20 \text{ GeV} \end{array} \right.$$

BACKGROUND PROCESSES

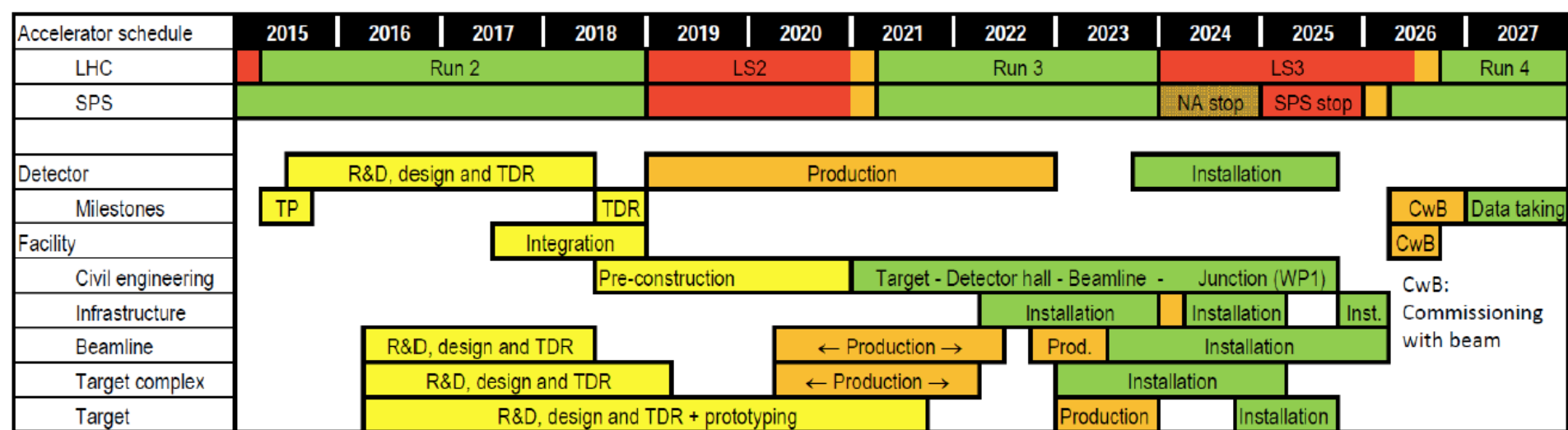
	ν_e	$\bar{\nu}_e$	ν_μ	$\bar{\nu}_\mu$	all
Elastic scattering on e^-	16	2	20	18	56
Quasi - elastic scattering	105	73			178
Resonant scattering	13	27			40
Deep inelastic scattering	3	7			10
Total	137	109	20	18	284

$\epsilon =$ dark photon coupling with e.m. current

$m_A =$ dark photon mass



Project schedule



~10 years from TP to data taking

- ✓ Schedule optimized to avoid interference with operation of North Area
 - ➔ Preparation of facility in four clear and separate work packages (target complex, detector hall, beam line and junction cavern)
- ✓ All TDRs by the end of 2018
- ✓ Four years for detector construction, plus two years for installation
- ✓ Updated schedule with new accelerator schedule (Run 2 up to end 2018, 2 years LS2) relaxes current schedule
 - ➔ Data taking 2026



Summary

SHIP to complement searches for New Physics at CERN in the largely unexplored domain of new, very weakly interacting particles with masses $O(10)$ GeV

- ✓ *Unique opportunity for ν_τ physics*
- ✓ *Sensitivity improves past experiments by $O(10000)$ for Hidden Sector and by $O(\sim 1000)$ for ν_τ physics*
- ✓ *The SHiP proposal submitted in April 2015 to the SPS Committee at CERN with recommendations delivered in January 2016*
- ✓ *SHiP could therefore constitute a key part of the CERN Fixed Target programme in the HL-LHC era. SPSC **recommends** that the SHiP proponents proceed with the preparation of a Comprehensive Design Report (CDR), and that this preparation be made in close contact with the planned Fixed Target working group.*
- ✓ *SHiP is an experiment recognised at CERN (grey book) since May 2016*
- ✓ *Optimisation of the design going on: many technological choices still waiting for your contribution!*